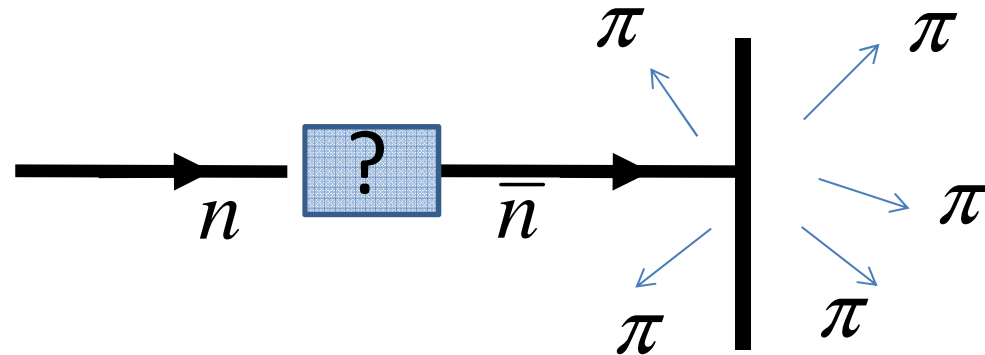


A search for free $n \rightarrow \bar{n}$ oscillations at the ESS



D. Milstead
Stockholm University

Why baryon number violation ?

Why baryon number violation ?

- Baryon number is not a “sacred” quantum number
 - Approximate conservation of BN in SM
 - “Accidental” global symmetry at perturbative level
 - Depends on specific matter content of the SM
 - BNV in SM by non-perturbative processes
 - Sphalerons
 - $B-L$ conserved in SM, not B, L separately.
 - Generic BNV in BSM theories, eg, SUSY.
 - BNV a Sakharov condition for baryogenesis

Why $n \rightarrow \bar{n}$?

$$n \rightarrow \bar{n}$$

- Theory
 - Baryogenesis via *BNV* (Sakharov conditions)
 - SM extensions from TeV mass scales scale-upwards
 - Complementarity with open questions in neutrino physics
- Experiment
 - One of the few means of looking for pure BNV
 - Stringent limit on stability of matter

Neutron oscillations – models

- Back-of-envelope dimensional reasoning:

$$6 \text{ q operator for } \Delta B = 2, \Delta L = 0 \Rightarrow \delta m_{n \rightarrow \bar{n}} = \frac{c \Lambda_{QCD}^6}{M^5} \Rightarrow M \sim 1000 \text{ TeV}$$

- *R*-parity violating supersymmetry
- GUT's , $M \sim 10^{15}$ GeV
- Extra dimensions models
- Post-sphaleron baryogenesis
- etc, etc: [arXiv:1410.1100]

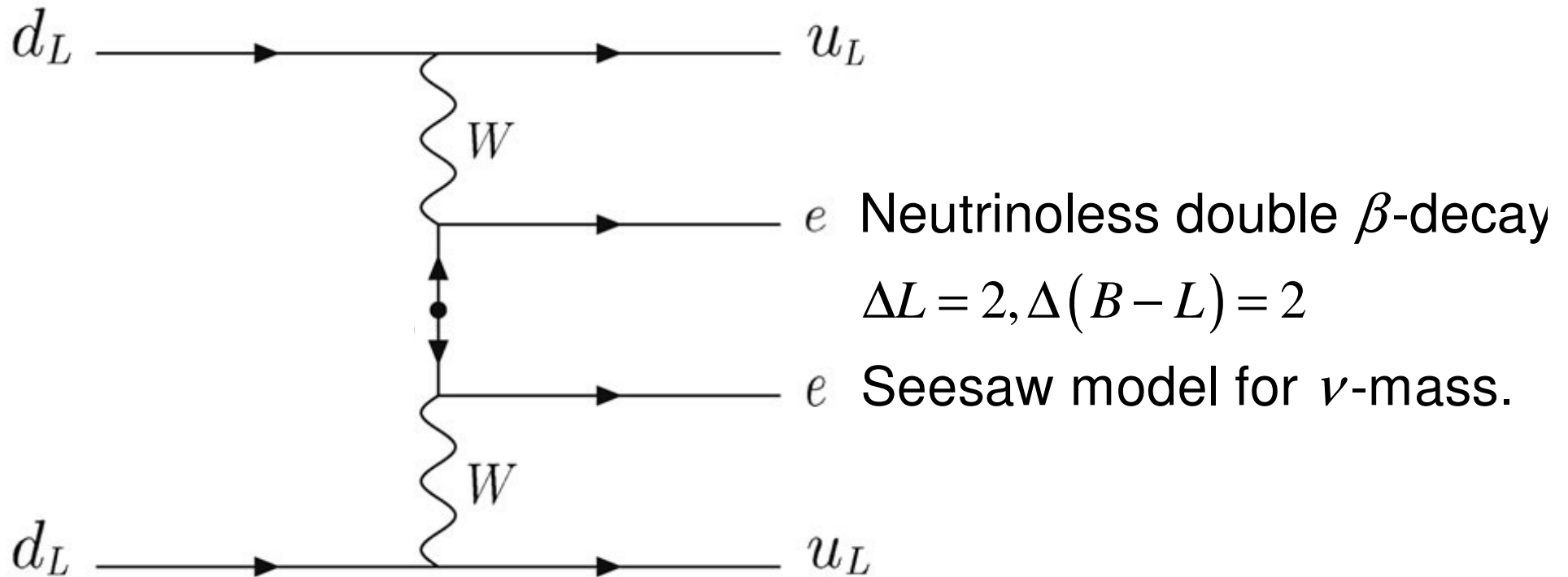
High precision $n \rightarrow \bar{n}$ search

\Rightarrow Scan over wide range of phase space for generic *BNV*

+

\Rightarrow model constraints.

Neutron-neutrino complementarity



Connection to $n \rightarrow \bar{n}$: $\Delta B = 2, \Delta(B - L) = 2$

Large class of seesaw models predict observable $n \rightarrow \bar{n}$.

More generally, both approaches test $B - L$ symmetry.

Neutron oscillations – an experimentalist's view

Hypothesis: BN is weakly violated.

How do we look for BNV ?

Single nucleon decay searches, eg, $p \rightarrow \pi^0 + e^+$?

\Rightarrow L -violation, another (likely weakly) violated quantity.

Decays without leptons, eg, $p \rightarrow \pi + \pi$, impossible due to angular momentum conservation.

$n \rightarrow \bar{n}$ and dinucleon decay searches sensitive to BNV -only processes.

Free $n \rightarrow \bar{n}$ searches \Rightarrow cleanest experimental and theoretical approach.

Previous searches for BNV and $n\bar{n}$ @ESS

Decay mode Partial mean life ($\times 10^{30}$ yrs)

$N \rightarrow e^+ \pi$	> 2000 (n), > 8200 (p)
$N \rightarrow \mu^+ \pi$	> 1000 (n), > 6600 (p)
$N \rightarrow \nu \pi$	> 1100 (n), > 390 (p)
$p \rightarrow e^+ \eta$	> 4200
$p \rightarrow \mu^+ \eta$	> 1300
$n \rightarrow \nu \eta$	> 158
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$p \rightarrow e^+ K^*(892)^0$	> 84
$N \rightarrow \nu K^*(892)$	> 78 (n), > 51 (p)
$p \rightarrow e^+ \pi^+ \pi^-$	> 82
$p \rightarrow e^+ \pi^0 \pi^0$	> 147
$n \rightarrow e^+ \pi^+ \pi^0$	> 52
$p \rightarrow \mu^+ \pi^+ \pi^-$	> 133
$p \rightarrow \mu^+ \pi^0 \pi^0$	> 101
$n \rightarrow \mu^+ \pi^+ \pi^0$	> 74
$n \rightarrow e^+ K^0 \pi^-$	> 18
$n \rightarrow e^- \pi^+$	> 65
$n \rightarrow \mu^- \pi^+$	> 49
$n \rightarrow e^- \rho^+$	> 62
$n \rightarrow \mu^- \rho^+$	> 7
$n \rightarrow e^- K^+$	> 32
$n \rightarrow \mu^- K^+$	> 57
$p \rightarrow e^- \pi^+ \pi^+$	> 30
$n \rightarrow e^- \pi^+ \pi^0$	> 29
$p \rightarrow \mu^- \pi^+ \pi^+$	> 17
$n \rightarrow \mu^- \pi^+ \pi^0$	> 34
$p \rightarrow e^- \pi^+ K^+$	> 75
$p \rightarrow \mu^- \pi^+ K^+$	> 245

(RPP)

$p \rightarrow e^+ \gamma$	> 670
$p \rightarrow \mu^+ \gamma$	> 478
$n \rightarrow \nu \gamma$	> 28
$p \rightarrow e^+ \gamma \gamma$	> 100
$n \rightarrow \nu \gamma \gamma$	> 219
$p \rightarrow e^+ e^+ e^-$	> 793
$p \rightarrow e^+ \mu^+ \mu^-$	> 359
$p \rightarrow e^+ \nu \nu$	> 170
$n \rightarrow e^+ e^- \nu$	> 257
$n \rightarrow \mu^+ e^- \nu$	> 83
$n \rightarrow \mu^+ \mu^- \nu$	> 79
$p \rightarrow \mu^+ e^- e^-$	> 529
$p \rightarrow \mu^+ \mu^+ \mu^-$	> 675
$p \rightarrow \mu^+ \nu \nu$	> 220
$p \rightarrow e^- \mu^+ \mu^+$	> 6
$n \rightarrow 3\nu$	> 0.0005
$N \rightarrow e^+ \text{anything}$	> 0.6 (n, p)
$N \rightarrow \mu^+ \text{anything}$	> 12 (n, p)
$N \rightarrow e^+ \pi^0 \text{anything}$	> 0.6 (n, p)
$pp \rightarrow \pi^+ \pi^-$	> 0.7
$pn \rightarrow \pi^+ \pi^0$	> 2
$nn \rightarrow \pi^+ \pi^-$	> 0.7
$nn \rightarrow \pi^0 \pi^0$	> 3.4
$pp \rightarrow K^+ K^+$	> 170
$pp \rightarrow e^+ e^+$	> 5.8
$pp \rightarrow e^+ \mu^+$	> 3.6
$pp \rightarrow \mu^+ \mu^+$	> 1.7
$pn \rightarrow e^- \bar{\nu}$	> 2.8
$pn \rightarrow \mu^+ \bar{\nu}$	> 1.6
$pn \rightarrow \tau^+ \bar{\nu}_\tau$	> 1.0
$nn \rightarrow \nu_e \bar{\nu}_e$	> 1.4
$nn \rightarrow \nu_\mu \bar{\nu}_\mu$	> 1.4

$\Delta B \neq 0, \Delta L \neq 0$

$\Delta B \neq 0, \Delta L = 0$

Few searches for $\Delta B \neq 0, \Delta L = 0$

Limits on τ_{life} from all searches

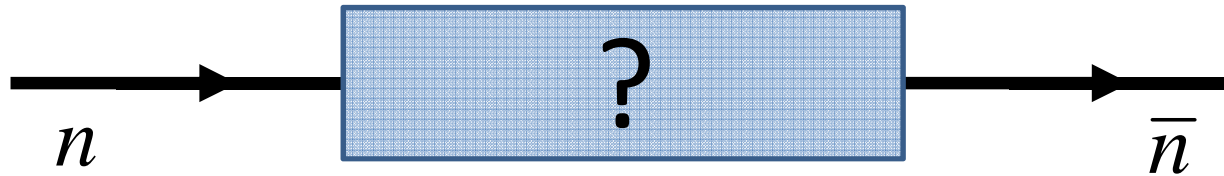
$\sim 10^{30} - 10^{34}$ yrs

New experiment: $\Delta B \neq 0, \Delta L = 0$

τ_{life} sensitivity $\sim 10^{35}$ yrs

Discovery or new stringent limit on stability of matter.

$n \rightarrow \bar{n}$ mixing formalism



$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} n \\ \bar{n} \end{pmatrix} = \begin{pmatrix} E_n & \delta m \\ \delta m & E_{\bar{n}} \end{pmatrix} \begin{pmatrix} n \\ \bar{n} \end{pmatrix}$$

$$\delta m = \langle \bar{n} | H_{eff} | n \rangle < 10^{-29} \text{ MeV} = n\bar{n} \text{ mixing physics}$$

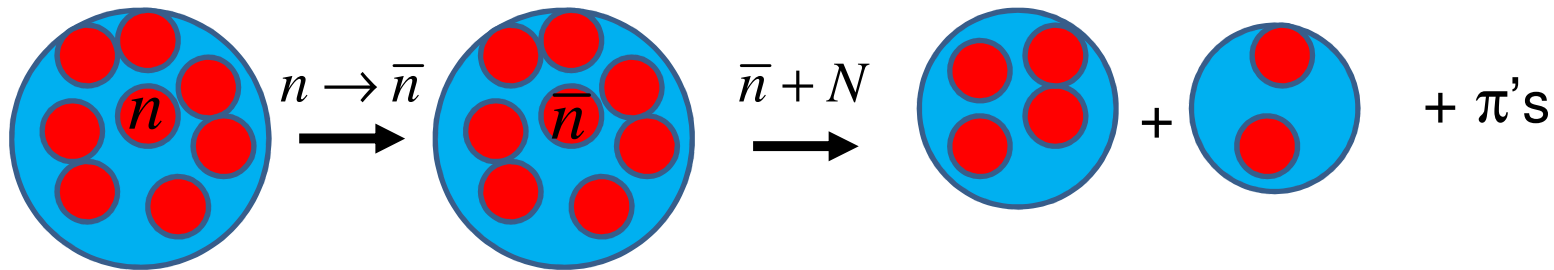
$$P_{n \rightarrow \bar{n}} = \left(\frac{\delta m}{\Delta E} \right)^2 \sin^2 (\Delta E \times t) \quad ; \quad \Delta E = E_n - E_{\bar{n}}$$

Two interesting cases:

- Free neutron oscillation: $\Delta E \times t \ll 1 \Rightarrow P \sim (\delta m \times t)^2$
- Bound neutron oscillation: $\Delta E \times t \gg 1$

Searching with bound neutrons

Nuclear disintegration after neutron oscillation



$$P_{n \rightarrow \bar{n}} = \left(\frac{\delta m}{\Delta E} \right)^2 \sin^2 (\Delta E \times t) ,$$

$$\Delta E \sim 100 \text{ MeV} .$$

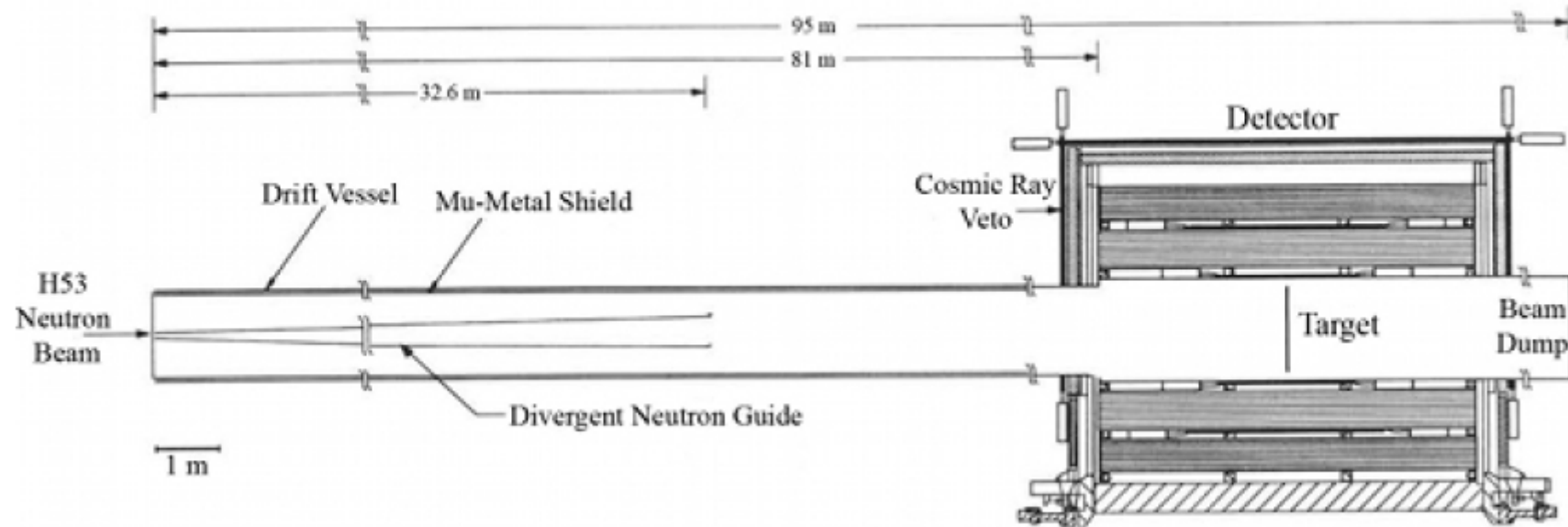
$$\Rightarrow \text{Suppression: } \left(\frac{\delta m}{\Delta E} \right)^2 < 10^{-60}$$

Best current limits (SuperKamiokande) $\Rightarrow \tau_{free} > 2.5 \times 10^8 \text{ s}$

Irreducible bg's prevent large improvements.

Model-dependent (nuclear interactions).

Free neutron search at ILL



Institute Laue-Langevin (Early 1990's).

Cold neutron beam from 58MW reactor.

~ 130 μ m thick carbon target

Signal of at least two tracks with $E > 850$ MeV

0 candidate events, 0 background.

$\Rightarrow \tau_{n \rightarrow \bar{n}} > 0.86 \times 10^8$ s.

The European Spallation Source

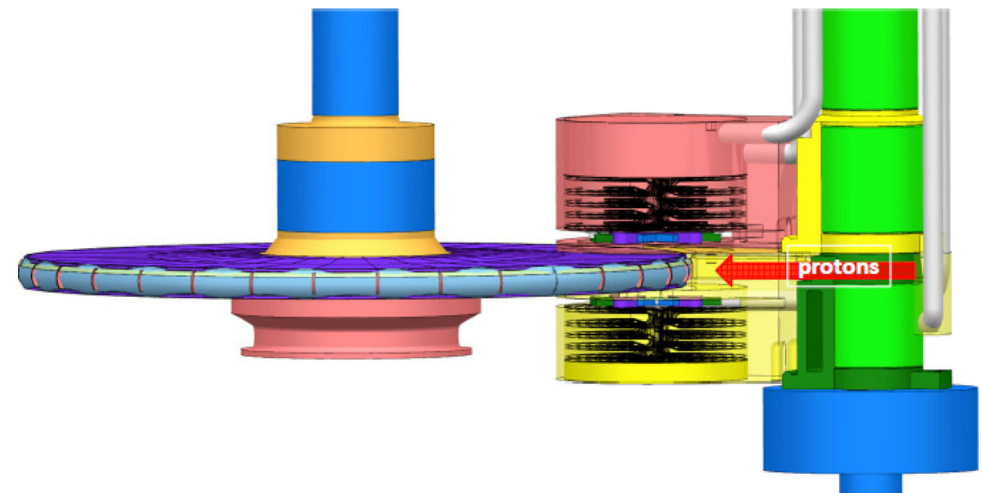
High intensity spallation
neutron source

Multidisciplinary research centre
with 17 European nations
participating.

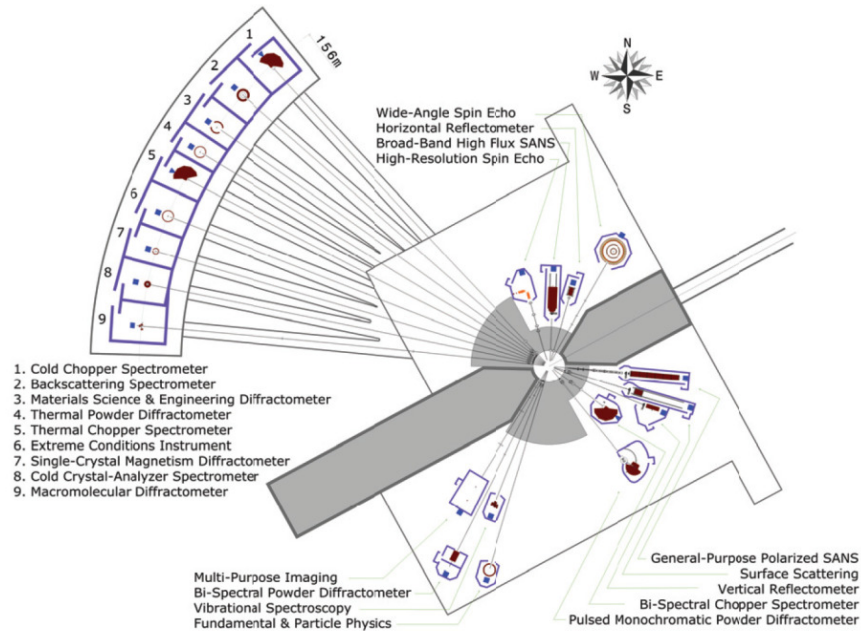
Lund, Sweden.
Start operations in 2019.

2 GeV protons (3ms long pulse,
14 Hz) hit rotating tungsten
target.

Cold neutrons after interaction
with moderators.

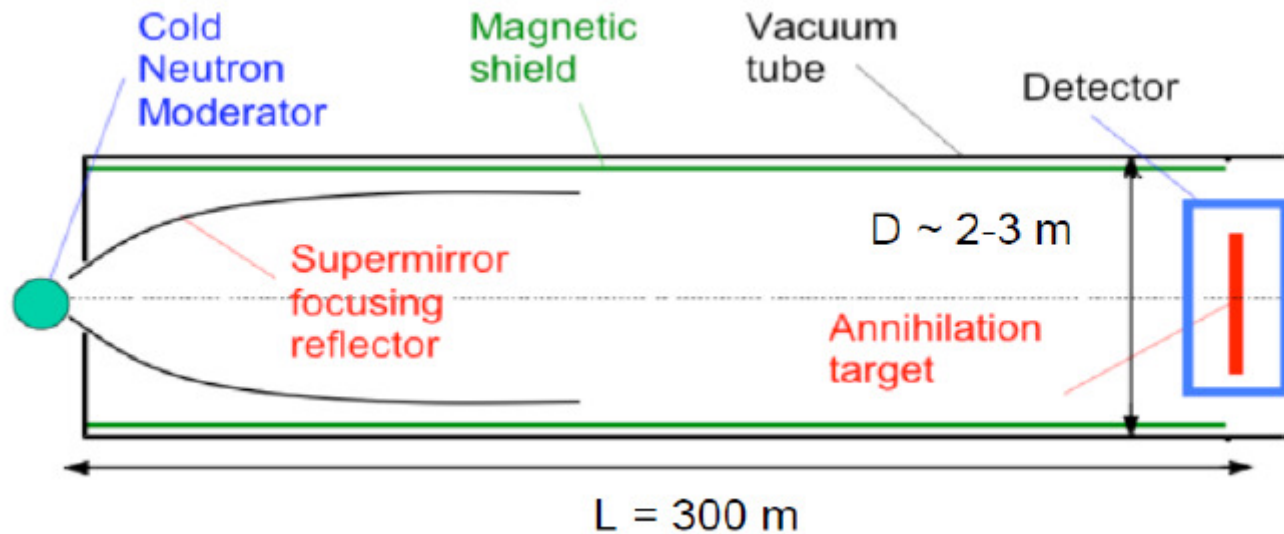


The European Spallation Source



~ 22 instruments/experiments with capability for more.

Overview of the Experiment

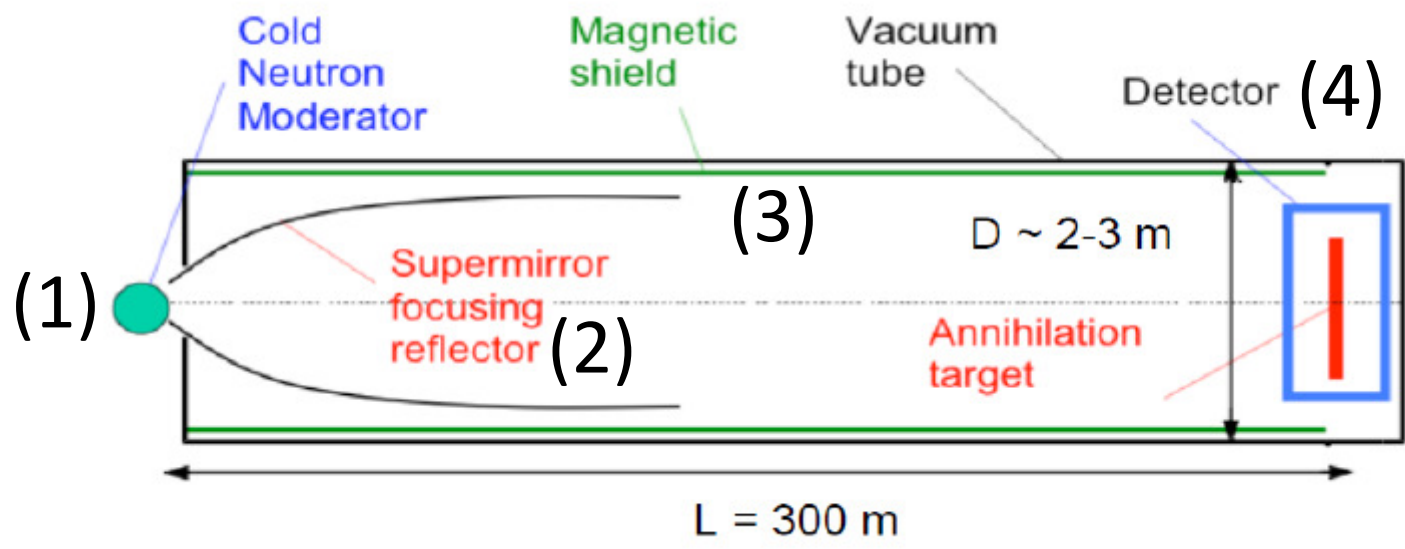


$$\text{Sensitivity} = (\text{free neutron flux at target}) \times P(n \rightarrow \bar{n}) \propto N_n t^2$$

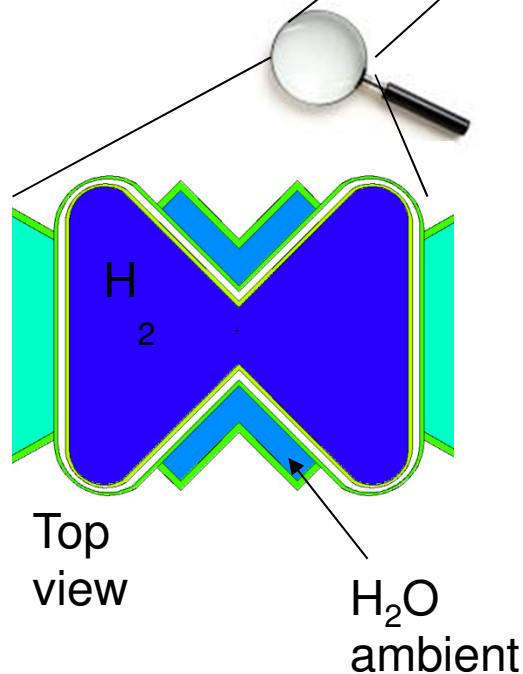
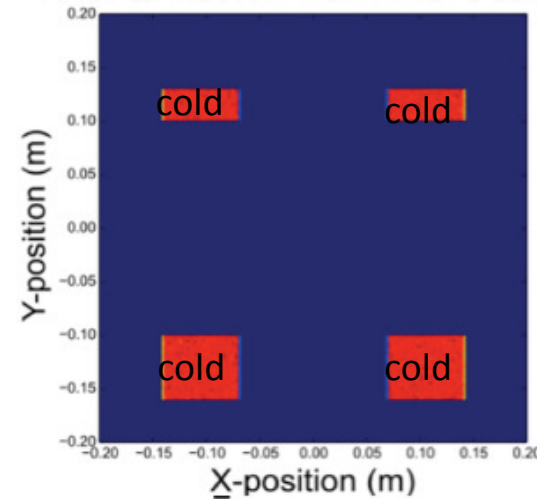
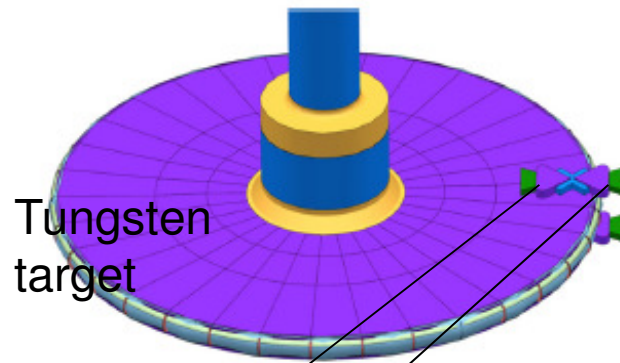
- Cold neutrons ($E < 5$ meV, $v < 1000$ ms $^{-1}$)
- Low neutron emission temperature (50-60 K)
- Supermirror transmission and transit time
- Large beam port option, large solid angle to cold moderator.

Increase in sensitivity for $P_{n\bar{n}} \sim 10^3$ compared to previous experiment (ILL)

- Neutron guiding, larger opening angle, higher flux, particle ID technologies, running time.



Neutronics (1)

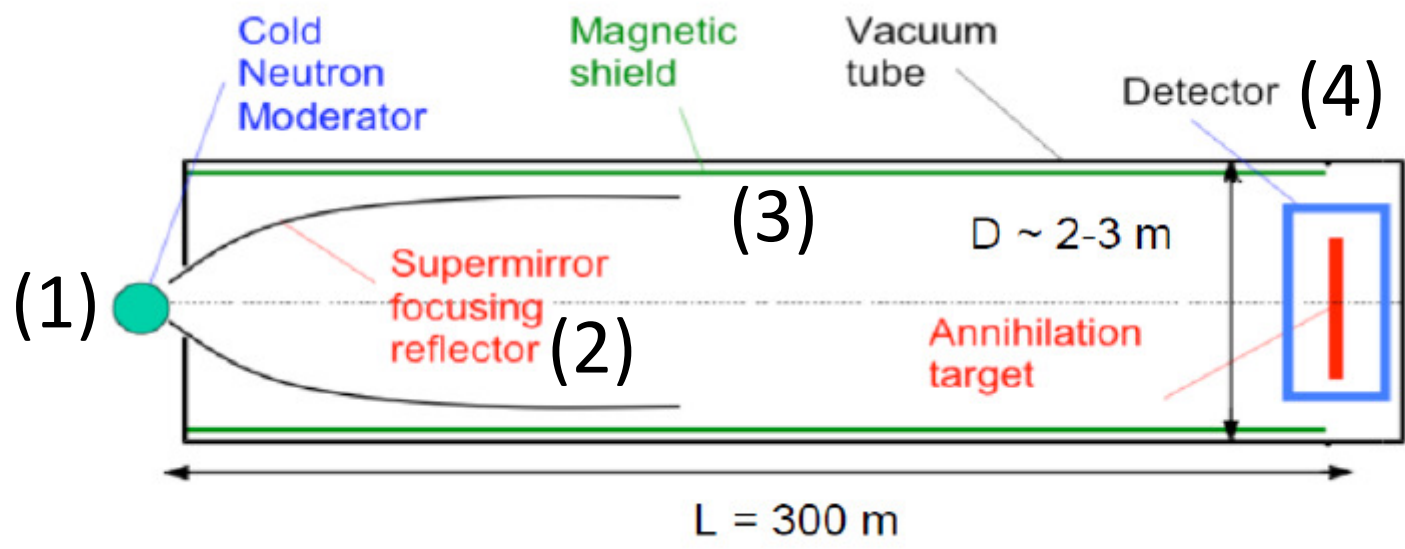


ESS moderators will be of “butterfly” design

- Increase cold yield
- Convenient beam extraction

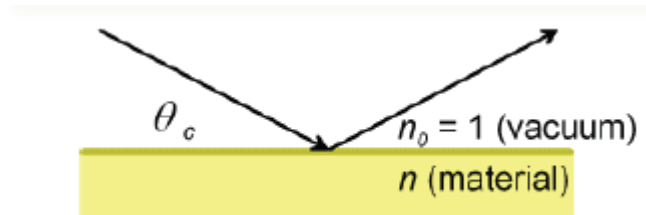
Additional challenge for nbar which could benefit from extracting neutrons from all four visible cold surfaces

- Conventional point-to-point focusing of a cold neutron beam using ellipsoidal mirrors inefficient.
- Ongoing studies on neutron optics



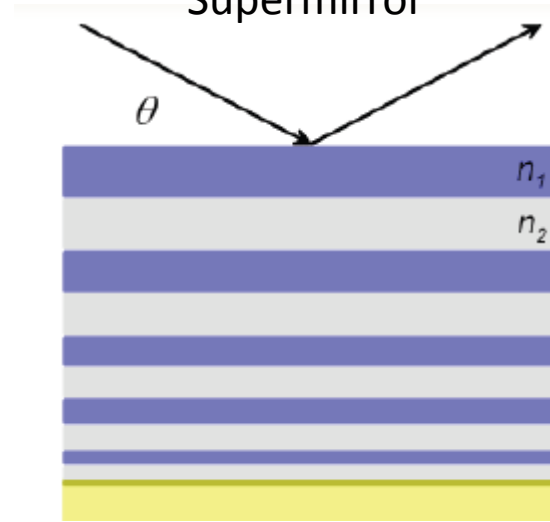
Neutron supermirror

Smooth surface

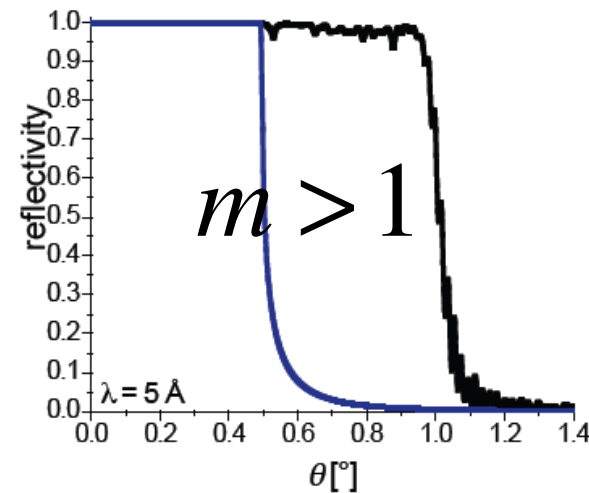
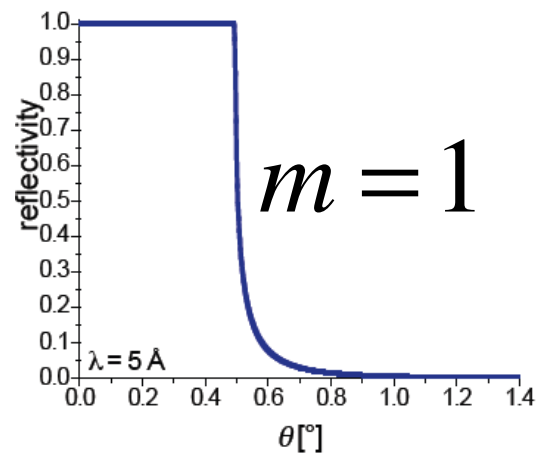


θ_c = Critical angle for total internal reflection

Supermirror

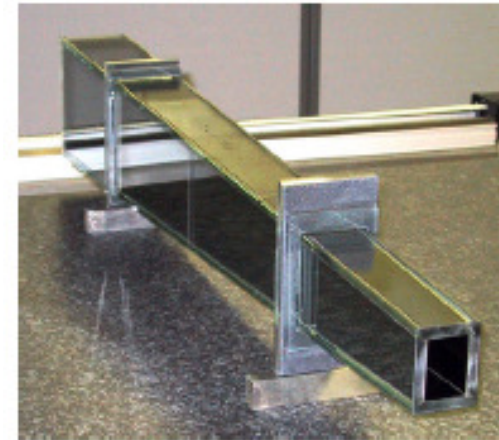
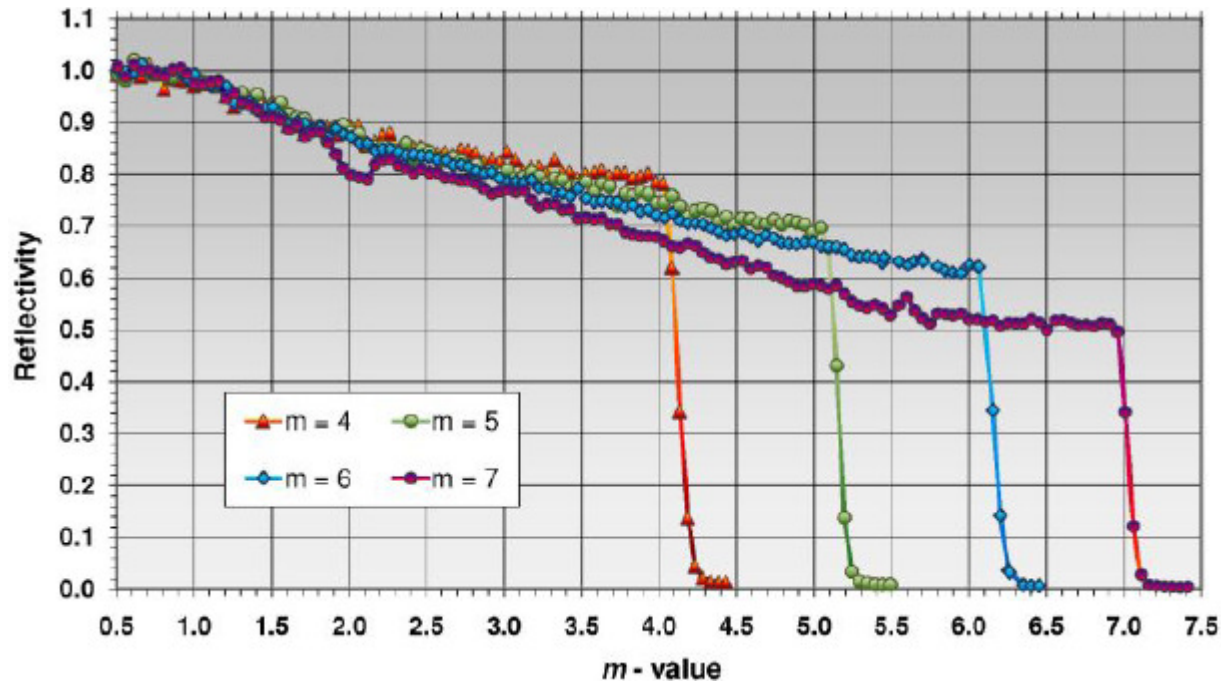


$\theta_c \rightarrow m\theta_C^{Ni}$



Need efficient focusing and minimal interactions
(each interaction "resets the n -clock")

Commercial supermirrors



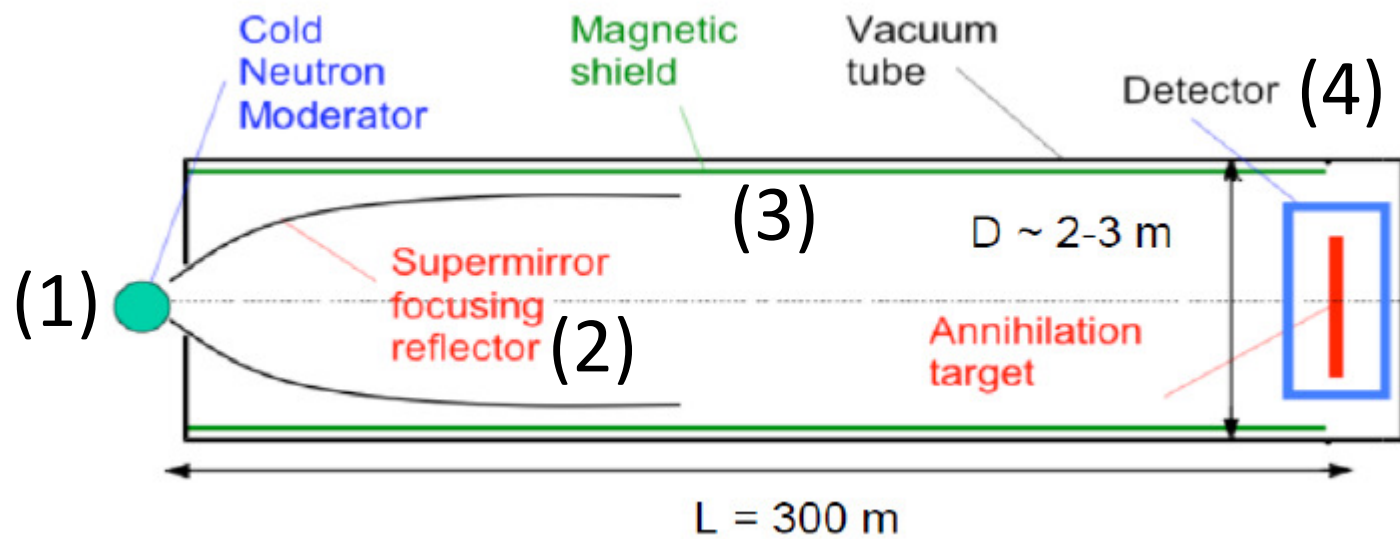
Commercial supermirrors with $m \sim 7$

Acceptance for straight guide $\propto m^2$

ILL experiment used $m \sim 1$ neutron optics.

Increase from use of focusing reflector and optimised mirror arrays.

Crucial contribution to increase of sensitivity wrt ILL.



The need for magnetic shielding

$$\frac{n(\mu \downarrow) \quad \bar{n}(\mu \uparrow)}{B \sim 0} \quad \begin{array}{c} n(\mu \downarrow) \\ \updownarrow 2\vec{\mu} \cdot \vec{B} \\ \bar{n}(\mu \uparrow) \end{array} \quad \uparrow E$$

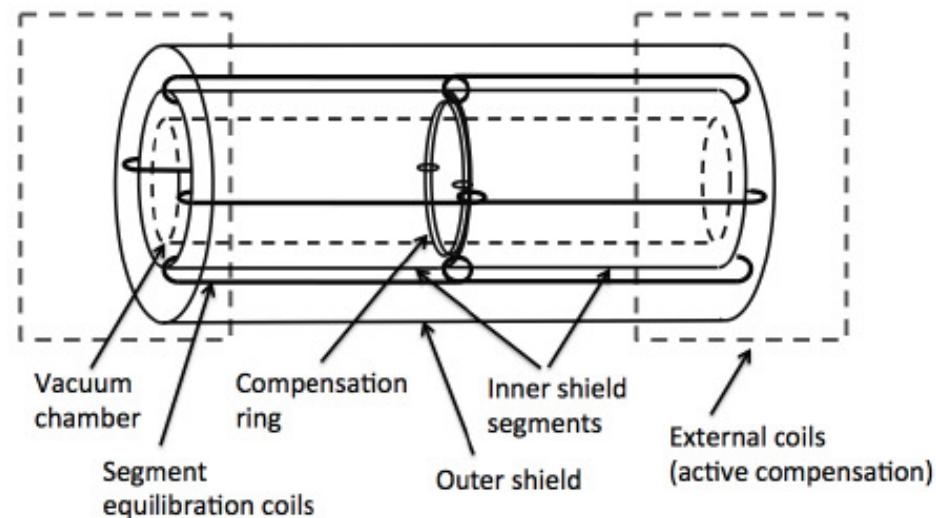
Degeneracy of n, \bar{n} broken in B-field due to dipole interactions: $\Delta E = 2\vec{\mu} \cdot \vec{B}$

Flight time $\leq 1\text{s}$

For quasi-free condition $\Delta E \times t \ll 1$

$\Rightarrow B \leq 5\text{nT}$ and vacuum $\leq 10^{-5} \text{ Pa}$.

Shielding



Magnetic shielding for flight volume

- $B < 5\text{nT}$, $P \sim 10^{-5}\text{mbar}$
- Aluminium vacuum chamber
- Passive magnetic shield from magnetizable alloy
- External coils for active compensation
- Background studied by turning on/off \vec{B} -field.

Maybe shielding isn't needed

PHYSICAL REVIEW D **91**, 096010 (2015)

Phenomenology of $n-\bar{n}$ oscillations revisited

S. Gardner* and E. Jafari

Department of Physics and Astronomy, University of Kentucky, Lexington, Kentucky 40506-0055, USA
(Received 14 August 2014; revised manuscript received 15 February 2015; published 22 May 2015)

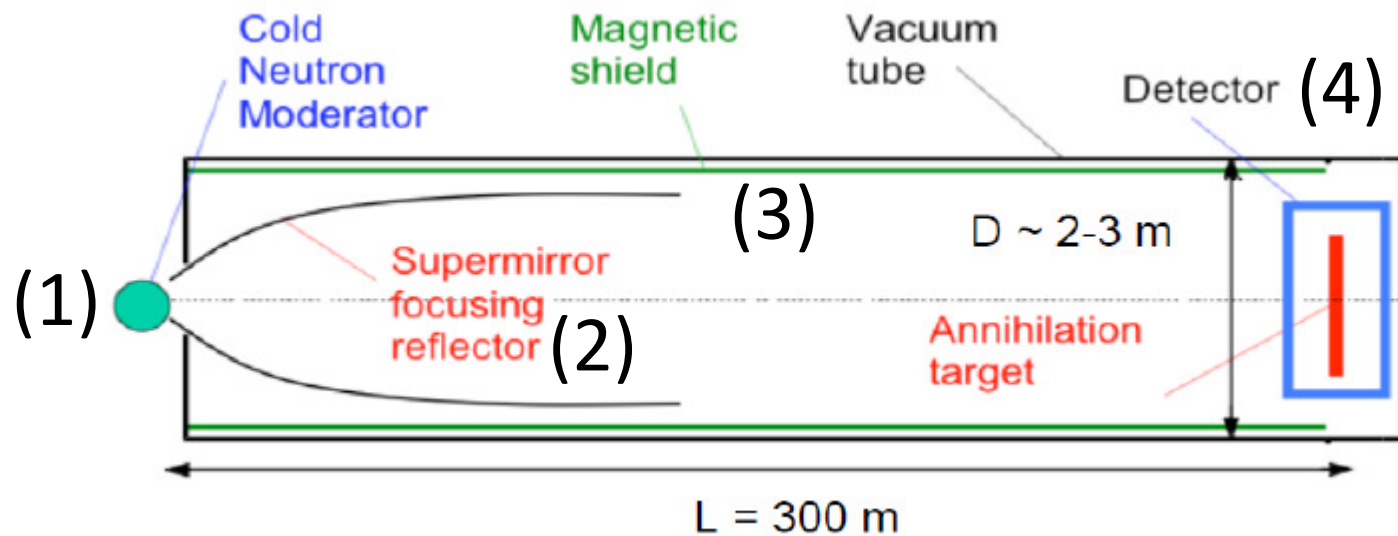
We revisit the phenomenology of $n-\bar{n}$ oscillations in the presence of external magnetic fields, highlighting the role of spin. We show, contrary to long-held belief, that the $n-\bar{n}$ transition rate need not be suppressed, opening new opportunities for its empirical study.

DOI: [10.1103/PhysRevD.91.096010](https://doi.org/10.1103/PhysRevD.91.096010)

PACS numbers: 11.30.Fs, 11.30.Er, 13.40.Em, 14.20.Dh

Interesting discussion in the literature.

Overview of the Experiment

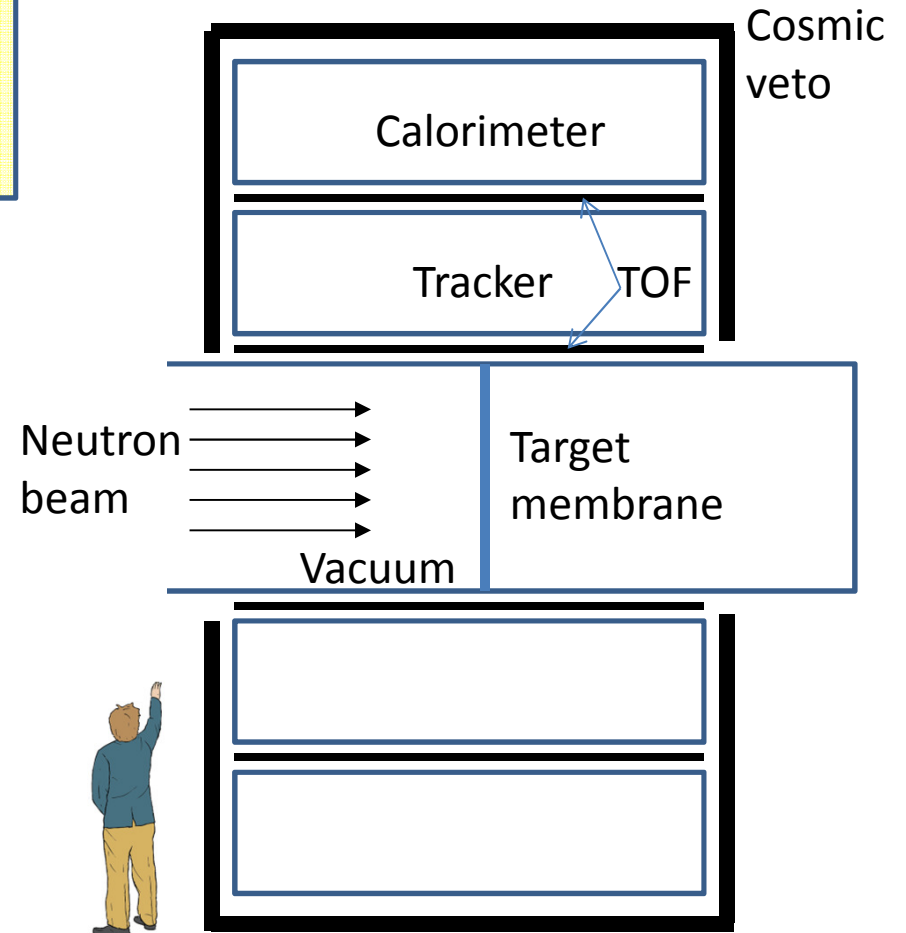


(4) Detector

Expect $n + \bar{n} \rightarrow \sim 5\pi$ at $\sqrt{s} \sim 2$ GeV.

Detector design for high efficiency ($\epsilon > 0.5$) and low bg (~ 0).

- Annihilation target - carbon sheet
- Tracker - vertex reconstruction
- Time-of-flight system
 - scintillators around tracker.
- Calorimeter
 - lead + scintillating and clear fibre.
- Cosmic veto - plastic scintillator pads
- Trigger - Track and cluster algorithms



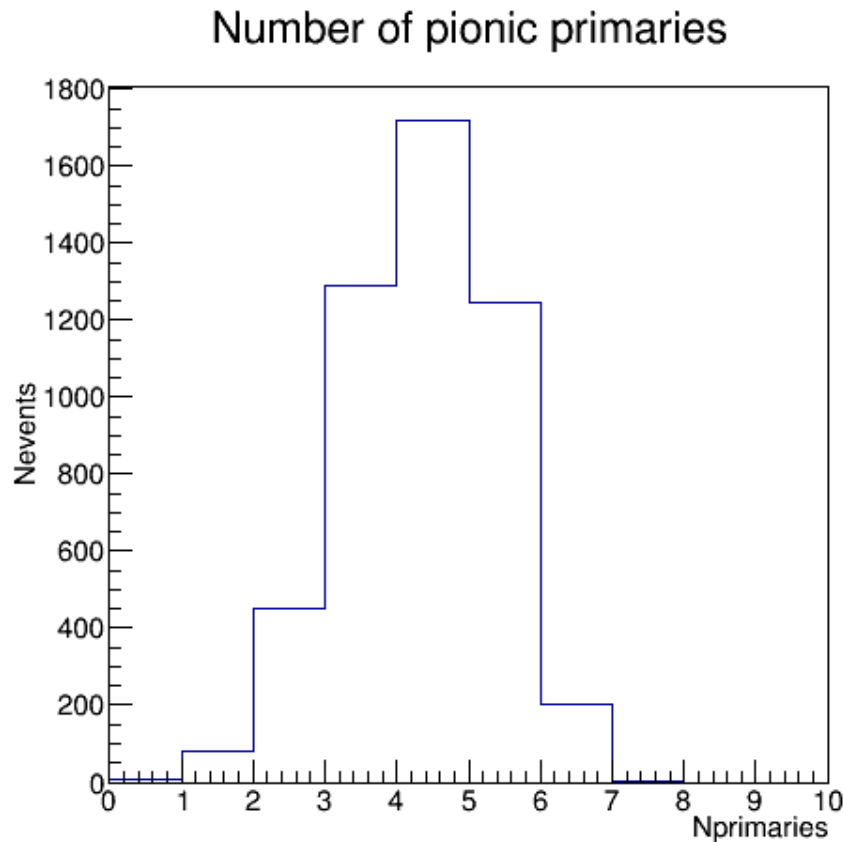
GENIE: NNBar Final State Primaries

Preliminary

Final state list prepared by R. W. Pattie

GENIE-2.0.0: intranuclear propagation based on INTRANUKE

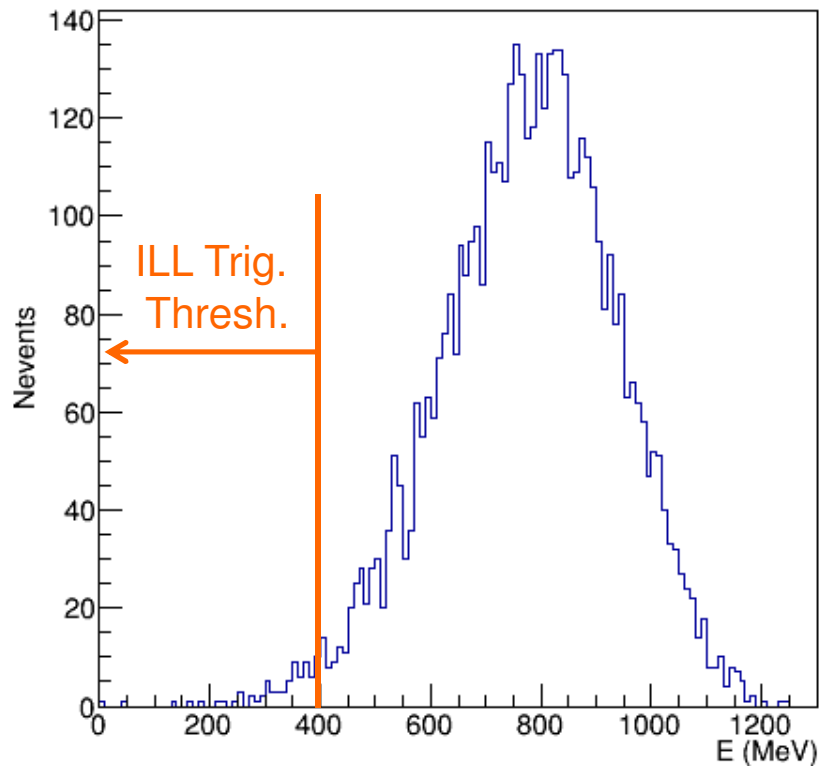
[C.Andreopoulos et al., The GENIE Neutrino Monte Carlo Generator, Nucl.Instrum.Meth.A614:87-104,2010.](#)



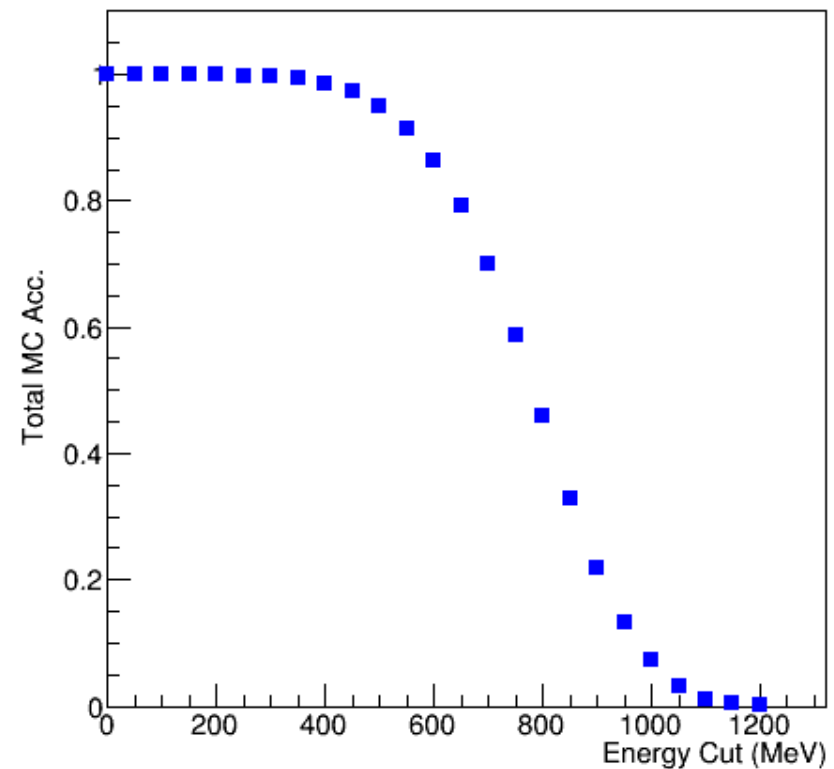
Final State Pionic Mode	Nevents	% Total
$\pi^+\pi^-2\pi^0$	530	10.60%
$2\pi^+\pi^-\pi^0$	486	9.72%
$\pi^+\pi^-\pi^0$	417	8.34%
$2\pi^+\pi^-2\pi^0$	409	8.18%
$\pi^+\pi^-3\pi^0$	329	6.58%
$2\pi^+2\pi^-\pi^0$	315	6.30%
$\pi^+2\pi^0$	290	5.80%
$\pi^+3\pi^0$	219	4.38%
$\pi^+\pi^-\omega$	145	2.90%
$\pi^+\pi^0$	137	2.74%
$\pi^+2\pi^-\pi^0$	132	2.64%
$2\pi^+2\pi^-$	124	2.48%

Energy Threshold Acceptance (Signal)

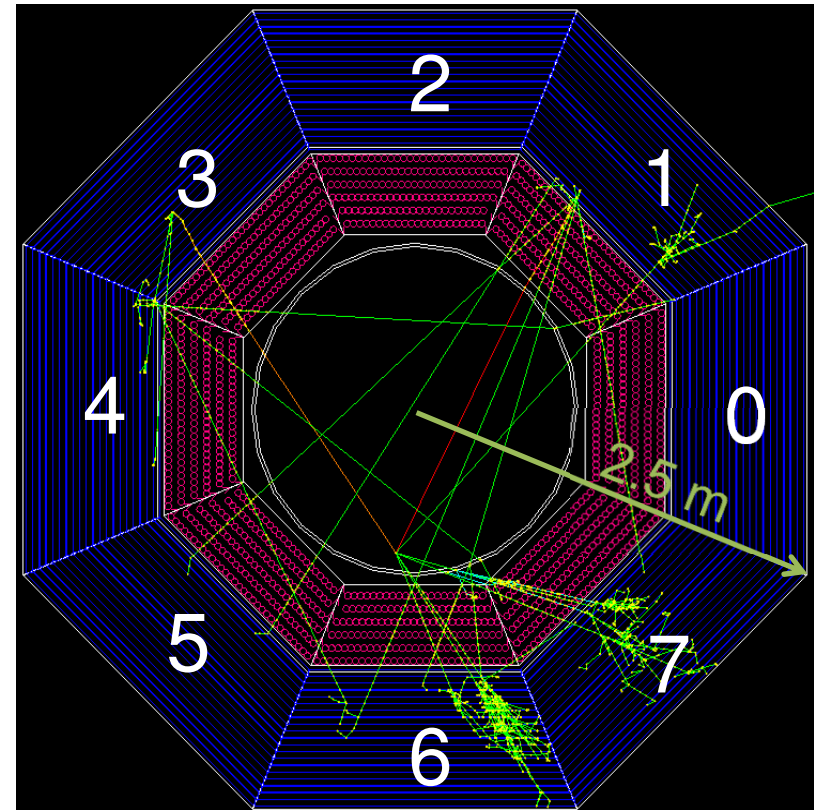
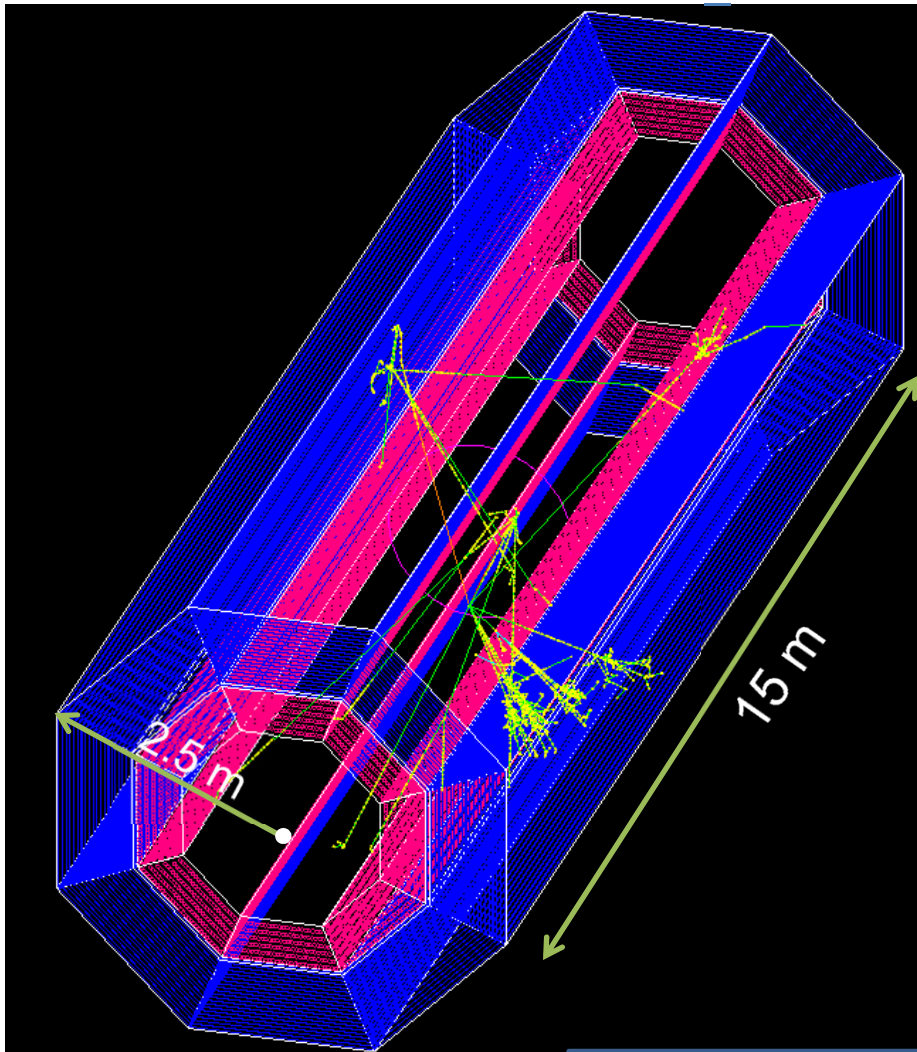
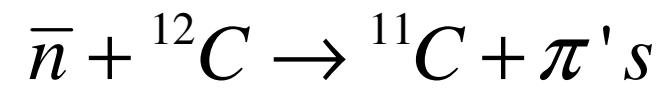
Tot. Energy Dep. in Active Cal. (signal)



MC Acc. vs. Active Cal. Energy Cut (signal)



Event Reconstruction



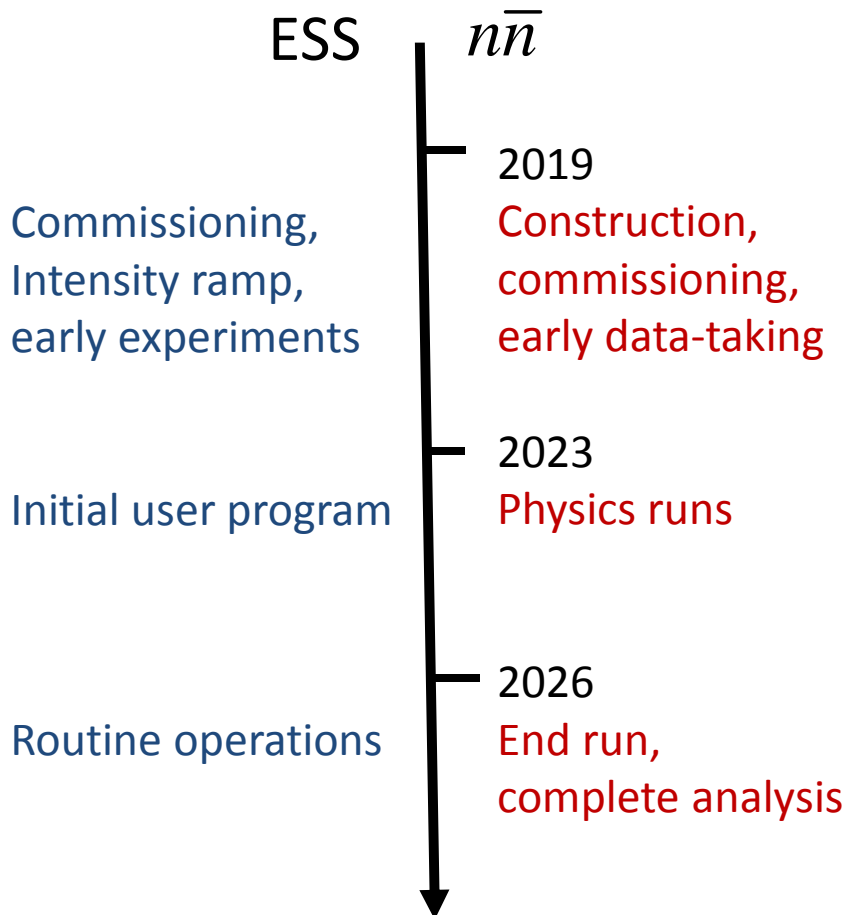
nnbar@ESS in Sweden

- Lund, SU, UU, Chalmers
- Tracking, Calo, Read-out, Geant, BG, interest
- Theory collaboration on LHC-nnbar complementarity (RPV-SUSY)

Collaboration and approximate timescales

Several workshops (CERN, Lund, Gothenburg)
Collaboration formed – interim spokesperson G. Brooijmans
Expression of Interest submitted to ESS.
Signatories from 26 institutes , 8 countries.

More collaborators are welcome!



Neutron-Anti-Neutron Oscillations at ESS

Lund, Feb 18-19, 2015

Neutral particle oscillations have proven to be extremely valuable probes of fundamental physics. Kaon oscillations provided us with our first insight into CP-violation, fast B_s oscillations provided the first indication that the top quark is extremely heavy, B oscillations form the most fertile ground for the continued study of CP-violation, and neutrino oscillations suggest the existence of a new, important energy scale well below the GUT scale. Neutrons oscillating into antineutrons could offer a unique probe of baryon number violation. The construction of the European Spallation Source in Lund, with first beam expected in 2019, together with modern neutron optical techniques, offers an opportunity to conduct an experiment with at least three orders of magnitude improvement in sensitivity to the neutron oscillation probability.

At this workshop the physics case for such an experiment will be discussed, together with the main experimental challenges and possible solutions. We hope the workshop will conclude with the first steps towards the formation of a collaboration to build and perform the experiment.

Organising committee:
G. Brooijmans (Columbia University)
S. Chakraborty (Cockcroft Institute)
R. Holt-Wilson (European Spallation Source)
Y. Kamenkov (University of Tennessee)
E. Hübner (Technical University of Denmark and European Spallation Source)
M. Lindner (European Spallation Source and Lund University)
L. Mepsted (ESS)
M. Mezzetto (INFN Padova)
H. M. Sorenson (Plymouth University)
W. M. Snow (Indiana University)
T. Seldner (Institute Laue-Langevin)
C. Thorne (European Spallation Source)

Register before
19 May on
www.nbar-at-ess.org

ESS EUROPEAN SPALLATION SOURCE CERN

Particle Physics Strategy

European:

h) Experiments studying quark flavour physics, investigating dipole moments, searching for charged-lepton flavour violation and performing other precision measurements at lower energies, such as those with neutrons, muons and antiprotons, may give access to higher energy scales than direct particle production or put fundamental symmetries to the test. They can be based in national laboratories, with a moderate cost and smaller collaborations. *Experiments in Europe with unique reach should be supported, as well as participation in experiments in other regions of the world.*

US P5 report:

- With a mix of large, medium, and small projects, important physics results will be produced continuously throughout the twenty-year P5 timeframe. In our budget exercises, we maintained a small projects portfolio to preserve budgetary space for a set of projects whose costs individually are not large enough to come under direct P5 review but which are of great importance to the field. This is in addition to the aforementioned small neutrino experiments portfolio, which is intended to be integrated into a coherent overall neutrino program.

Consensus in the field is to pursue experiments with unique capabilities and physics reach.

Summary

- The search for neutron-antineutron oscillations addresses open questions in modern physics.
- An experiment at the ESS offers a new opportunity to extend sensitivity to neutron oscillation probability by several orders of magnitude and set a new limit on the stability of matter.
- Collaboration formed and EOI submitted
- Provisional schedule made.

Brightness		≥ 1
Moderator Temperature	<TOF> driven by colder neutrons, \sim quadratic (t^2)	≥ 1
Moderator Area	Needs large aperture	2
Angular Acceptance	2D, so quadratic sensitivity	40
Length	Scale with t^2 , so L^2	5
Run Time	ILL run was 1 year	3
Total		≥ 1000

x 1000 in probability, reach $\tau \sim 2-3 \times 10^9$ s
(simulations with various moderator options underway)

Potential gains

Factor	Gain wrt ILL
Brightness	≥ 1
Moderator temperature	≥ 1
Moderator area	2
Angular acceptance/neutron transmission	40
Length	5
Run time	3
Total	≥ 1000

Baryon number violation searches

Decay mode Partial mean life ($\times 10^{30}$ yrs)

Few searches for
 $\Delta B \neq 0, \Delta L = 0$

τ limits $\sim 10^{30} - 10^{34}$ yrs

τ limit from new
 experiment

$\sim 10^{35}$ yrs

Decay mode	Partial mean life ($\times 10^{30}$ yrs)
$N \rightarrow e^+ \pi$	$> 2000 (n), > 8200 (p)$
$N \rightarrow \mu^+ \pi$	$> 1000 (n), > 6600 (p)$
$N \rightarrow \nu \pi$	$> 1100 (n), > 390 (p)$
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$p \rightarrow e^+ K^*(892)^0$	> 84
$N \rightarrow \nu K^*(892)$	$> 78 (n), > 51 (p)$
$p \rightarrow e^+ \pi^+ \pi^-$	> 82
$p \rightarrow e^+ \pi^0 \pi^0$	> 147
$n \rightarrow e^+ \pi^- \pi^0$	> 52
$p \rightarrow \mu^+ \pi^+ \pi^-$	> 133
$p \rightarrow \mu^+ \pi^0 \pi^0$	> 101
$n \rightarrow \mu^+ \pi^- \pi^0$	> 74
$n \rightarrow e^+ K^0 \pi^-$	> 18
$n \rightarrow e^- \pi^+$	> 65
$n \rightarrow \mu^- \pi^+$	> 49
$n \rightarrow e^- \rho^+$	> 62
$n \rightarrow \mu^- \rho^+$	> 7
$n \rightarrow e^- K^+$	> 32
$n \rightarrow \mu^- K^+$	> 57
$p \rightarrow e^- \pi^+ \pi^+$	> 30
$n \rightarrow e^- \pi^+ \pi^0$	> 29
$p \rightarrow \mu^- \pi^+ \pi^+$	> 17
$n \rightarrow \mu^- \pi^+ \pi^0$	> 34
$p \rightarrow e^- \pi^+ K^+$	> 75
$p \rightarrow \mu^- \pi^+ K^+$	> 245

(RPP)

$p \rightarrow e^+ \gamma$	> 670
$p \rightarrow \mu^+ \gamma$	> 478
$n \rightarrow \nu \gamma$	> 28
$p \rightarrow e^+ \gamma \gamma$	> 100
$n \rightarrow \nu \gamma \gamma$	> 219
$p \rightarrow e^+ e^+ e^-$	> 793
$p \rightarrow e^+ \mu^+ \mu^-$	> 359
$p \rightarrow e^+ \nu \nu$	> 170
$n \rightarrow e^+ e^- \nu$	> 257
$n \rightarrow \mu^+ e^- \nu$	> 83
$n \rightarrow \mu^+ \mu^- \nu$	> 79
$p \rightarrow \mu^+ e^+ e^-$	> 529
$p \rightarrow \mu^+ \mu^+ \mu^-$	> 675
$p \rightarrow \mu^+ \nu \nu$	> 220
$p \rightarrow e^- \mu^+ \mu^+$	> 6
$n \rightarrow 3\nu$	> 0.0005
$N \rightarrow e^+ \text{anything}$	$> 0.6 (n, p)$
$N \rightarrow \mu^+ \text{anything}$	$> 12 (n, p)$
$N \rightarrow e^+ \pi^0 \text{anything}$	$> 0.6 (n, p)$
$pp \rightarrow \pi^+ \pi^+$	> 0.7
$pn \rightarrow \pi^+ \pi^0$	> 2
$nn \rightarrow \pi^+ \pi^-$	> 0.7
$nn \rightarrow \pi^0 \pi^0$	> 3.4
$pp \rightarrow K^+ K^+$	> 170
$pp \rightarrow e^+ e^+$	> 5.8
$pp \rightarrow e^+ \mu^+$	> 3.6
$pp \rightarrow \mu^+ \mu^+$	> 1.7
$pn \rightarrow e^+ \bar{\nu}$	> 2.8
$pn \rightarrow \mu^+ \bar{\nu}$	> 1.6
$pn \rightarrow \tau^+ \bar{\nu}_\tau$	> 1.0
$nn \rightarrow \nu_e \bar{\nu}_e$	> 1.4
$nn \rightarrow \nu_\mu \bar{\nu}_\mu$	> 1.4

$\Delta B \neq 0, \Delta L \neq 0$

$\Delta B \neq 0, \Delta L = 0$

BNV searches

RPP

BARYON NUMBER

$\Gamma(Z \rightarrow p e)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-6}$, CL = 95%
$\Gamma(Z \rightarrow p \mu)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-6}$, CL = 95%
$\Gamma(\tau^- \rightarrow \bar{p} \gamma)/\Gamma_{\text{total}}$	$<3.5 \times 10^{-6}$, CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \pi^0)/\Gamma_{\text{total}}$	$<1.5 \times 10^{-5}$, CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} 2\pi^0)/\Gamma_{\text{total}}$	$<3.3 \times 10^{-5}$, CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \eta)/\Gamma_{\text{total}}$	$<8.9 \times 10^{-6}$, CL = 90%
$\Gamma(\tau^- \rightarrow \bar{p} \pi^0 \eta)/\Gamma_{\text{total}}$	$<2.7 \times 10^{-5}$, CL = 90%
$\Gamma(\tau^- \rightarrow \Lambda \pi^-)/\Gamma_{\text{total}}$	$<7.2 \times 10^{-8}$, CL = 90%
$\Gamma(\tau^- \rightarrow \bar{\Lambda} \pi^-)/\Gamma_{\text{total}}$	$<1.4 \times 10^{-7}$, CL = 90%
$\Gamma(D^0 \rightarrow p e^-)/\Gamma_{\text{total}}$	[s] $<1.0 \times 10^{-5}$, CL = 90%
$\Gamma(D^0 \rightarrow \bar{p} e^+)/\Gamma_{\text{total}}$	[t] $<1.1 \times 10^{-5}$, CL = 90%
$\Gamma(B^+ \rightarrow \Lambda^0 \mu^+)/\Gamma_{\text{total}}$	$<6 \times 10^{-8}$, CL = 90%
$\Gamma(B^+ \rightarrow \Lambda^0 e^+)/\Gamma_{\text{total}}$	$<3.2 \times 10^{-8}$, CL = 90%
$\Gamma(B^+ \rightarrow \bar{\Lambda}^0 \mu^+)/\Gamma_{\text{total}}$	$<6 \times 10^{-8}$, CL = 90%
$\Gamma(B^+ \rightarrow \bar{\Lambda}^0 e^+)/\Gamma_{\text{total}}$	$<8 \times 10^{-8}$, CL = 90%
$\Gamma(B^0 \rightarrow \Lambda_c^+ \mu^-)/\Gamma_{\text{total}}$	$<1.8 \times 10^{-6}$, CL = 90%
$\Gamma(B^0 \rightarrow \Lambda_c^+ e^-)/\Gamma_{\text{total}}$	$<5 \times 10^{-6}$, CL = 90%

p mean life

[u] $>2.1 \times 10^{29}$ years, CL = 90%

A few examples of proton or bound neutron decay follow. For limits on many other nucleon decay channels, see the Baryon Summary Table.

$\tau(N \rightarrow e^+ \pi)$	$> 2000 (n), > 8200 (p) \times 10^{30}$ years, CL = 90%
$\tau(N \rightarrow \mu^+ \pi)$	$> 1000 (n), > 6600 (p) \times 10^{30}$ years, CL = 90%
$\tau(N \rightarrow e^+ K)$	$> 17 (n), > 1000 (p) \times 10^{30}$ years, CL = 90%
$\tau(N \rightarrow \mu^+ K)$	$> 26 (n), > 1600 (p) \times 10^{30}$ years, CL = 90%

limit on $n\bar{n}$ oscillations (free n) $>0.86 \times 10^8$ s, CL = 90%

limit on $n\bar{n}$ oscillations (bound n) [v] $>1.3 \times 10^8$ s, CL = 90%

L and B violated

B violated

Poor experimental coverage of "pure" B violation tests