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TALYS: A unified approach towards nuclear reaction calculations

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Swedish Physical Society, Nuclear Physics, Uppsala, Sweden, November 12, 2014



Outline

- Introduction to the TALYS code
- Today's examples:
 - Fission cross sections and yields
 - Residual production cross sections
 - Astrophysics: Solar abundances
- Parameter uncertainties and links to applications
- Summary



- TALYS
- Nuclear model code by NRG Petten, CEA Bruyeres-le-Chatel, UL Bruxelles
- g,n,p,d,t,h,a induced reactions from 10⁻⁵ eV up to 250 MeV

Release:

- www.talys.eu
- Latest official version, TALYS-1.6, released december 21, 2013.
- Estimated 500-1000 users, > 800 publications

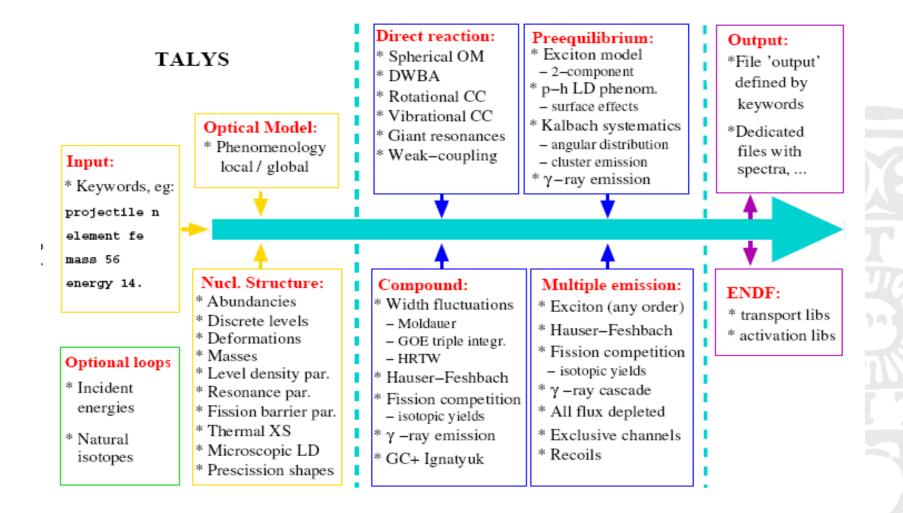
Software issues:

- TALYS is ready to be used by persons other than the authors
- Very flexible:
 - 4-line idiot-proof input (element, mass, projectile, energy), but also...
 - > 250 keywords to change models, parameters, level of output, etc.
- 300 page manual, 20 widely varying sample cases
- Readable modular programming (extensions by others is relatively easy)
- Very robust, thanks to dripline-to-dripline and random input testing





TALYS code scheme





UPPSALA TALYS physics

General use:

• TALYS can be used for

In-depth single nuclide/reaction analyses

Global multi-nuclide calculations

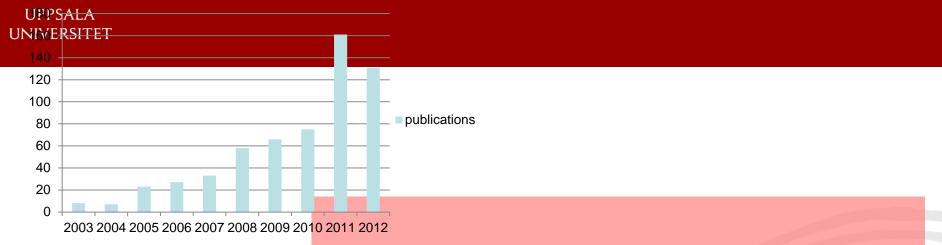
Complete output:

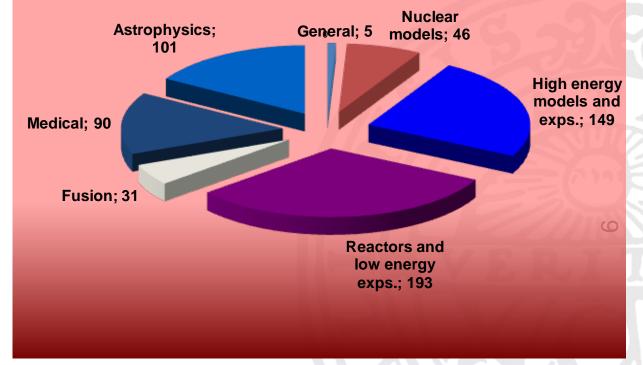
 Total, partial and residual production cross sections, (Double)differential spectra, Angular distributions per discrete level, Fission yields, Recoils, Isomeric production, Astrophysical reaction rates Gamma production, etc,

Recent accomplishment: option to use all optical, level density, fission and pre-equilibrium models phenomenological (Woods-Saxon, Fermi gas, Hill-Wheeler, exciton) or microscopical (Hartree-Fock-Bogolyubov-based, by Hilaire, Goriely, Bauge)



TALYS publications

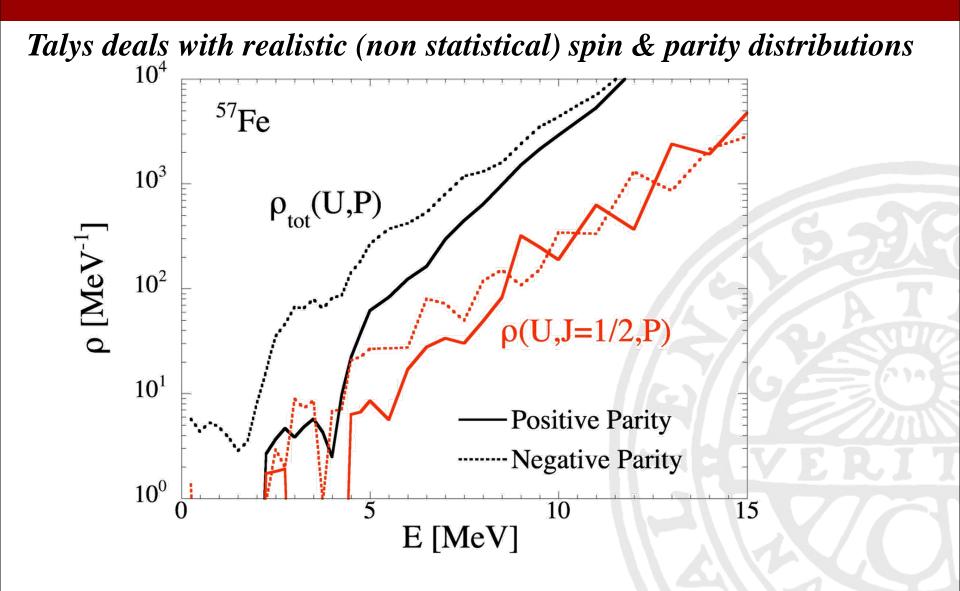






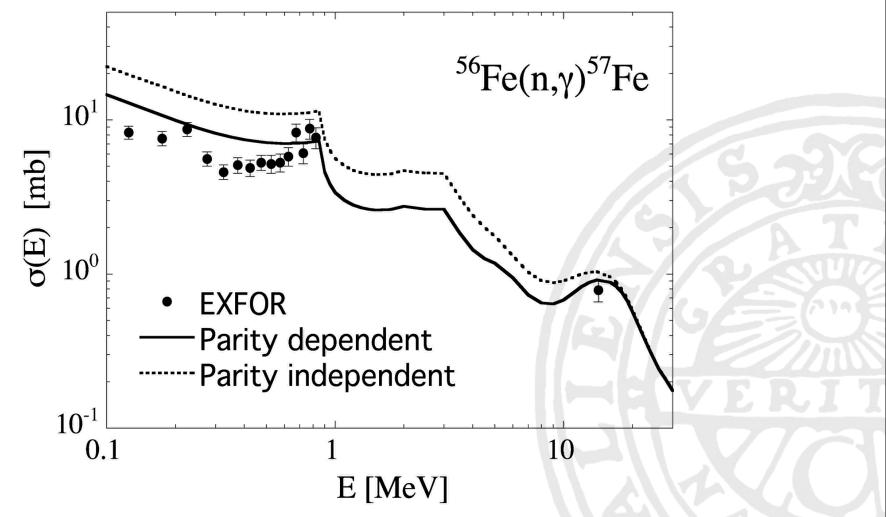
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Non statistical effects in nuclear level densities (HFB + D1S Gogny)



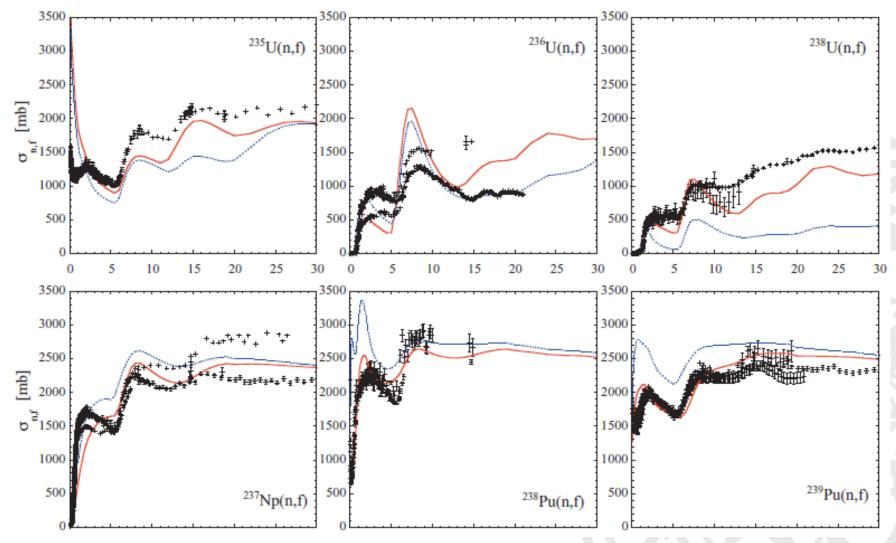


Talys deals with realistic (non statistical) spin & parity distributions





UPPSALA TALYS: Microscopic fission



S. Goriely, S. Hilaire, A.J. Koning, M. Sin and R. Capote, Phys Rev C79, 024612 (2009)



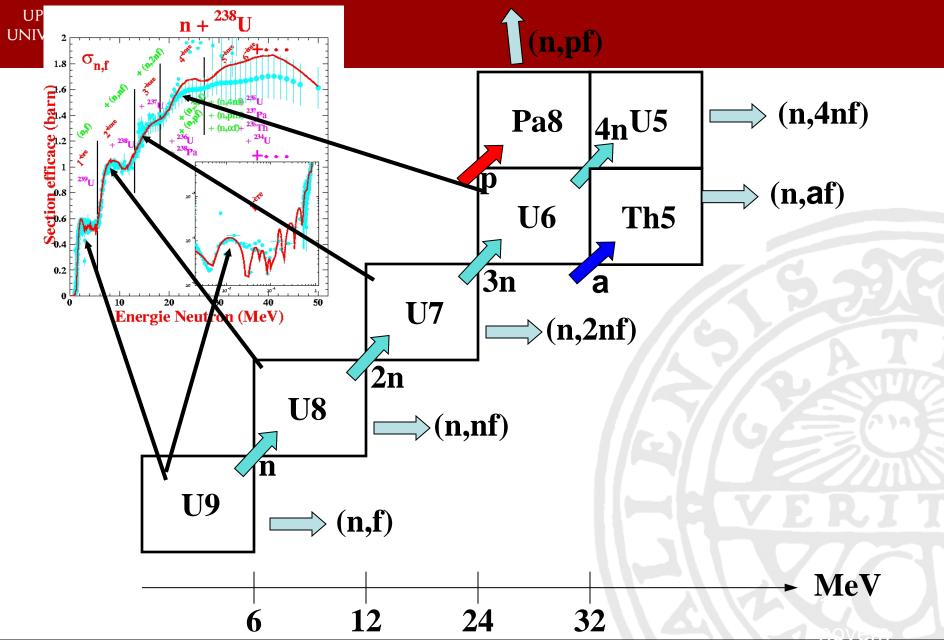
Coherent fission cross sections with phenomenological approach

Neutron induced fission on ²³⁸U

- several hundreds of parameters
- unique set for all fission chances or U targets

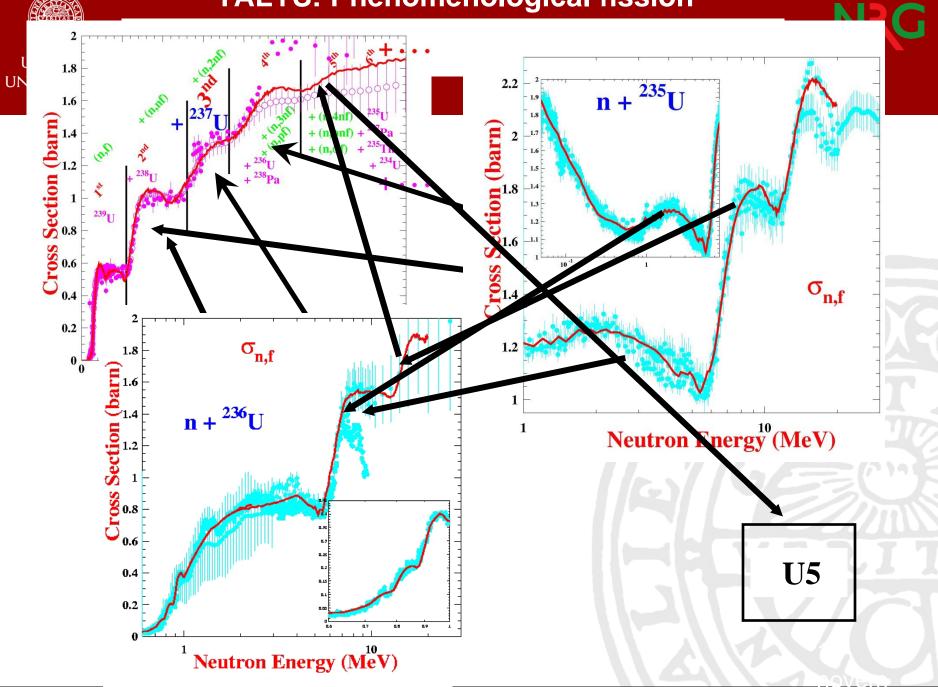


TALYS: phenomenological fission



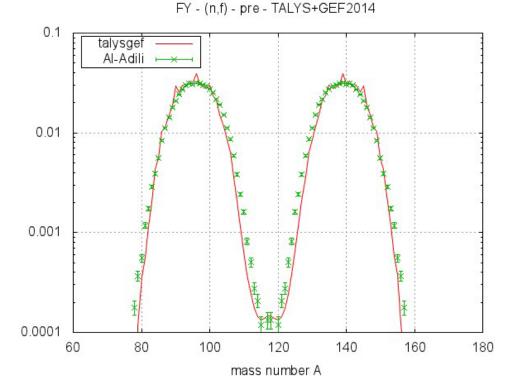
TALYS







- TALYS is now coupled with GEF, a fission code by Karl-Heinz Schmidt (GSI), which calculates fission observables on a microscopic basis, starting from an excitation energy, spin and parity
- TALYS feeds GEF with that starting point



²³⁴U(n,f)

Experimental data from PhD thesis of Ali Al-Adili, Uppsala University 2013.

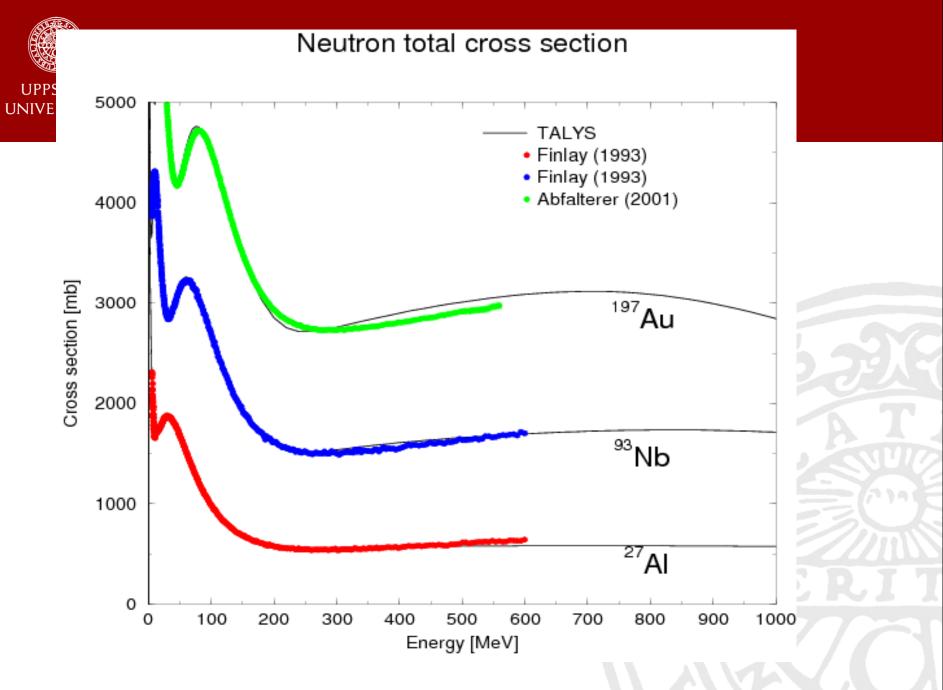
More examples: S. Pomp Presentation at FIESTA 2014, Santa Fe 13

yield

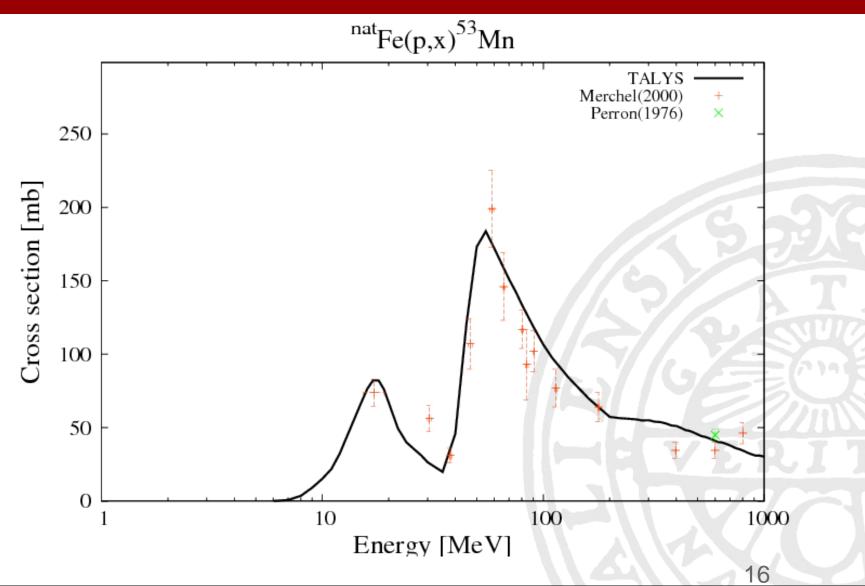


UNIVERSITET (computational) extension to 1 GeV

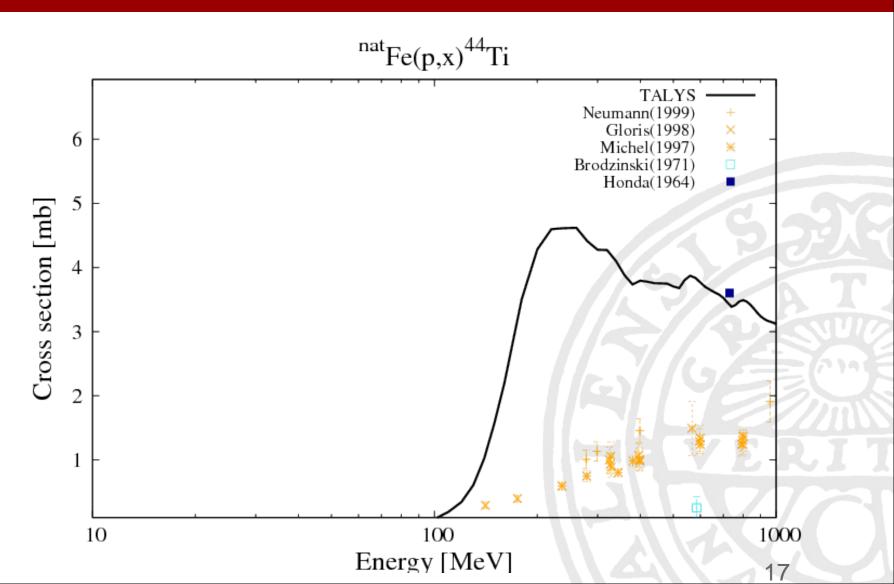
- Extension of KD03 OMP to higher energies: smooth extension of Vv(E) to negative values, double increase of Wv(E) above 200 MeV (ND2013)
- Exciton model for pre-equilibrium reactions: multiple pre-equilibrium emission up to any order already in TALYS, exciton model works at higher energies (see e.g. Mashnik-CEM, Chadwick-Hybrid MC).
- New mechanisms Pion prediction and fragmentation: Not (yet) taken into account, at which energy do we pay the price?
- Multiple Hauser-Feshbach decay: Equidistant excitation energy grid no longer appropriate. Logarithmic grid introduced for all residual nuclides in the reaction chain.
- TALYS-1.6
 - Total and residual production cross sections tested to some extent
 - Spectra and angular distributions not tested
 - Which is all acceptable.....as long as the code doesn't crash!
 - Gives user more flexibility to switch from a high-energy (INC) code to TALYS





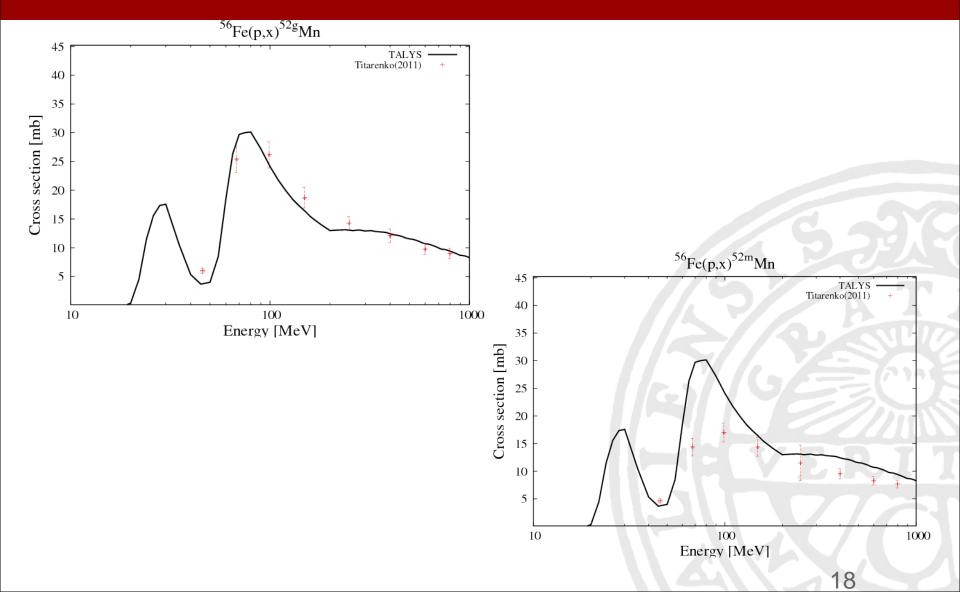




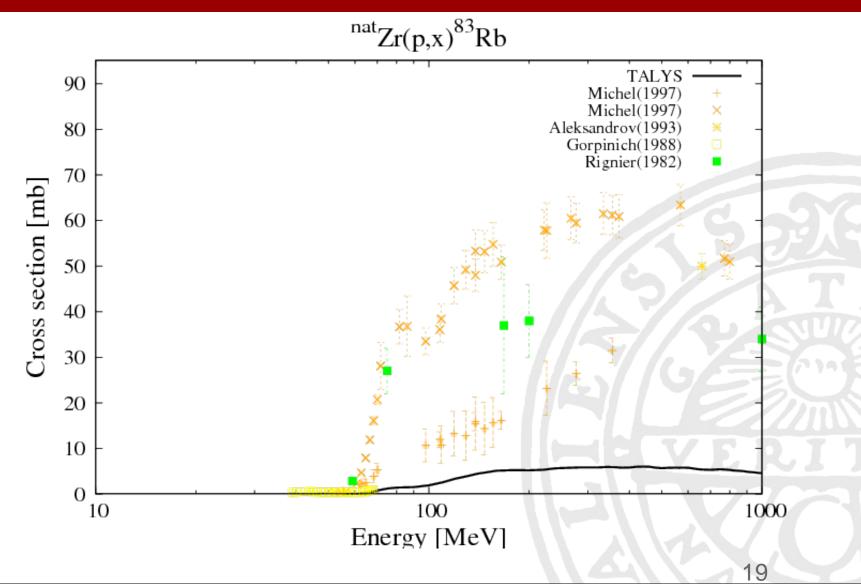




UPPSALA Not bad at all!



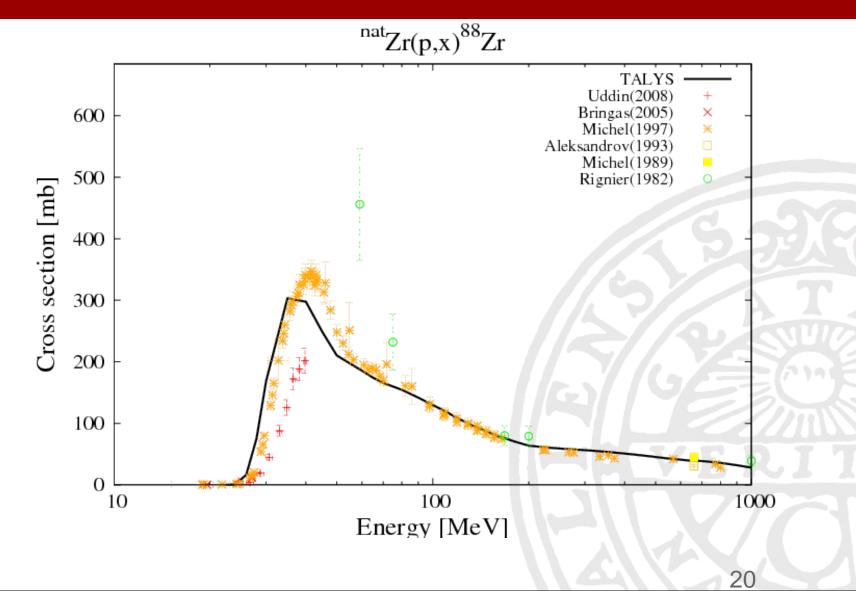






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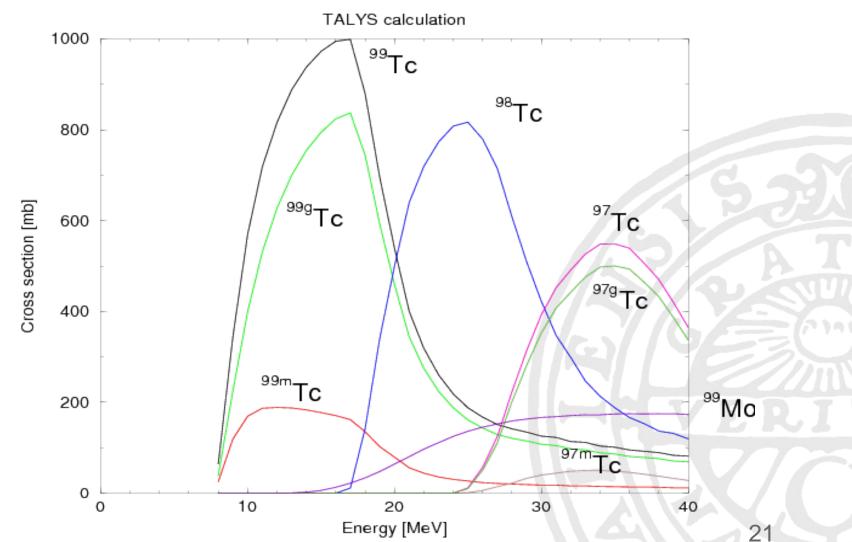
And finally: rather good!





Medical isotope production

p + Mo-100

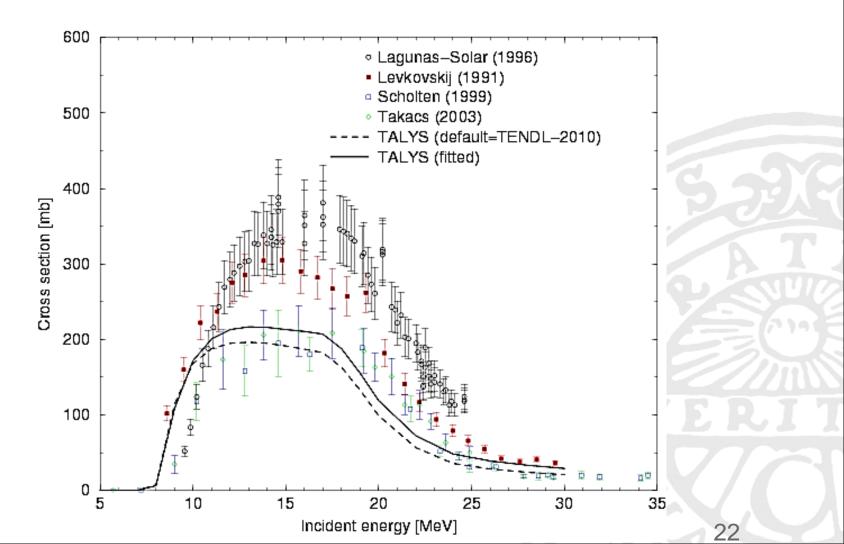






Medical isotope production

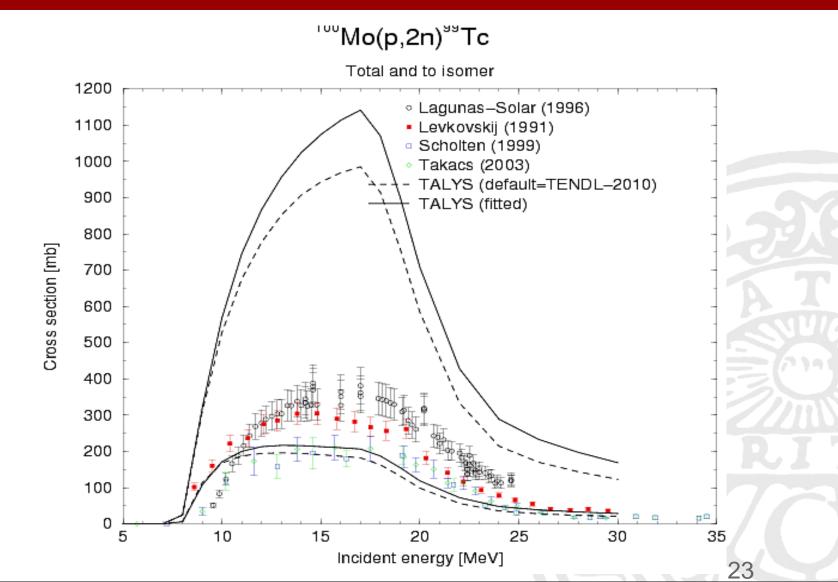
[™]Mo(p,2n)^{99m}Tc





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Isomeric ratio is essential: about 0.22!





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GLOBAL SYSTEMATIC APPROACHES Astrophysical r-process (1/4)

Solar abundancies for 140 < A < 200 : r-nuclei ?

Two options :

- Supernovae :

Matter ejection without any problem Great sensitivity to thermodynamical conditions No clearly identified astrophysical site Failure of explosion models

- Neutron stars collisions :

Enough neutrons Binary systems abundancies ? Matter ejection ? Large sensitivity to fission models

Recently solved

+++ +++

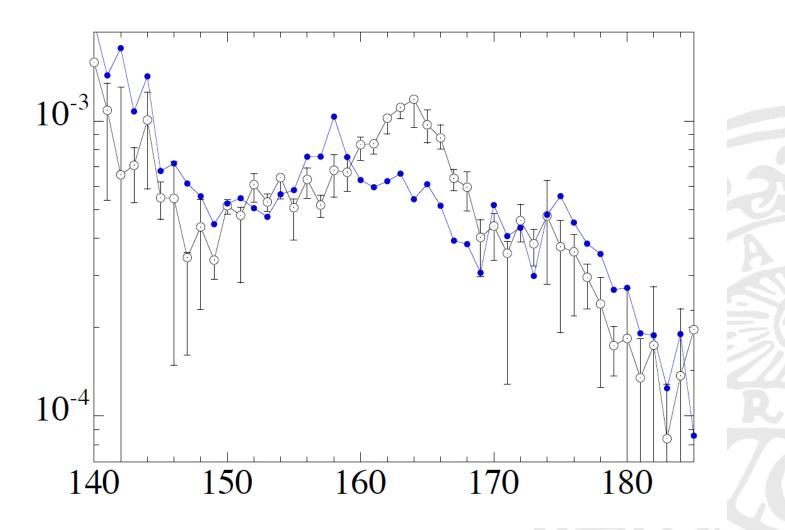
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GLOBAL SYSTEMATIC APPROACHES Astrophysical r-process (2/4)

Solar abundancies for 140 < A < 200 : situation before 2013 ?

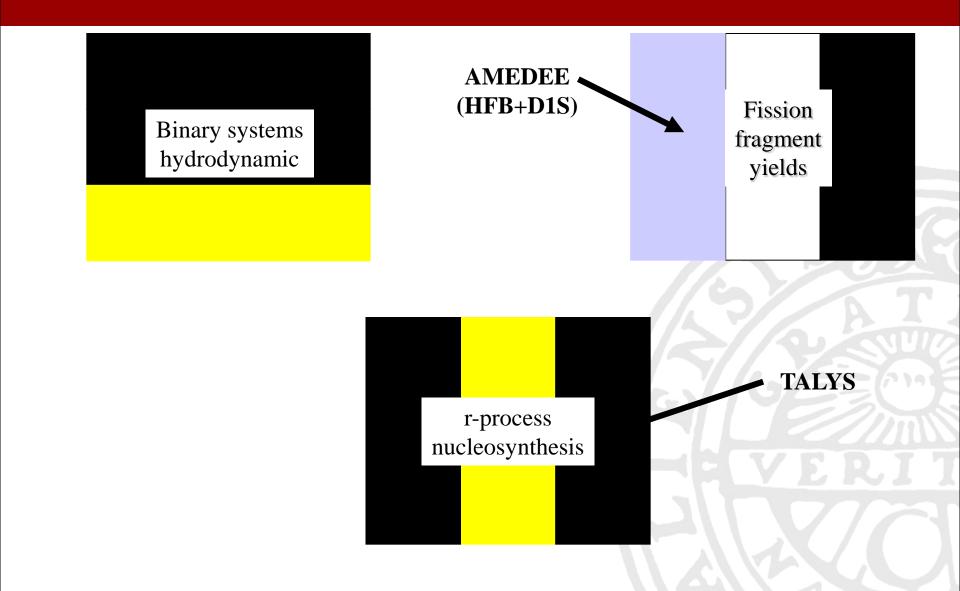




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GLOBAL SYSTEMATIC APPROACHES Astrophysical r-process (3/4)





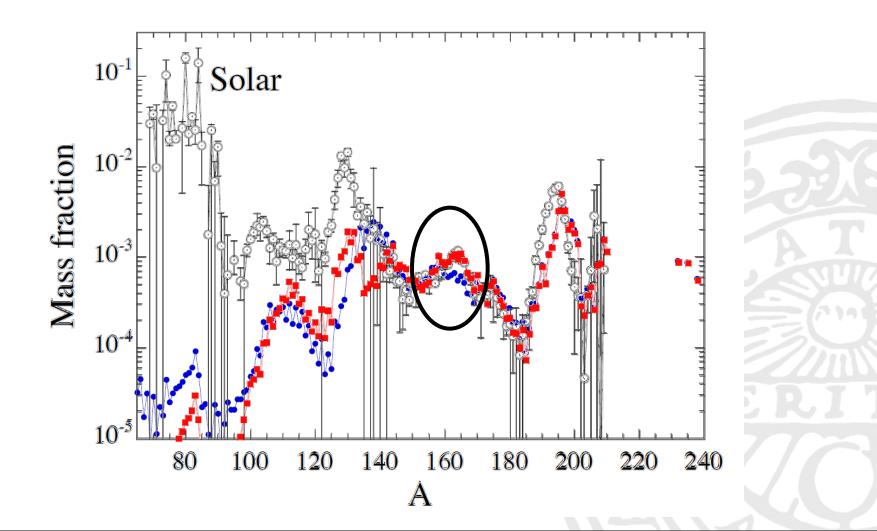


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GLOBAL SYSTEMATIC APPROACHES Astrophysical r-process (4/4)



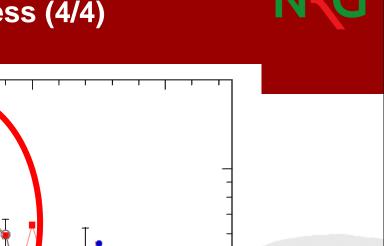
Solar abundancies for $140 \le A \le 200$: situation in 2013 2013

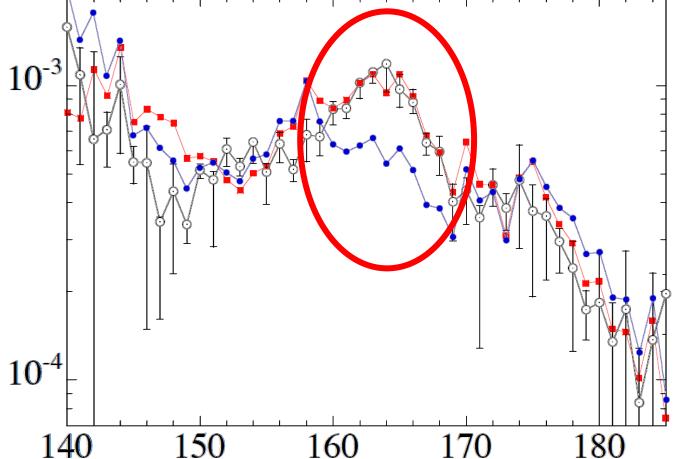




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GLOBAL SYSTEMATIC APPROACHES Astrophysical r-process (4/4)





More details for discussions/explanations in Goriely et al, PRL 111, 242502 (2013)



Covariances from nuclear model calculations

Experimentalists:

- · are well educated: give uncertainties with their results.
- Theoreticians:
 - Bad behaviour: most of them say their models are good (or bad), but none of them says how good (or bad): x-y instead of x-y-dy

No excuse possible:

- the strong nucleon-nucleon force is not known,
- the exact many-body problem is not solved,
- \longrightarrow all nuclear models are limited
- so where are the uncertainties of all the cross sections, spectra, angular distributions, etc.?

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Approach

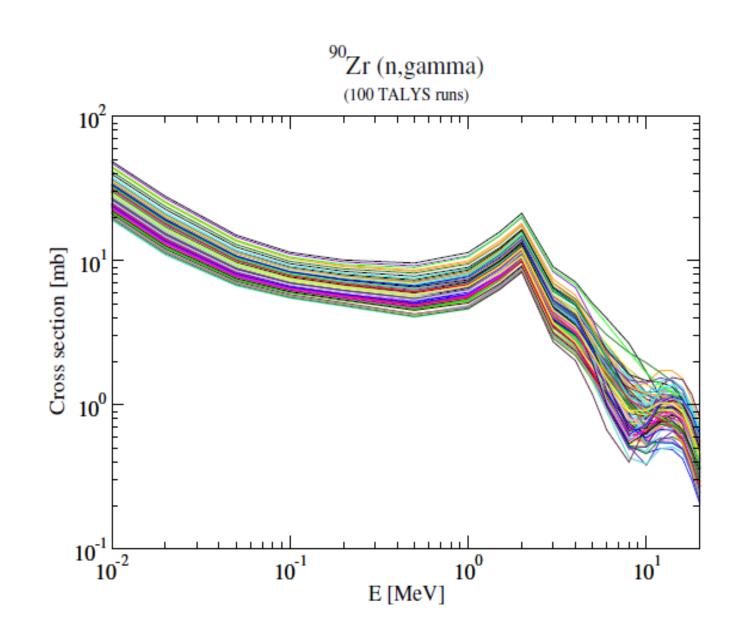


- Find a nuclear model code that predicts all open reaction channels, and is very flexible in input and output.
- Assess realistic uncertainties for the input, i.e. nuclear model, parameters.
- Propagate these uncertainties directly to the cross sections, angular distributions, gamma production, energy spectra, etc. using a Monte Carlo method.
- Obtain full covariance matrix (diagonal elements → uncertainties).

Covariances

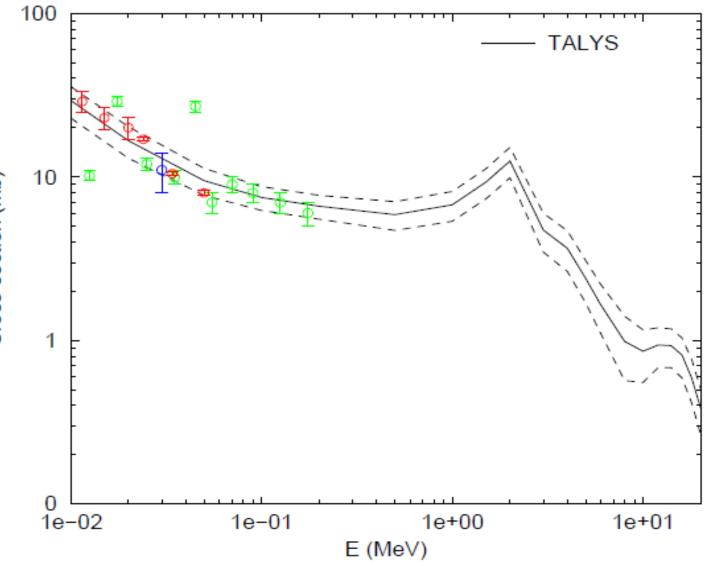


- Nuclear model parameter vector p: e.g. $p^1 = a_{ld}(26, 56)$, $p^2 = a_{ld}(26, 57)$, $p^3 = r_V$, etc.
- Physical quantity vector σ of length N: e.g. $\sigma^1 = \sigma_{n\gamma}(E_1), \ldots, \sigma^i = d\sigma_{el}/d\Omega(E_1, \Theta_1), \ldots, \sigma^N$
- $\sigma = T(\mathbf{p})$, where the function T stands for TALYS.
- Let \mathbf{p}_0, σ_0 be the best parameter/quantity set.
- Perform k = 1, K (=1000) TALYS calculations with p drawn at random from a Gaussian distribution.
- Covariance matrix $V_{ij} = \frac{1}{K} \sum_{k=1,K} (\sigma_k^i \sigma_0^i) (\sigma_k^j \sigma_0^j)$ for i,j=1,N
- Relative covariance matrix: $R_{ij} = V_{ij}/(\sigma_0^i \sigma_0^j)$ for i,j=1,N.



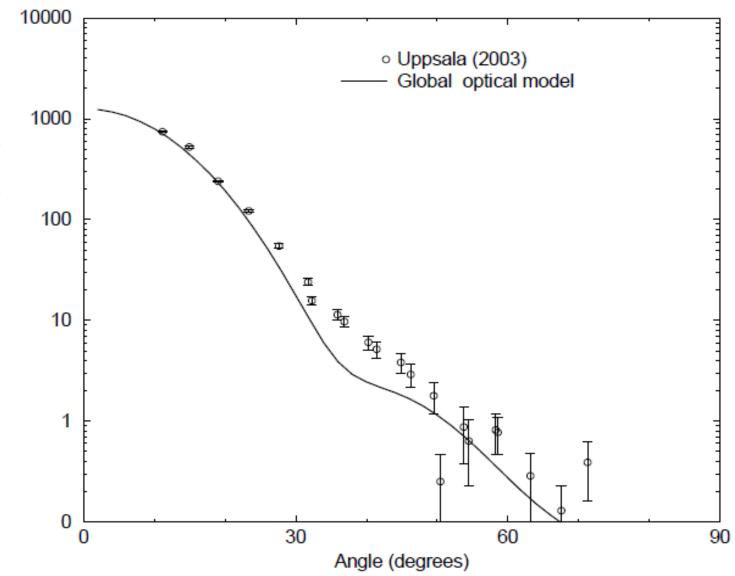
5-56

 $^{90}Zr(n,\gamma)$



33

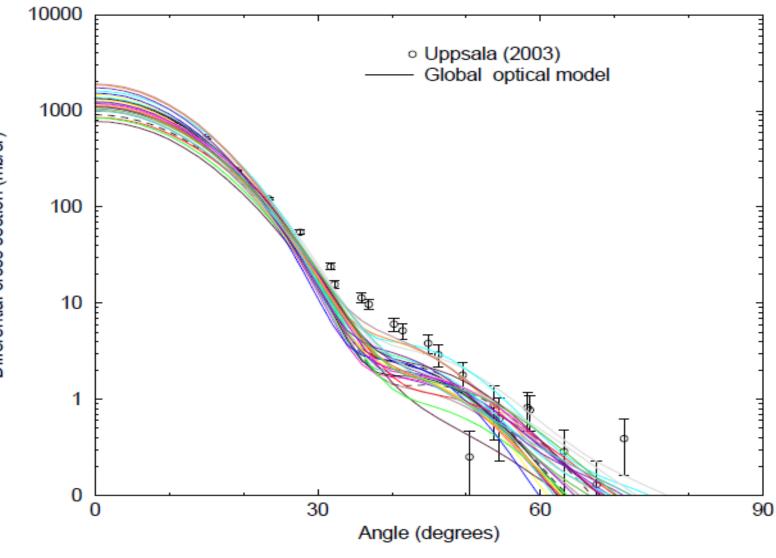
¹²C(n,el) at 96 MeV



Differential cross section (mb/sr)

15385.

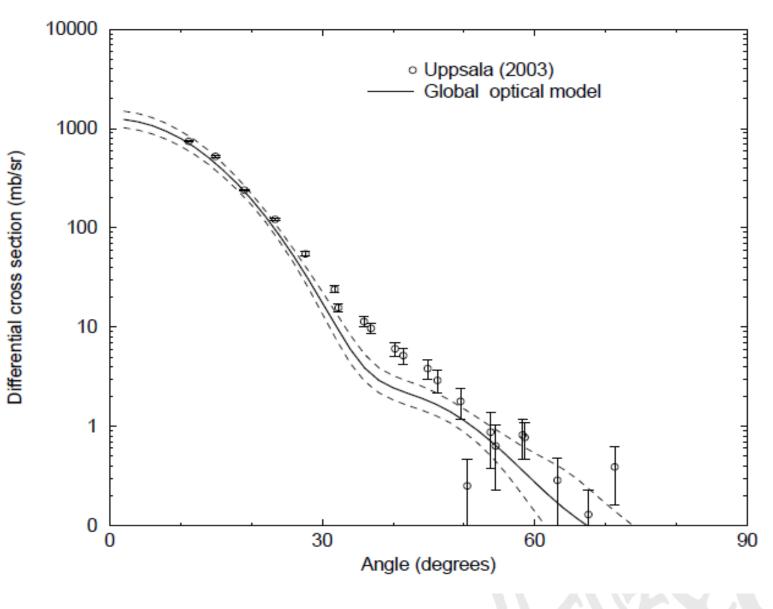
¹²C(n,el) at 96 MeV



35

5355.

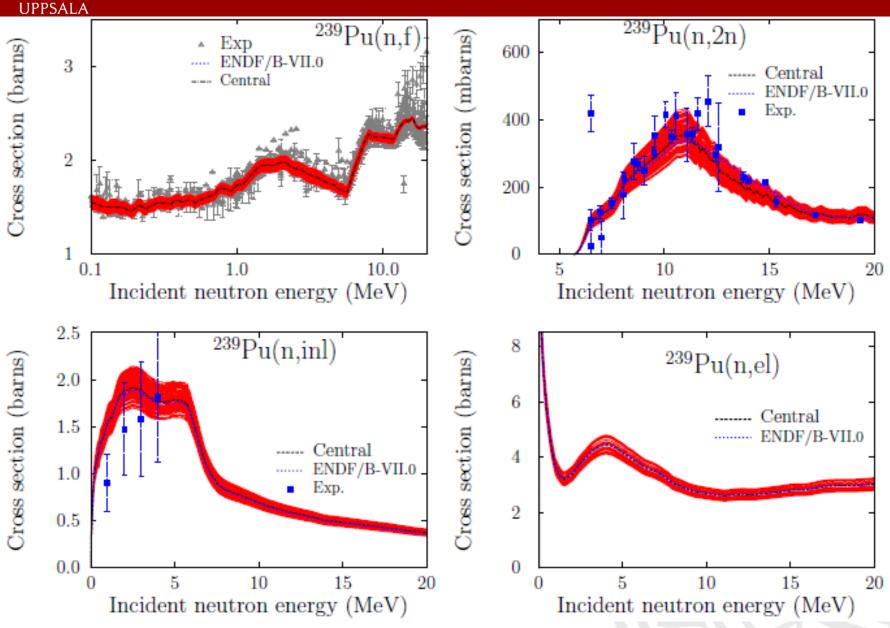
¹²C(n,el) at 96 MeV



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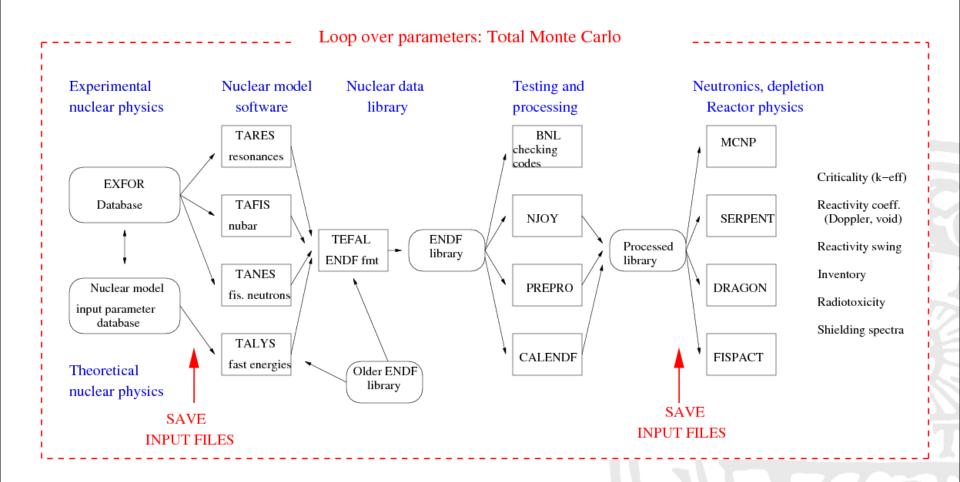
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Random cross sections for ²³⁹Pu

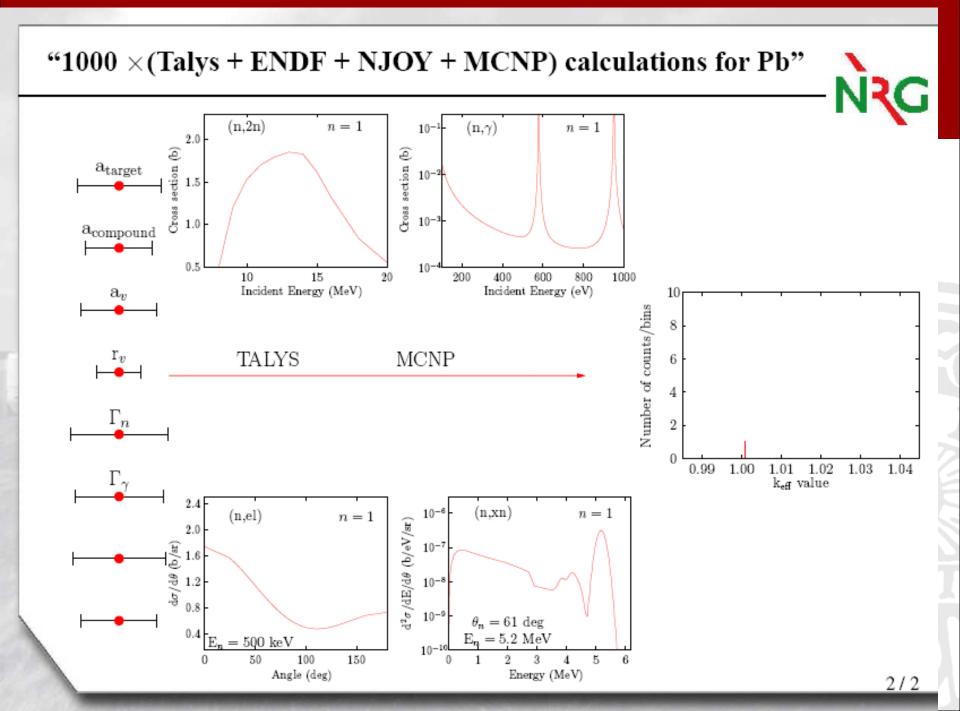


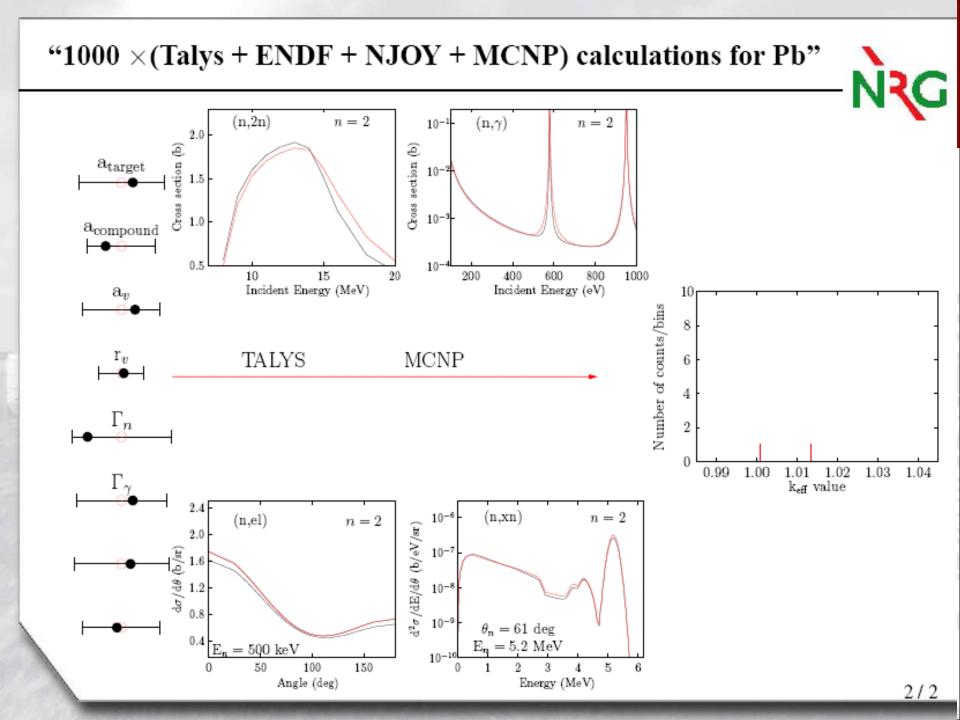


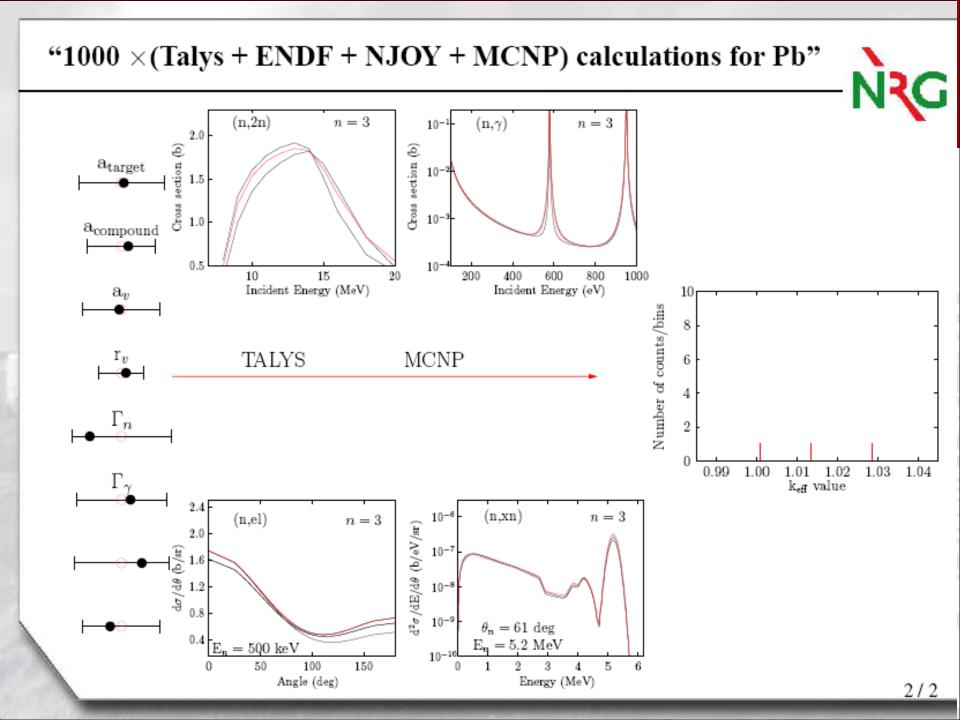
UPPSALA Total Monte Carlo

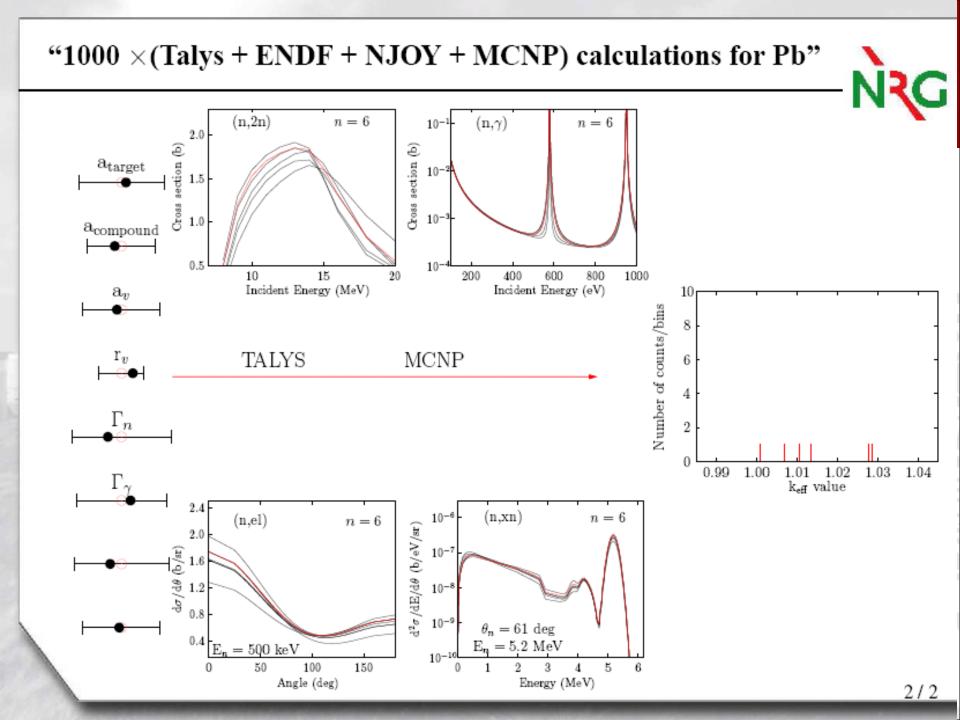


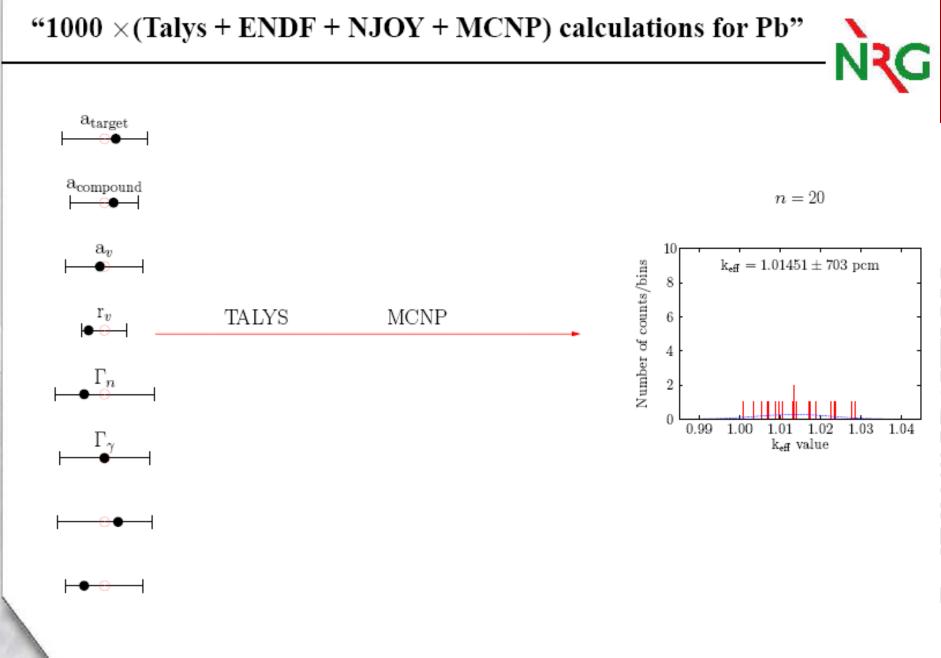
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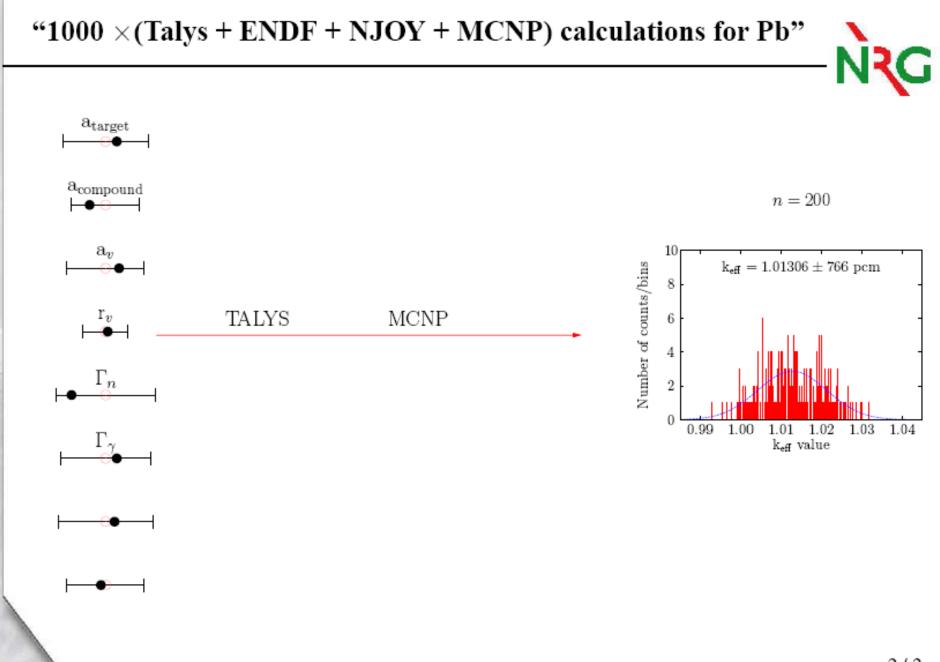


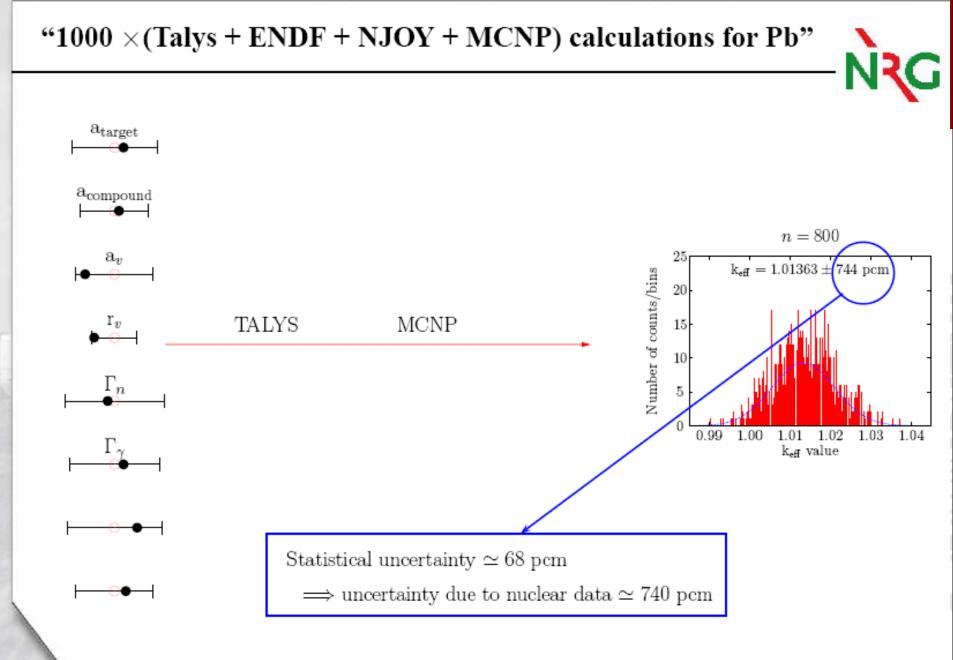














UPPSALA UNIVERSITUSION: Optimized Cu-65 file: differential performance

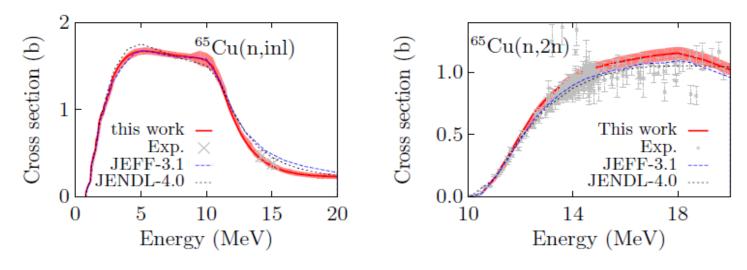
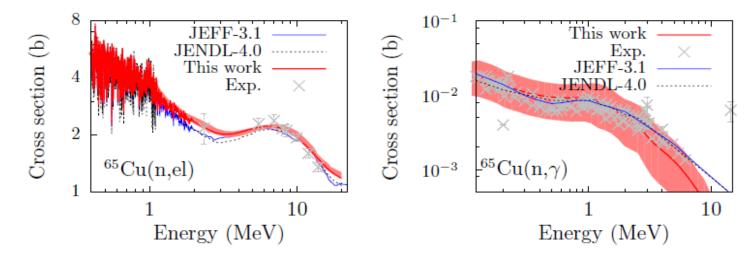


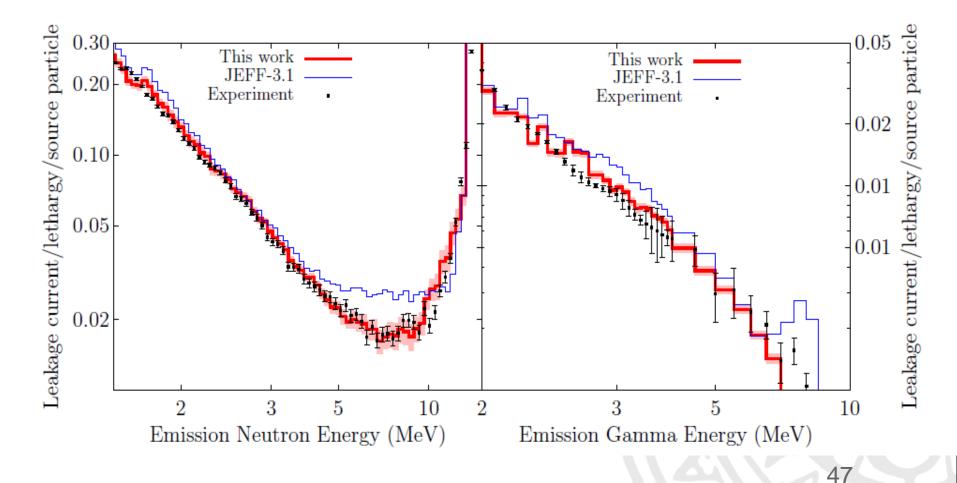
FIG. 12: Inelastic and (n,2n) cross sections for ^{65}Cu .





Fusion: Optimized Cu63,65 file vs Oktavian: integral performance

D. Rochman, A.J. Koning and S.C. van der Marck, <u>Exact nuclear data uncertainty propagation for</u> <u>fusion design</u>", Fusion Engineering and Design 85, 669-682 (2010).





SALA Summary

- TALYS (<u>www.talys.eu</u>) is
 - a user-friendly nuclear model code for reactions involving all light particles up to 200 MeV,
 - used, and validated, by a worldwide community,
 - competitive at all energies
 - Used to analyze nuclear reaction experiments
 - Capable of producing complete nuclear reaction data sets of reasonable to good quality for all kinds of applications