

# *The current state of the solar modelling problem*

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## The Sun as a benchmark star

### The role of the Sun:

Well-studied, **helioseismic constraints**, **neutrino fluxes**, testbed for **physical ingredients**. The Sun is used as a **reference**:

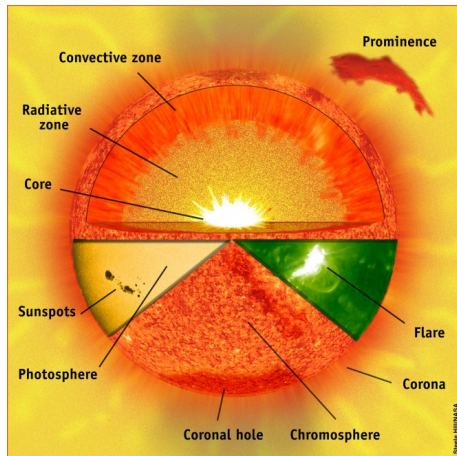
- **Metallicity scale**,
- **Enrichment laws**,
- **SSM framework**,
- **Paved the way for asteroseismology using solar-like oscillations.**

**Most of our models will include some ingredients that have been calibrated on the Sun. Thus, if you change the way you model the Sun, you impact stellar physics as a whole.**

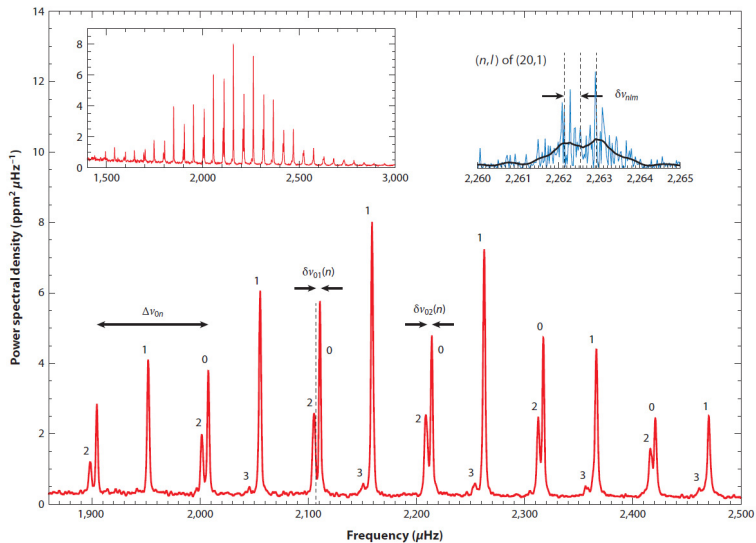
**But how well do we know the Sun?**

## Stellar Structure (Low-mass main-sequence star, like the Sun)

- **Convective layers:** macroscopic motions of the fluid,
- **Radiative layers:** transport of energy by photons,
- **Core:** Thermonuclear reactions.



# Observational Context - Solar-like oscillations



What is the problem with the Sun?

## The solar modelling problem

### *A brief history of Standard Solar Models*

Before 2004, high metallicity solar models ( $Z = 0.0182$ ):

- ① **Correct** position of the BCZ,
- ② **Correct** Helium abundance in the CZ,
- ③ Sound Speed profile relative differences of up to **0.006**.

(From Kosovichev & Fedorova 1991, 1993, Vorontsov et al. 1991)

**But:** slow degradation as physical ingredients were updated.

From 2004, downward revision of the solar  $Z$ :

- ① **Wrong** position of the BCZ,
- ② **Wrong** Helium abundance in the CZ,
- ③ Sound Speed profile relative differences of up to **0.02**.

## The controversy in "brief"

### Spectroscopy

- ① Asplund et al. 2005: **Low Z.**
- ② Asplund et al. 2009: **Low Z.**
- ③ Caffau et al. 2011: **Higher Z.**
- ④ Scott et al. 2015a,b: **Low Z.**
- ⑤ Amarsi et al. 2017, 2019, 2020, 2021: **Low Z.**
- ⑥ Young 2018: **Higher Ne.**
- ⑦ Asplund et al. 2021: **Low Z, Ne Confirmed.**
- ⑧ Magg et al. 2022: **High Z.**

*Some "disagreements" remain...  
(Nabar 2022)*

### Helioseismology & Neutrinos

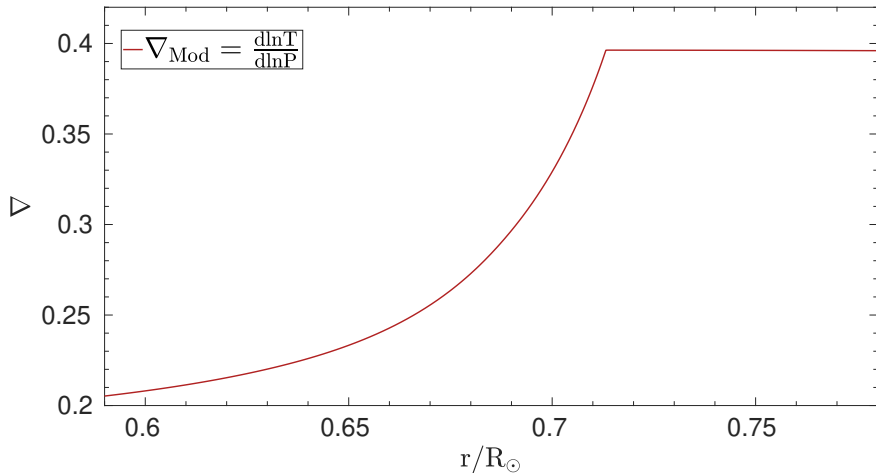
- ① Sound speed is wrong!
- ② Helium is wrong!
- ③ Neutrino fluxes are wrong!
- ④ BCZ is wrong!

⇒ Abundances **HAVE TO** change!  
They **ARE** higher! Helioseismology says so!

Or are they? What does helioseismology actually say?

## The problem of the BCZ

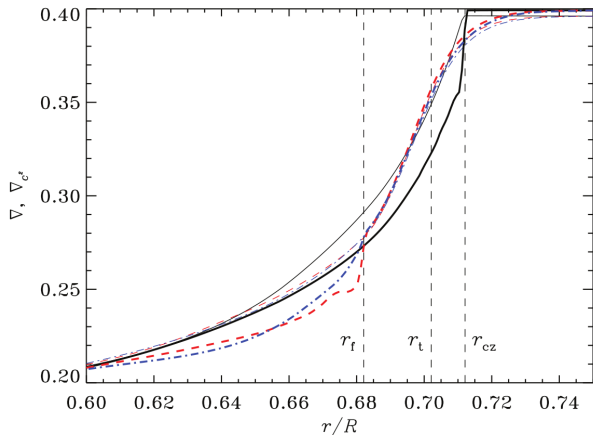
$r_{CZ} = 0.713 \pm 0.001 R_{\odot}$  (Basu & Antia 1997), SSM High Z:  $r_{CZ} \approx 0.713$ , SSM Low Z:  $r_{CZ} \approx 0.720 \Rightarrow$  Low Z is wrong!





## The problem of the BCZ

### What actually influences the BCZ position?



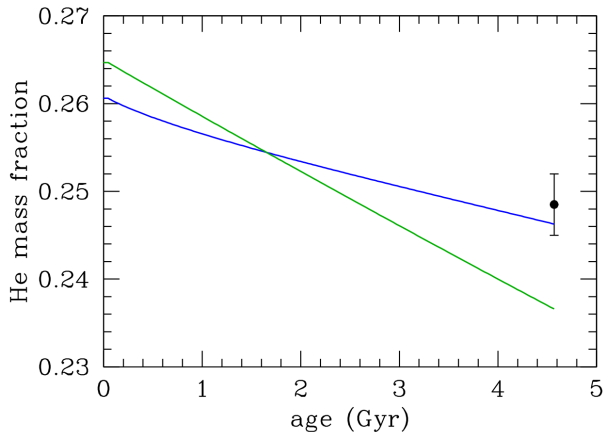
BCZ position is affected by:

- Overshooting;
- Opacities at the BCZ;
- Nuclear reaction rates;
- Transport of chemicals;
- ...;

Christensen-Dalsgaard et al. 2011; see also Zhang et al. 2019 and Monteiro et al. 1994 for older references!

## The problem of Helium

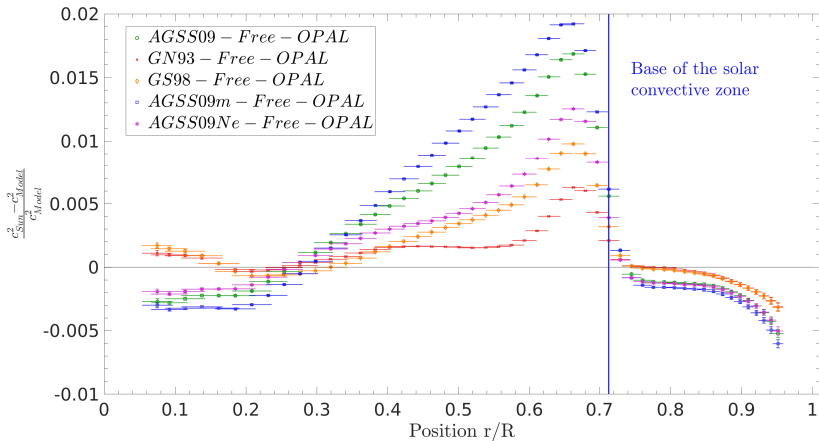
$Y_{CZ} = 0.2485 \pm 0.0035$  (Basu et al. 1995), SSM High Z:  $Y_{CZ} \approx 0.245$ ,  
SSM Low Z:  $Y_{CZ} \approx 0.235 \Rightarrow$  Low Z is wrong!



Helium is affected by:

- Abundances ( $Z_{\odot}$  value);
- Opacities at high T;
- Nuclear reaction rates;
- Transport of chemicals;

# The problem of Sound Speed



**Sound speed is NOT a direct measurement of abundances.**

## What is a Standard Solar Model?

*Definition of SSMs (Bahcall et al. 1982)*

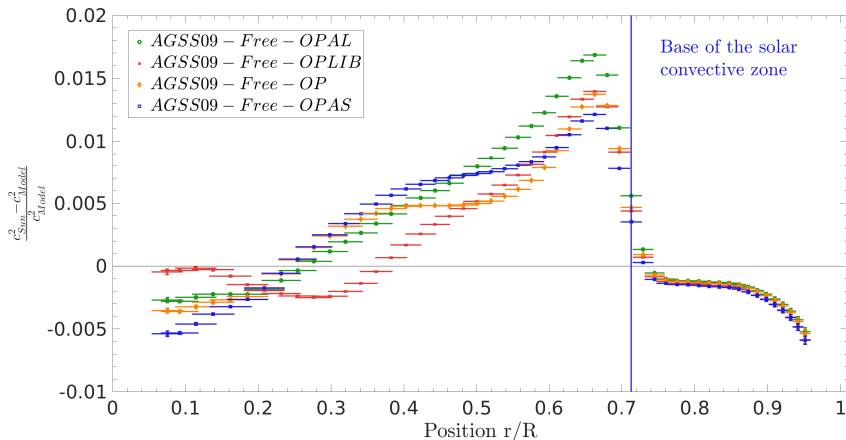
Recipe to compute solar models (Magg et al. 2022):

- ① Standard opacities (OP, Badnell et al. 2005),
- ② Standard EOS (FreeEOS, Irwin 2012),
- ③ Mixing length theory of convection (Böhm-Vitense 1958).
- ④ Effects of microscopic diffusion (Thoul et al. 1994).
- ⑤ Standard nuclear rates (Adelberger et al. 2011).

**Other ingredients:** OPLIB (Colgan et al. 2016), OPAS (Mondet et al. 2015), Saha-S (Baturin et al. 2013,2017,2019), ...

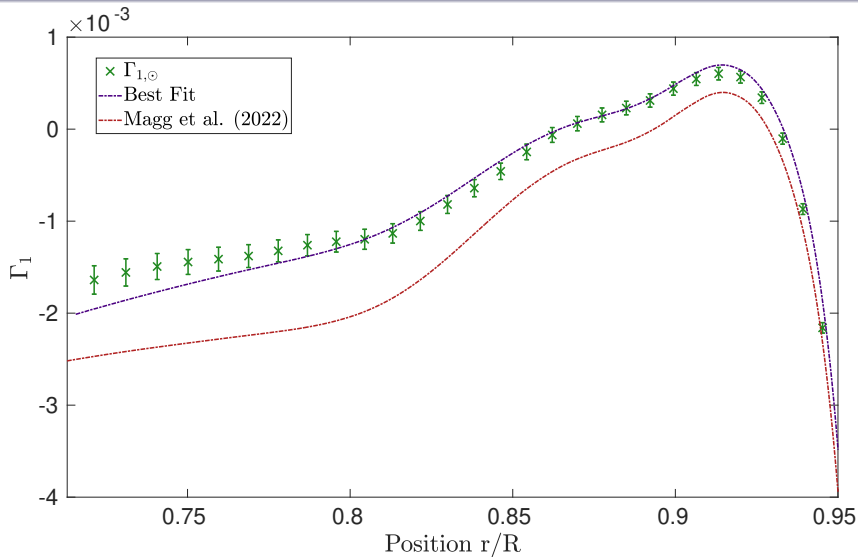
**Simplified physics:** Rotation, overshooting, radiative accelerations, convection ...

## What does helioseismology actually say?

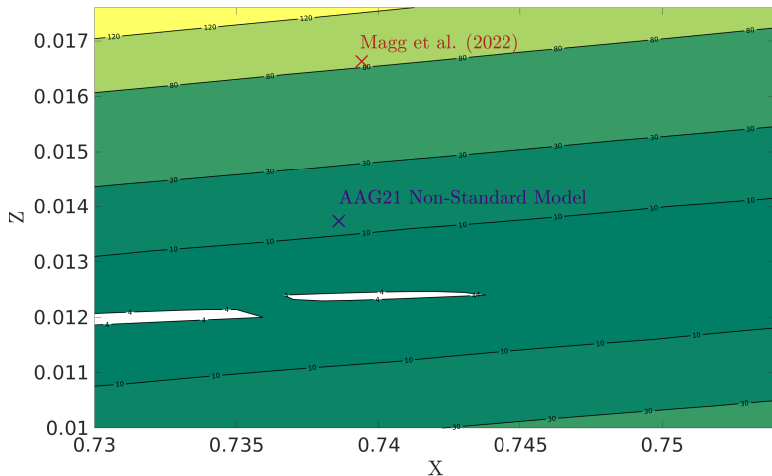


Opacity modifications have the same effect as abundance modifications. Sound speed provides degenerate information.

## Helioseismology and abundances

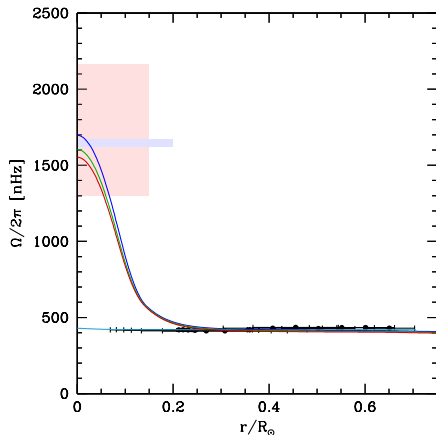
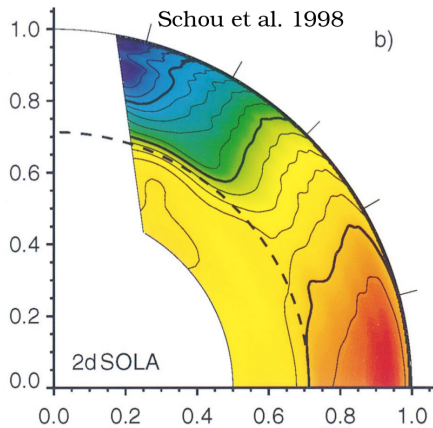


## Helioseismology and abundances (Buldgen et al. (in prep))



**Studied since early 2000s:** Takata & Shibahashi 2001; Lin & Däppen 2005; Lin et al. 2005; Antia & Basu 2005; Vorontsov et al. 2013, 2014; Buldgen et al. 2017; Baturin et al. 2022, Buldgen et al. (in prep).

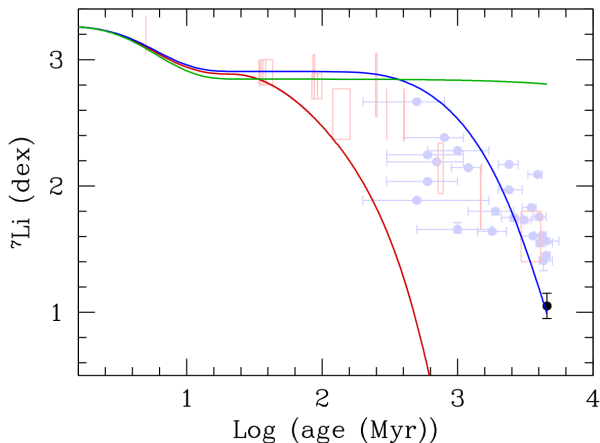
## What does helioseismology actually say?



The internal rotation of Sun is known since 1988 (Kosovichev 1988), explained for a while (e.g. Eggenberger et al. 2005).

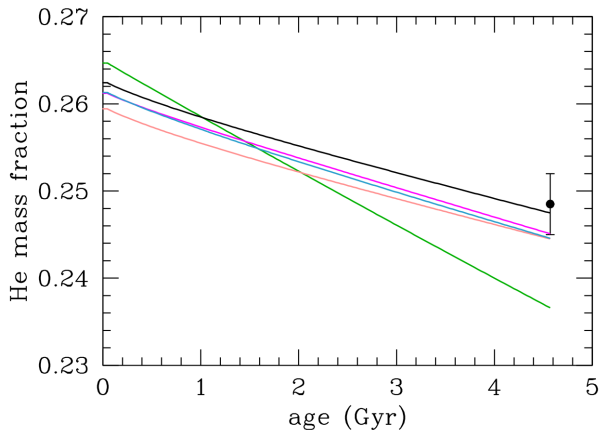


## Depletion of light elements (Eggenberger et al. 2022)



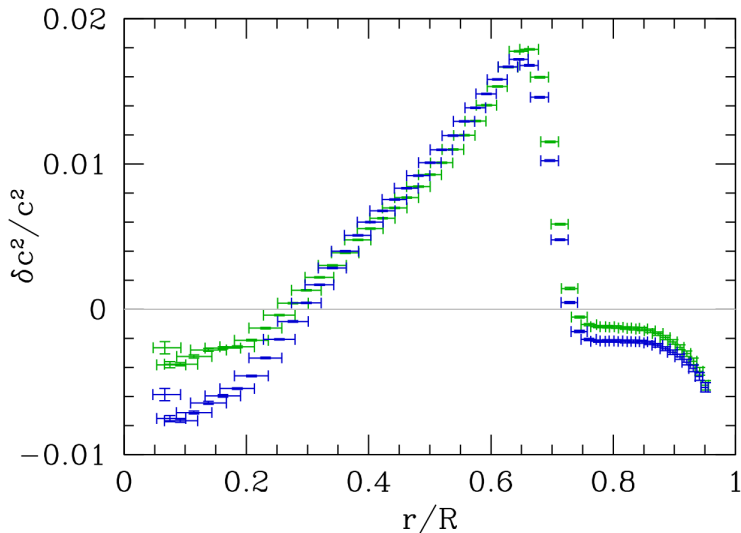
Lithium depletion is an issue since 1990s (Proffitt & Michaud 1991, Richard et al. 1996).

## Non-standard models and helium



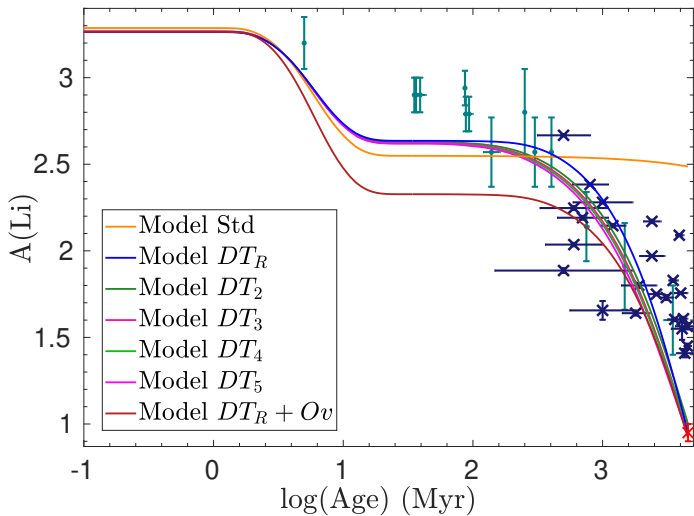
The helium-lithium correlation exists for multiple shapes of the transport coefficients!

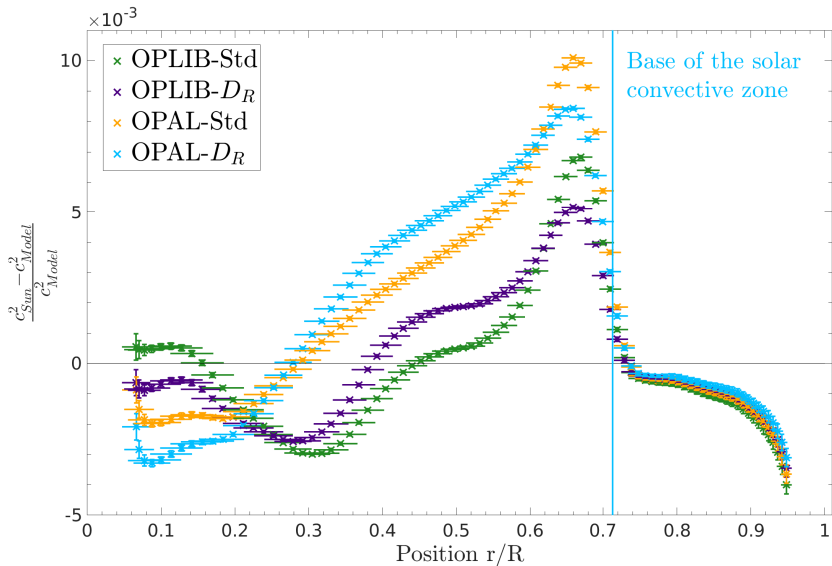
## Non-standard models and the BCZ



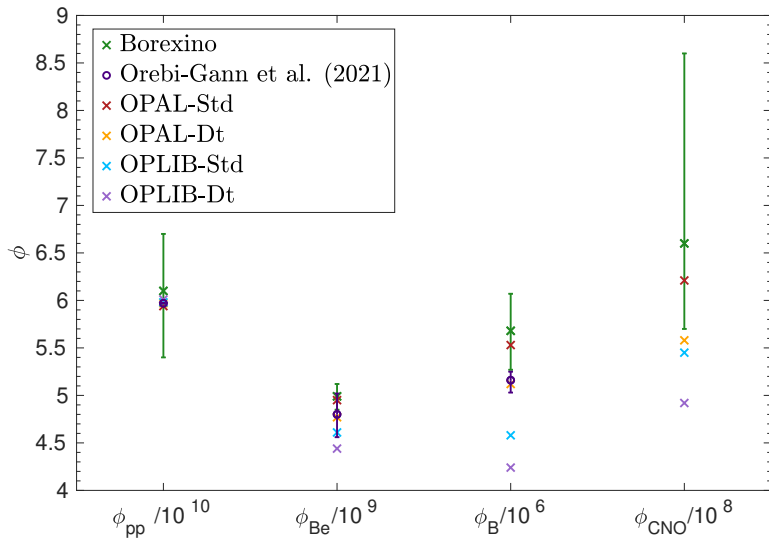
Sound speed at the BCZ not strongly affected by mixing.

Is the problem solved with higher abundances?

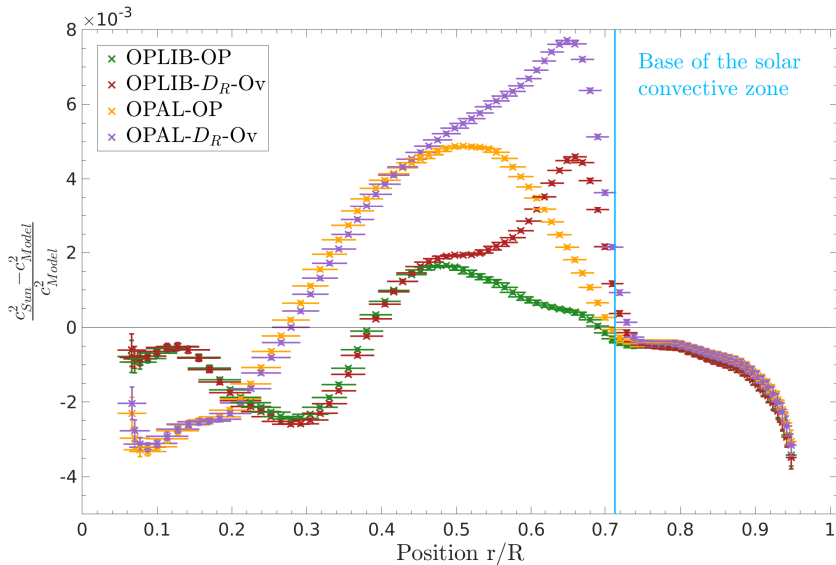




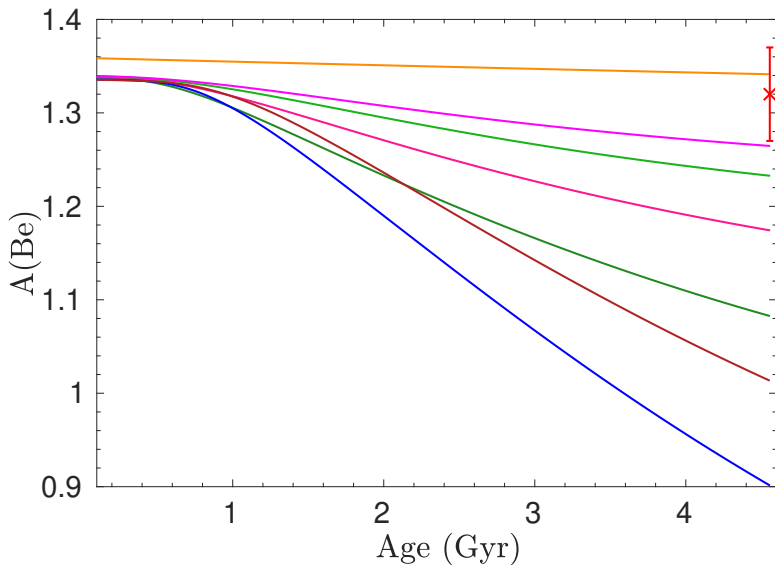
# Magg et al. 2022 Models - Neutrino Fluxes



# Magg et al. 2022 Models - Sound Speed

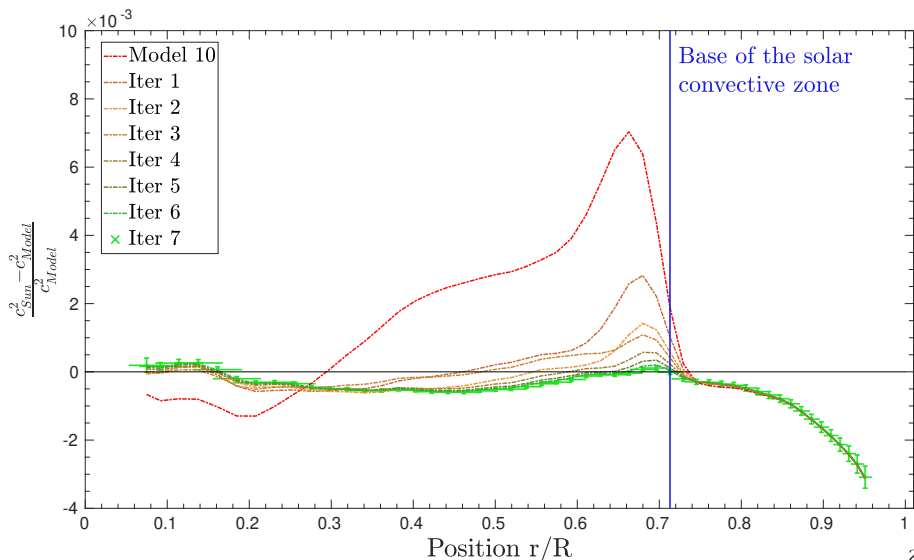






Can you push models further?

# Seismic models: providing a full structure (Buldgen et al. 2020)



### Wishes for coming years:

#### From modelling side:

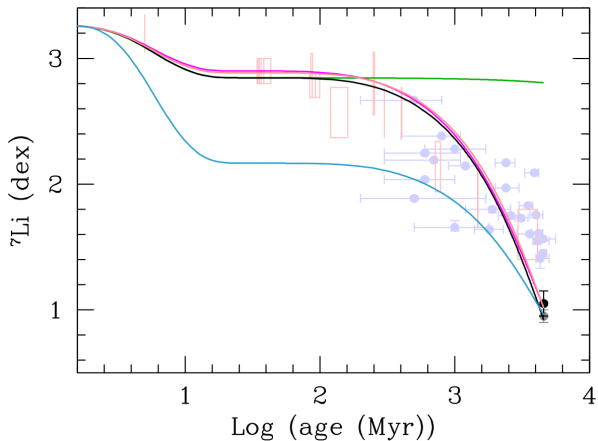
- Beryllium determination and transport calibration.
- Differentiation of overshooting and  $\kappa$  modification for entropy of the CZ.
- Determination of the metallicity in the envelope.
- Analysis of numerical robustness of solar models.

#### From physics side:

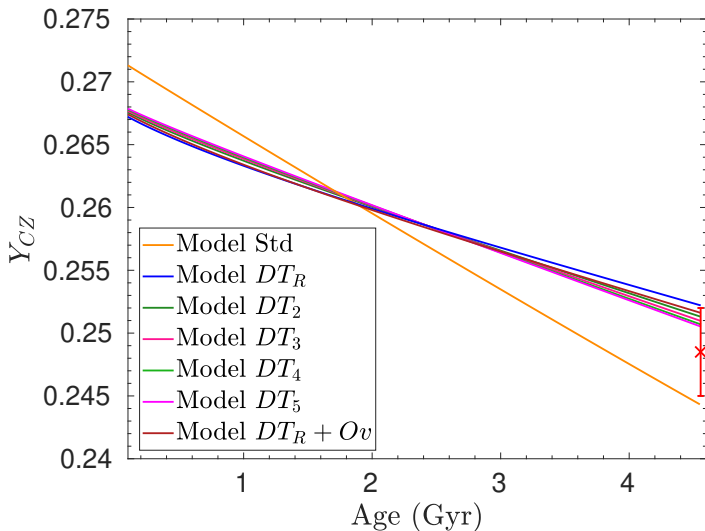
- New opacities (SCO-RCG tables, OP), comparisons on more TD paths.
- New EOS (MHD2020).

Thank you for your attention!

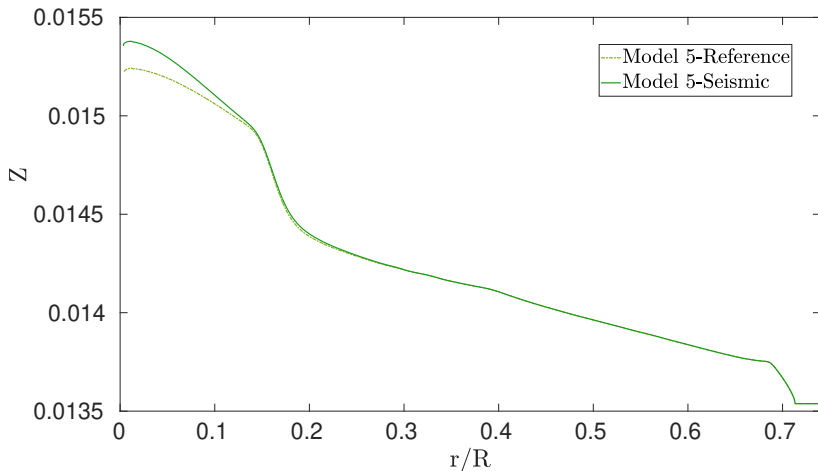
## Non-standard models and overshooting



You also get constraints **on overshooting at the BCZ!**



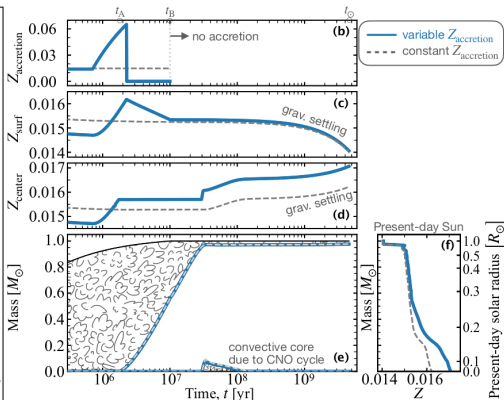
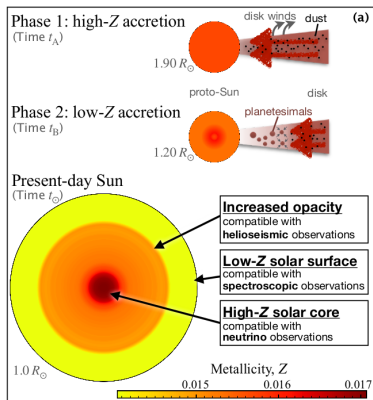
## Seismic models: providing a full structure



**Solution:** from  $P$  and  $\rho$  for a given  $X$  and  $Z$ , determine from the EOS  $T(\rho, P, X, Z)$ . Then compute  $\varepsilon(T, \rho, X, Z)$  so that  $L = L_{\odot}$ .



# Non-standard models and planetary formation



Solar core influenced by planetary formation mechanisms!