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# Based on the work with Yibing Zhang arXiv: 1705.09500

#### Table of Contents

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- Motivation:
  - Puzzles in neutrino physics
  - Links between NSIs and neutrino oscillation
  - Status of the MOMENT proposal
- Remarks on neutrino oscillation channels with CC-NSIs
- Impact on precision measurements from CC-NSIs
- Correlations and constraints of CC-NSIs
- Summary

### Puzzles in neutrino physics





• Non-unitary neutrino mixing?



## New physics beyond SM: new particles, new couplings, new phenomenon...

- Flavor violating interactions with neutrinos:  $\nu_{\alpha}f \rightarrow \nu_{\beta}f, l_{\alpha}^{-} \rightarrow \nu_{\beta}e^{-}\bar{\nu}_{e}\cdots$
- 4-fermion vertices:  $L_{\text{eff}} = 2\sqrt{2} G_F \left(\epsilon^{L/R}\right)^{\alpha\gamma}_{\beta\delta} \left(\bar{\nu}^{\beta}\gamma^{\rho} P_L \nu_{\alpha}\right) \left(\bar{\ell}^{\delta}\gamma^{\rho} P_{L/R} \ell_{\gamma}\right)$



NSI happens to neutrino propagation in matter



NSI at neutrino productions

#### **Current bounds on NSIs**



Constraints by experiments with neutrinos and charged leptons (Davidson et al., 2003):

$$\begin{bmatrix} -0.9 < \varepsilon_{ee} < 0.75 & |\varepsilon_{e\mu}| \lesssim 3.8 \times 10^{-4} & |\varepsilon_{e\tau}| \lesssim 0.25 \\ -0.05 < \varepsilon_{\mu\mu} < 0.08 & |\varepsilon_{\mu\tau}| \lesssim 0.25 \\ |\varepsilon_{\tau\tau}| \lesssim 0.4 \end{bmatrix}$$

Constraints including loops (Biggio, Blennow, Fernández-Martínez, 2009):

Model independent bound for  $\mathcal{E}_{e\mu}$  increases by a factor of 10^3!

$$|\epsilon^{\mu e}| < \begin{pmatrix} 0.025 & 0.03 & 0.03 \\ 0.025 & 0.03 & 0.03 \\ 0.025 & 0.03 & 0.03 \end{pmatrix} \quad 90\% \text{C.L.} \qquad |\epsilon^{ud}| < \begin{pmatrix} 0.041 & 0.025 & 0.041 \\ 0.026 & 0.078 & 0.013 \\ 0.12 & 0.013 & 0.13 \end{pmatrix} \quad 90\% \text{C.L.}$$

Chinese proposal: MuOn-decay Medium baseline NeuTrino beam experiment (MOMENT)

- Neutrino flux: 200-300 MeV
- Baseline: 150 km
- Proton beam: 1.5-2.0 GeV
- Intensity of proton beam: 10 mA
- **POT/y:** 1.1\*e+24
- Neutrino flavors:
- $+ : \mu^- \to e^- + \bar{\nu}_e + \nu_\mu$  $: \mu^+ \to e^+ + \nu_e + \bar{\nu}_\mu$

The proposal sells the design of accelerators with plenty of room for imagination of detectors.

- Which type of detector?
- What kind of physics can we do?



v / m<sup>2</sup> / 0.02 GeV/ 200 days 01 17 19

10<sup>13</sup>

1012

10<sup>1</sup>

Proposal of neutrino detectors for MOMENT



- In MOMENT, neutrino flux peaked at low energies → require a massive detector to compensate for the small cross section.
  - **Good flavor and charge identifications are needed!**  $N_{v}(E) \sim \Phi_{v}(E) \times \sigma_{v}(E) \times target$ v flux detector (# neutrinos) (# targets) depends on your v source make this large! make this large! v cross section tiny (~10<sup>-38</sup> cm<sup>2</sup>)  $\sigma_{v}^{tot} \sim E_{v}$ go to higher energies this is how we Charged Current (CC) detected neutrinos neutrino in in the first place W+ charged lepton out - flavor of outgoing lepton "tags"  $\begin{array}{ccc} \nu_{e} \rightarrow e^{-} & \overline{\nu_{e}} \rightarrow e^{+} \\ \nu_{\mu} \rightarrow \mu^{-} & \overline{\nu_{\mu}} \rightarrow \mu^{+} \\ \nu_{\tau} \rightarrow \tau^{-} & \overline{\nu_{\tau}} \rightarrow \tau^{+} \end{array}$ q flavor of incoming neutrino - charge of outgoing lepton determines whether v or anti-v

Proposal of neutrino detectors for MOMENT



- In MOMENT, neutrino flux peaked at low energies → require a massive detector to compensate for the small cross section.
- Good flavor and charge identifications are needed!  $N_{\nu}(E) \sim \Phi_{\nu}(E) \times \sigma_{\nu}(E) \times target$  $V_{\rho} + n \rightarrow p + e$ v flux detector  $\overline{v}_{\mu} + p \rightarrow n + \mu^+$ (# neutrinos) (# targets) depends on your v source make this large! make this large! v cross section  $\overline{V}_{e} + p \rightarrow n + e^{+}$ tiny (~10<sup>-38</sup> cm<sup>2</sup>)  $\sigma_{v}^{tot} \sim E_{v}$  $v_{\mu} + n \rightarrow p + \mu^{-}$ go to higher energies
  - Choose the Water Cherenkov detector due to its excellent flavor identifications and high efficiencies at the low-energy range.
- Distinguishing neutrinos and anti-neutrinos solved by doping the water with Gd (at the 0.1-0.2% level)
  Ref: hep-ph/0309300

Overview of physics study at MOMENT



• First standard neutrino oscillation physics study performed by Pilar, Matthias and Erique in arXiv:1511.02859

→ CP discovery reach in a complementarity with T2K&NOvA.

• NC-NSIs in matter considered by Pouya and Yasaman in arXiv: 1602.07099





*Ref: arXiv:1511.02859*  Overview of physics study at MOMENT



• First standard neutrino oscillation physics study performed by Pilar, Matthias and Erique in arXiv:1511.02859

→ CP discovery reach in a complementarity with T2K&NOvA.

- NC-NSIs in matter considered by Pouya and Yasaman in arXiv: 1602.07099
  - $\rightarrow$  Degeneracy induced by NSIs and measurements of CPV.



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#### Formalism of neutrino oscillation with NSIs





Ref: Slides borrowed from Tommy Ohlsson



#### Follow the same strategy and extend to 3 flavors + CC-NSIs:

$$\begin{split} P_{\bar{\nu}_{e}^{s} \to \bar{\nu}_{e}^{d}}^{ND} &\approx 1 + 2|\epsilon_{ee}^{s}|\cos\phi_{ee}^{s} + 2|\epsilon_{ee}^{d}|\cos\phi_{ee}^{d} \\ P_{\bar{\nu}_{\mu}^{s} \to \bar{\nu}_{\mu}^{d}}^{ND} &\approx 1 + 2|\epsilon_{\mu\mu}^{s}|\cos\phi_{\mu\mu}^{s} + 2|\epsilon_{\mu\mu}^{d}|\cos\phi_{\mu\mu}^{d} \\ P_{\nu_{\mu}^{s} \to \nu_{e}^{d}}^{ND} &\approx |\epsilon_{\mu e}^{s}|^{2} + |\epsilon_{\mu e}^{d}|^{2} + 2|\epsilon_{\mu e}^{s}||\epsilon_{\mu e}^{d}|\cos(\phi_{\mu e}^{s} - \phi_{\mu e}^{d}) \\ P_{\nu_{e}^{s} \to \nu_{\mu}^{d}}^{ND} &\approx |\epsilon_{e\mu}^{s}|^{2} + |\epsilon_{e\mu}^{d}|^{2} + 2|\epsilon_{e\mu}^{s}||\epsilon_{e\mu}^{d}|\cos(\phi_{e\mu}^{s} - \phi_{e\mu}^{d}) \end{split}$$

- Zero-distance effects show up for CC-NSIs at the near detector
- It is important for a near detector to discover New physics

#### Neutrino oscillation channels at a far detector



 $P_{\nu_{\mu} \to \nu_{e}}^{vac} = s_{2 \times 13}^{2} s_{23}^{2} \sin^{2} \Delta_{31}$  $+\alpha \Delta_{31} s_{2\times 12} s_{2\times 23} s_{13} [\sin(2\Delta_{31})\cos\delta - 2\sin\delta\sin^2\Delta_{31}]$  $+\alpha^2 \Delta_{21}^2 c_{22}^2 s_{2\times 12}^2$  $-\left|\epsilon_{\mu e}^{s}\right|s_{2\times 13}s_{23}\sin(\delta+\phi_{\mu e}^{s})\sin(2\Delta_{31})$  $-2|\epsilon_{\mu e}^{s}|s_{2\times 13}s_{23}\cos(\delta+\phi_{\mu e}^{s})\sin^{2}\Delta_{31}$  $-|\epsilon_{\mu e}^{d}|s_{2 \times 13}s_{23}\sin(\delta+\phi_{\mu e}^{d})\sin(2\Delta_{31})$  $-2|\epsilon_{\mu e}^{d}|s_{2\times 13}s_{23}c_{2\times 23}\cos(\delta+\phi_{\mu e}^{d})\sin^{2}\Delta_{31}$  $+4|\epsilon_{\tau e}^{d}|c_{23}s_{2\times 13}s_{23}^{2}\cos(\delta+\phi_{\tau e}^{d})\sin^{2}\Delta_{31}$  $-2\alpha\Delta_{31}|\epsilon_{\mu e}^{s}|c_{13}s_{2\times 12}c_{23}\sin\phi_{\mu e}^{s}$  $+2\alpha\Delta_{31}|\epsilon_{\mu e}^{d}|s_{2\times 12}c_{13}c_{23}s_{23}^{2}\cos\phi_{\mu e}^{d}\sin(2\Delta_{31})$  $-2\alpha\Delta_{31}|\epsilon^d_{\mu e}|s_{2\times 12}c_{13}c_{23}\sin\phi^d_{\mu e}(1-2s^2_{23}\sin^2\Delta_{31})$  $+4\alpha\Delta_{31}\epsilon_{\tau e}^{d}c_{13}c_{23}^{2}s_{2\times 12}s_{23}\sin\phi_{\tau e}^{d}\sin^{2}\Delta_{31}$  $+2\alpha\Delta_{31}|\epsilon_{\tau e}^{d}|c_{13}c_{23}^{2}s_{2\times 12}s_{23}\cos\phi_{\tau e}^{d}\sin(2\Delta_{31})$  $+O(\alpha^{3}) + O(\alpha^{2}s_{13}) + O(\alpha s_{13}^{2}) + O(s_{13}^{3})$  $+O(\epsilon\alpha^2)+O(\epsilon s_{13}^2)+O(\epsilon^2)$ 

 $P_{\nu_e \to \nu_{\mu}}^{vac} = s_{2 \times 13}^2 s_{23}^2 \sin^2 \Delta_{31}$  $+\alpha \Delta_{31} s_{2\times 12} s_{2\times 23} s_{13} [\sin(2\Delta_{31})\cos\delta + 2\sin\delta\sin^2\Delta_{31}]$  $+\alpha^2 \Delta_{31}^2 c_{23}^2 s_{2\times 12}^2$  $+ |\epsilon_{e\mu}^s| s_{2\times 13} s_{23} \sin(\delta - \phi_{e\mu}^s) \sin(2\Delta_{31})$  $-4|\epsilon_{e\mu}^s|s_{2\times13}s_{23}c_{2\times23}\cos(\delta-\phi_{e\mu}^s)\sin^2\Delta_{31}$  $+4|\epsilon_{e\tau}^{s}|c_{23}s_{2\times 13}s_{23}^{2}\cos(\delta-\phi_{e\tau}^{s})\sin^{2}\Delta_{31}$  $+ |\epsilon_{e\mu}^{d}| s_{2 \times 13} s_{23} \sin(\delta - \phi_{e\mu}^{d}) \sin(2\Delta_{31})$  $-2|\epsilon_{e\mu}^d|s_{2\times 13}s_{23}\cos(\delta-\phi_{e\mu}^d)\sin^2\Delta_{31}$  $+2\alpha\Delta_{31}|\epsilon_{e\mu}^{s}|s_{2\times12}c_{13}c_{23}s_{23}^{2}\cos\phi_{e\mu}^{s}\sin(2\Delta_{31})$  $-2\alpha\Delta_{31}|\epsilon_{e\mu}^{s}|s_{2\times12}c_{13}c_{23}\sin\phi_{e\mu}^{s}(1-2s_{23}^{2}\sin^{2}\Delta_{31})$  $+2\alpha\Delta_{31}|\epsilon_{e\tau}^{s}|c_{13}c_{23}^{2}s_{2\times 12}s_{23}\cos\phi_{e\tau}^{s}\sin(2\Delta_{31})$  $+4\alpha\Delta_{31}|\epsilon_{e\tau}^{s}|c_{13}c_{23}^{2}s_{2\times12}s_{23}\sin\phi_{e\tau}^{s}\sin^{2}\Delta_{31}$  $-2\alpha\Delta_{31}|\epsilon^d_{e\mu}|c_{13}s_{2\times12}c_{23}\sin(\phi^d_{e\mu})$  $+O(\alpha^{3}) + O(\alpha^{2}s_{13}) + O(\alpha s_{13}^{2}) + O(s_{13}^{3})$  $+O(\epsilon\alpha^2)+O(\epsilon s_{12}^2)+O(\epsilon^2)$ 

- Selections of proper CC-NSI parameters to be studied in simulation
- Understand correlations of CC-NSI and standard parameters

#### Neutrino oscillation channels at a far detector



 $P_{\nu_e \to \nu_{\mu}}^{vac} = s_{2 \times 13}^2 s_{23}^2 \sin^2 \Delta_{31}$  $P_{\nu_{\mu} \to \nu_{e}}^{vac} = s_{2 \times 13}^{2} s_{23}^{2} \sin^{2} \Delta_{31}$  $+\alpha \Delta_{31} s_{2\times 12} s_{2\times 23} s_{13} [\sin(2\Delta_{31})\cos\delta + 2\sin\delta\sin^2\Delta_{31}]$  $+\alpha \Delta_{31} s_{2\times 12} s_{2\times 23} s_{13} [\sin(2\Delta_{31})\cos\delta - 2\sin\delta\sin^2\Delta_{31}]$  $+\alpha^2 \Delta_{21}^2 c_{22}^2 s_{2\times 12}^2$  $+\alpha^2 \Delta_{21}^2 c_{22}^2 s_{2\times 12}^2$  $+\epsilon_{e\mu}^{s}s_{2\times13}s_{23}\sin(\delta-\phi_{e\mu}^{s})\sin(2\Delta_{31})$  $-\epsilon_{\mu e}^{s}s_{2\times 13}s_{23}\sin(\delta+\phi_{\mu e}^{s})\sin(2\Delta_{31})$ Correlations and adjust the location of maximum  $-2\epsilon_{\mu e}^{s}s_{2\times 13}s_{23}\cos(\delta+\phi_{\mu e}^{s})\sin^{2}\Delta_{31}$  $-4|\epsilon_{e\mu}^{s}|s_{2\times13}s_{23}c_{2\times23}\cos(\delta-\phi_{e\mu}^{s})\sin^{2}\Delta_{31}$ Vanishing for maximal  $-\epsilon^d_{\mu e} s_{2\times 13} s_{23} \sin(\delta + \phi^d_{\mu e}) \sin(2\Delta_{31})$  $+4(\epsilon_{e\tau}^{s})c_{23}s_{2\times 13}s_{23}^{2}\cos(\delta-\phi_{e\tau}^{s})\sin^{2}\Delta_{31}$ mu-tau symmetry  $-2|\epsilon_{\mu e}^{d}|s_{2\times 13}s_{23}c_{2\times 23}\cos(\delta+\phi_{\mu e}^{d})\sin^{2}\Delta_{31}$  $+ \epsilon^d_{e\mu} s_{2\times 13} s_{23} \sin(\delta - \phi^d_{e\mu}) \sin(2\Delta_{31})$  $-2\overline{\epsilon_{e\mu}^d}s_{2\times 13}s_{23}\cos(\delta-\phi_{e\mu}^d)\sin^2\Delta_{31}$  $+4(\epsilon_{\tau e}^{d})c_{23}s_{2\times 13}s_{23}^{2}\cos(\delta+\phi_{\tau e}^{d})\sin^{2}\Delta_{31}$ Correlations and change  $+2\alpha\Delta_{31}|\epsilon_{e\mu}^{s}|s_{2\times12}c_{13}c_{23}s_{23}^{2}\cos\phi_{e\mu}^{s}\sin(2\Delta_{31})$  $-2\alpha\Delta_{31}|\epsilon_{\mu e}^{s}|c_{13}s_{2\times 12}c_{23}\sin\phi_{\mu e}^{s}$ the amplitude of peak  $+2\alpha\Delta_{31}|\epsilon^{d}_{\mu e}|s_{2\times 12}c_{13}c_{23}s^{2}_{23}\cos\phi^{d}_{\mu e}\sin(2\Delta_{31})$  $-2\alpha\Delta_{31}|\epsilon_{e\mu}^{s}|s_{2\times12}c_{13}c_{23}\sin\phi_{e\mu}^{s}(1-2s_{23}^{2}\sin^{2}\Delta_{31})$  $-2\alpha\Delta_{31}|\epsilon^d_{\mu e}|s_{2\times 12}c_{13}c_{23}\sin\phi^d_{\mu e}(1-2s^2_{23}\sin^2\Delta_{31})$  $+2\alpha\Delta_{31}|\epsilon_{e\tau}^{s}|c_{13}c_{23}^{2}s_{2\times 12}s_{23}\cos\phi_{e\tau}^{s}\sin(2\Delta_{31})$  $+4\alpha\Delta_{31}|\epsilon_{e\tau}^{s}|c_{13}c_{23}^{2}s_{2\times12}s_{23}\sin\phi_{e\tau}^{s}\sin^{2}\Delta_{31}$  $+4\alpha\Delta_{31}\epsilon_{\tau e}^{d}c_{13}c_{23}^{2}s_{2\times 12}s_{23}\sin\phi_{\tau e}^{d}\sin^{2}\Delta_{31}$  $-2\alpha\Delta_{31}|\epsilon^d_{e\mu}|c_{13}s_{2\times12}c_{23}\sin(\phi^d_{e\mu})$  $+2\alpha\Delta_{31}\epsilon_{\pi e}^{d}c_{13}c_{23}^{2}s_{2\times 12}s_{23}\cos\phi_{\pi e}^{d}\sin(2\Delta_{31})$  $+O(\alpha^{3}) + O(\alpha^{2}s_{13}) + O(\alpha s_{13}^{2}) + O(s_{13}^{3})$  $+O(\alpha^{3}) + O(\alpha^{2}s_{13}) + O(\alpha s_{13}^{2}) + O(s_{13}^{3})$  $+O(\epsilon\alpha^2)+O(\epsilon s_{13}^2)+O(\epsilon^2)$  $+O(\epsilon\alpha^2)+O(\epsilon s_{12}^2)+O(\epsilon^2)$ 

- Selections of proper CC-NSI parameters to be studied in simulation
- Understand correlations of CC-NSI and standard parameters

#### Numerical tests of oscillation probabilities/events



**Jian Tang** 

**School of Physics** 

#### Table of Contents



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#### Impacts on precision measurements by CC-NSIs





**Degeneracy shows up after an introduction of CC-NSIs at some parameter space.** 

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#### Constraints of CC-NSIs with a near detector at MOMENT

SUN UNITE

- Colorful regions are allowed after running a near detector at MOMENT.
- Appearance channels can well constrain the magnitude of related NSI parameters while have no impacts on their phases.
- Disappearance channels can exclude a large region allowed in current bounds.
- A strong correlation between the magnitude and phase of NSIs.



#### Constraints of CC-NSIs with a far detector at MOMENT



- Colorful regions are allowed after running a far detector at MOMENT.
- The e-mu sector of NSI are the best constrained
- Almost all NSI-induced CP phases change the exclusion limits severely except the e-mu sector.
- Limits from other sectors are not as good as those from the e-mu sector of NSI.



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#### Summary



- We have studied CC-NSI effects with ND+FD at MOMENT.
- We have found some improvements of bounds on CC-NSIs at MOMENT.
- NSIs destroy the precision measurements of standard mixing parameters.
- Degeneracies between NSI and standard mixing parameters deserve further study.

Parameter	ND constraints	FD constraints	ND+FD constraints	Current bounds
$ \epsilon^s_{ee} $	0.027	0.028	0.018	0.025
$ \epsilon^s_{e\mu} $	0.023	0.018	0.014	0.03
$ \epsilon^s_{e\tau} $	n/a	0.065	0.065	0.03
$ \epsilon^s_{\mu e} $	0.025	0.021	0.015	0.025
$ \epsilon^s_{\mu\mu} $	0.028	0.029	0.019	0.03
$ \epsilon^s_{\mu au} $	n/a	0.054	0.054	0.03
$ \epsilon^d_{ee} $	0.027	0.028	0.027	0.041
$ \epsilon^d_{e\mu} $	0.023	0.015	0.013	0.026
$ \epsilon^d_{\mu e} $	0.025	0.022	0.025	0.025
$ \epsilon^d_{\mu\mu} $	0.028	0.03	0.028	0.078
$ \epsilon^d_{ au e} $	n/a	0.065	0.065	0.041
$ \epsilon^d_{ au\mu} $	n/a	0.054	0.054	0.013

#### Thank you for your attention!



#### • Near and far detectors are needed.

Fiducial mass (ND/FD)	Gd-doping Water cherenkov		
FIGUCIAI MASS (ND/FD)	(100 t/500 kton)		
Baseline $(ND/FD)$	$500 { m m}/150 { m km}$		
Channels	$\nu_e(\bar{\nu}_e) \rightarrow \nu_e(\bar{\nu}_e),  \nu_\mu(\bar{\nu}_\mu) \rightarrow \nu_\mu(\bar{\nu}_\mu)$		
Onamiers	$\nu_e(\bar{\nu}_e) \rightarrow \nu_\mu(\bar{\nu}_\mu),  \nu_\mu(\bar{\nu}_\mu) \rightarrow \nu_e(\bar{\nu}_e)$		
Energy resolution	$8.5\%/\sqrt{E}$		
Runtime	$\mu^-$ mode 5 yrs+ $\mu^+$ mode 5 yrs		
Energy range	$100~{\rm MeV}$ to $800~{\rm MeV}$		
Ffficioney	$\nu_{\mu} \ (\bar{\nu}_{\mu})$ seclection: 50%		
Enciency	$\nu_e \ (\bar{\nu}_e)$ selection: 40%		
Normalization	appearance channels: $2.5\%$		
error on signal	disappearance channels: $5\%$		
Normalization	5% (all channels)		
error on backgroud			
	Neutral current		
Backgound sources	charge misidentifications		
	atmospheric neutrinos		

#### Backup: Oscillation channels faked by CC-NSIs



$$\mu^+ \to \nu_e \xrightarrow{\text{Osc.}} \nu_\mu \to \mu^-$$
 (1)

$$\mu^+ \to \nu_e \xrightarrow{\text{Osc.}} \nu_e \to e^- \tag{2}$$

$$\mu^+ \to \bar{\nu}_\mu \xrightarrow{\text{Osc.}} \bar{\nu}_\mu \to \mu^+$$
(3)

$$\mu^+ \to \bar{\nu}_\mu \xrightarrow{\text{Osc.}} \bar{\nu}_e \to e^+$$
 (4)

With CC-NSIs at the source, things are changed:

$$\mu^+ \xrightarrow{\text{NSI}} \nu_\mu \xrightarrow{\text{No osc.}} \nu_\mu \to \mu^-,$$
 (5)

$$\mu^+ \xrightarrow{\text{NSI}} \bar{\nu}_e \xrightarrow{\text{No osc.}} \bar{\nu}_e \to e^+,$$
(6)

Similarly, we obtain the processes including CC-NSIs at the detector:

$$\mu^+ \to \nu_e \xrightarrow{\text{No osc.}} \nu_e \xrightarrow{\text{NSI}} \mu^-.$$
 (7)

$$\mu^+ \to \bar{\nu}_\mu \xrightarrow{\text{No osc.}} \bar{\nu}_\mu \xrightarrow{\text{NSI}} e^+.$$
 (8)

#### Chirality discussions about NSI effective Lagrangian



$$\begin{split} \mathcal{L}_{\text{CC-NSI}}^{s,d} &= \frac{G_F}{\sqrt{2}} \sum_{f,f'} \varepsilon_{\alpha\beta}^{CC} [\bar{\nu}_{\alpha} \gamma^{\mu} (1 \mp \gamma^5) \ell_{\beta}] [\bar{f} \gamma_{\mu} (1 \mp \gamma^5) f'] + h.c. \begin{bmatrix} f \\ f \\ \bar{\nu}_{\alpha} \\ \bar{\nu}_{\alpha\beta} \end{bmatrix} \\ \mathcal{L}_{\text{CC-NSI}} &= \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{\mu\mp} \right)^s \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) \nu_{\mu} \right\} \left\{ \bar{\mu} \gamma_{\rho} (1 \mp \gamma^5) e \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^s \left\{ \bar{\nu}_{e} \gamma^{\rho} (1 - \gamma^5) \nu_{e} \right\} \left\{ \bar{\mu} \gamma_{\rho} (1 \mp \gamma^5) e \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{\mu\mp} \right)^d \left\{ \bar{\mu} \gamma^{\rho} (1 - \gamma^5) \nu_{e} \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left( \epsilon_{e\mu}^{e\mp} \right)^d \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{u} \gamma_{\rho} (1 \mp \gamma^5) d \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left\{ \epsilon_{e\mu}^{e\mp} \right\} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{\nu}_{\mu} \gamma_{\mu} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \\ &+ \frac{G_F}{\sqrt{2}} \left\{ \bar{\nu}_{\mu} \gamma^{\rho} (1 - \gamma^5) e \right\} \\ &+ \frac{G_F}{\sqrt{2}}$$

	source		detector	
	superbeam	muon-decay beam	superbeam	muon-decay beam
f	d	$\mu/e$	d	d
f'	u	$ u_{\mu}/ u_{e}$	u	u
$\nu_{lpha}$	$ u_e,  u_\mu,  u_ au$	$ u_e,  u_\mu,  u_ au$	$ u_e, \nu_\mu, \nu_ au$	$ u_e,  u_\mu,  u_ au$
$\ell_{eta}$	$e/\mu$	$e/\mu$	$e,\mu$	e, $\mu$

(V-A)(V+A) NSIs suppressed by

the helicity factors  $\mathcal{O}\left(\frac{m_{\mu}m_{e}}{E_{\mu}E_{e}}\right)$  and  $\mathcal{O}\left(\frac{m_{u}m_{d}}{E_{u}E_{d}}\right)$