

Particle physics – where to next ?



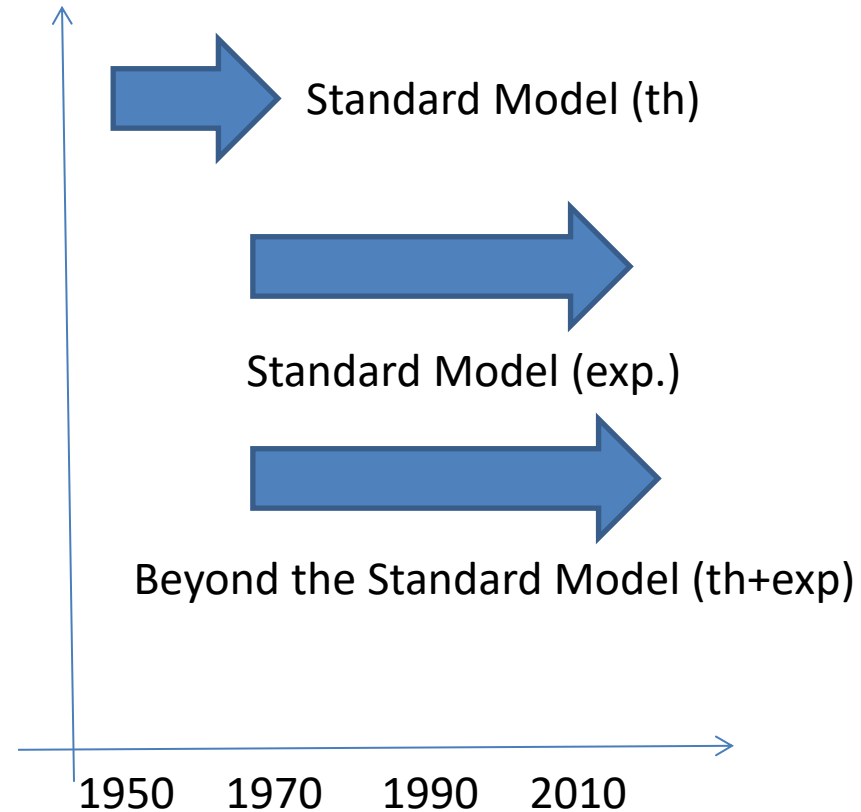
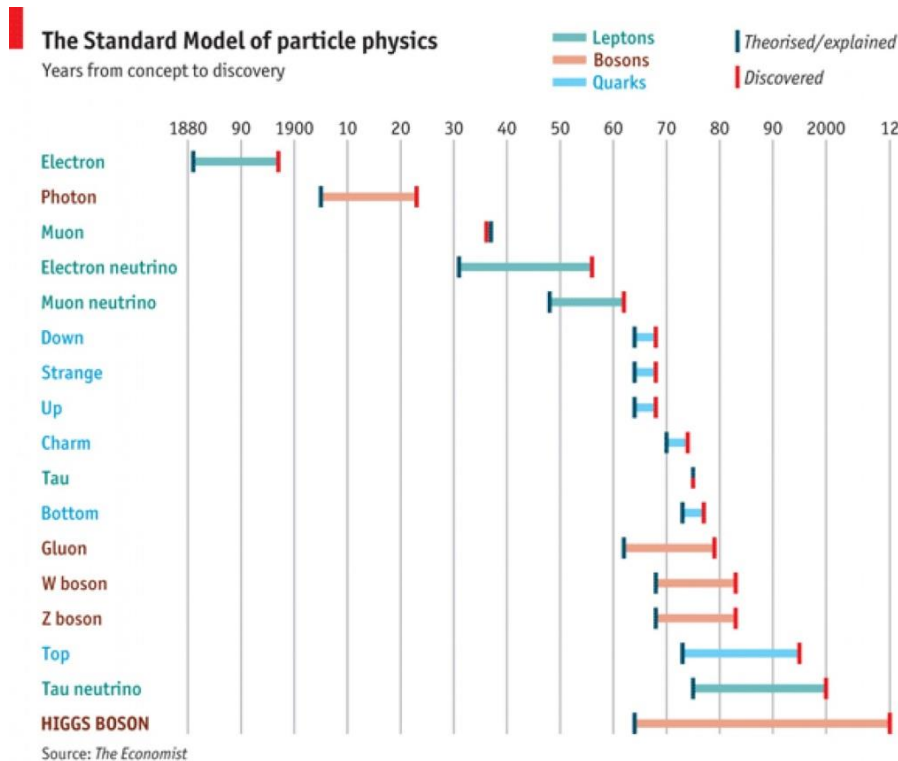
David Milstead

- Where are we ?
- What are we doing now ?
- What can/should we do in the future?
- The European PP Strategy Update

Disclaimer

- *PP is a very broad field.*
- *Focus on the issues surrounding EWSB and new physics*
- *Dark matter plays an important role in collider physics but is not discussed (much) here.*
- *Neutrinos are also largely ignored.*

A potted history of particle physics



Guiding paradigms

(1) SM development (~ 1950-1970)

Theory: gauge symmetries, renormalizability, EW symmetry breaking.

Experiment: Exploratory and guided by theory.

(2) SM validation (~ 1970->2035 ?)

Gradual process as increasing energy reveals more massive particles.

Experiment: Mostly guided by theory.

(3) Beyond-the-Standard Model searches (~ 1970->)

Theory: GUTs, naturalness, supersymmetry etc.

Experiment: Mostly ?? guided by theory (naturalness).

(2) is alive (EWSB+top poorly measured) but the end is coming.

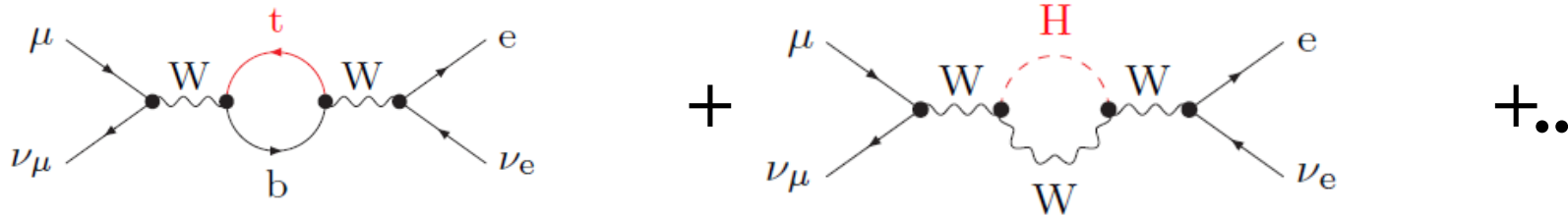
(3) is the present + future. What should our strategy be ?

Collider+non-collider + theory=beautiful SM physics

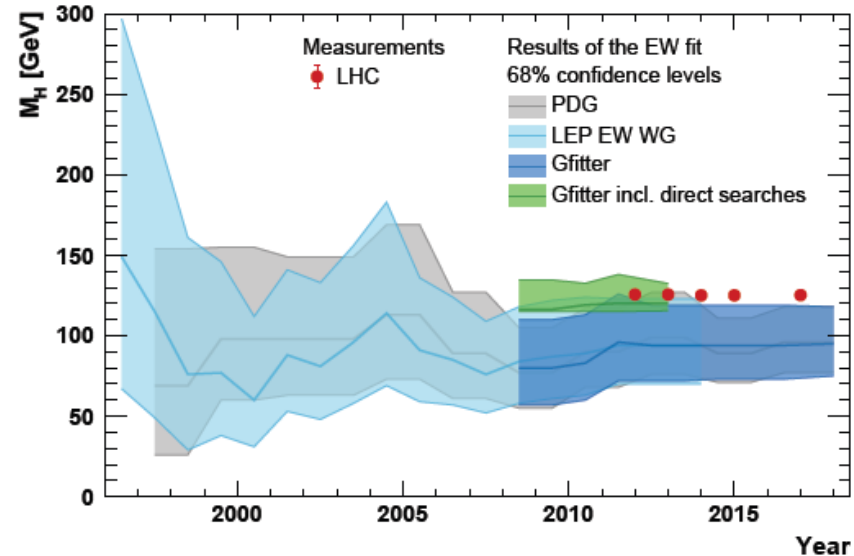
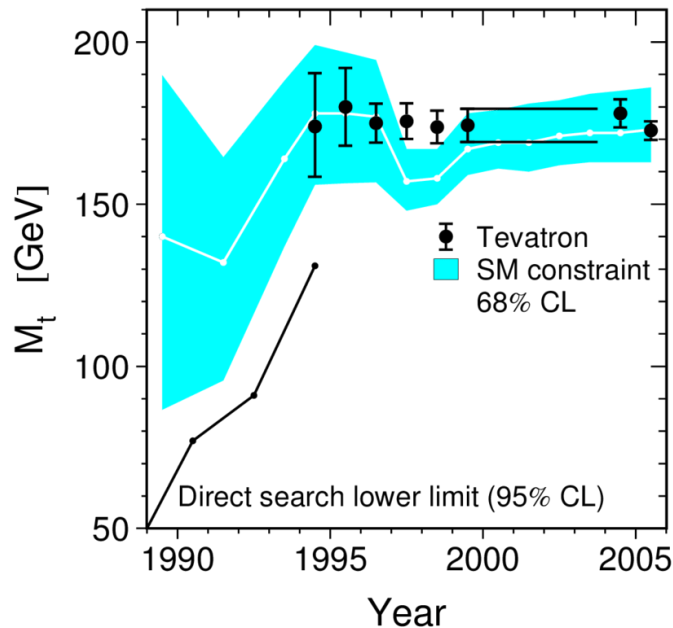
Precision muon lifetime : $\tau_\mu = (2.1969811 \pm 0.0000022) \times 10^{-6} \text{ s}$

+ collider measurements (eg $M_W = 80.385 \pm 0.015 \text{ GeV}$)

\Rightarrow unobserved particle masses.

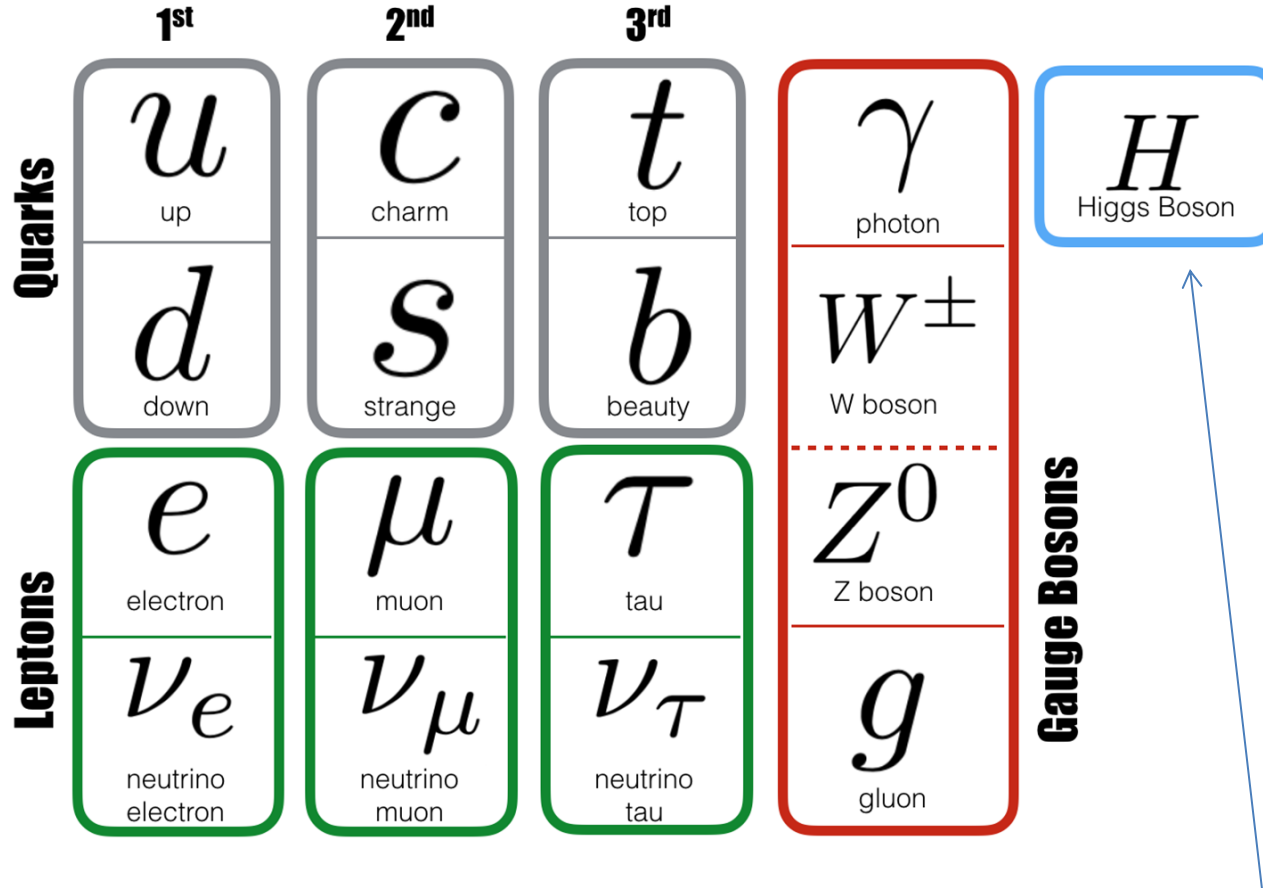


\Rightarrow Predict top quark and Higgs boson masses ahead of discovery.



"No lose" experiments - SM particle discovery or new physics.

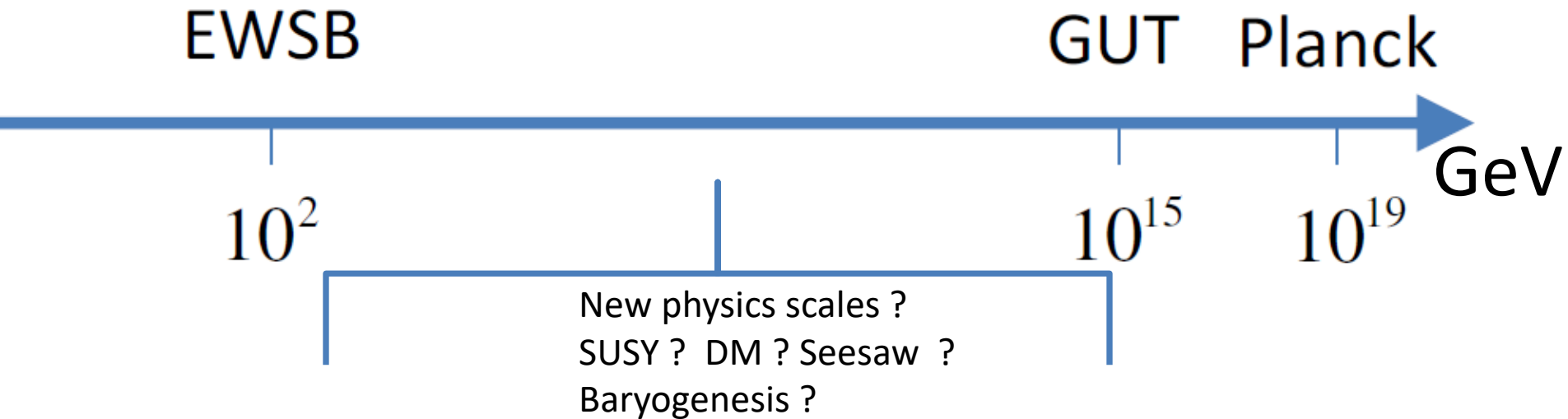
The Standard Model



Fundamental scalar !
We've never seen one of those before
And an entirely new field!

“Priority” issues for PP

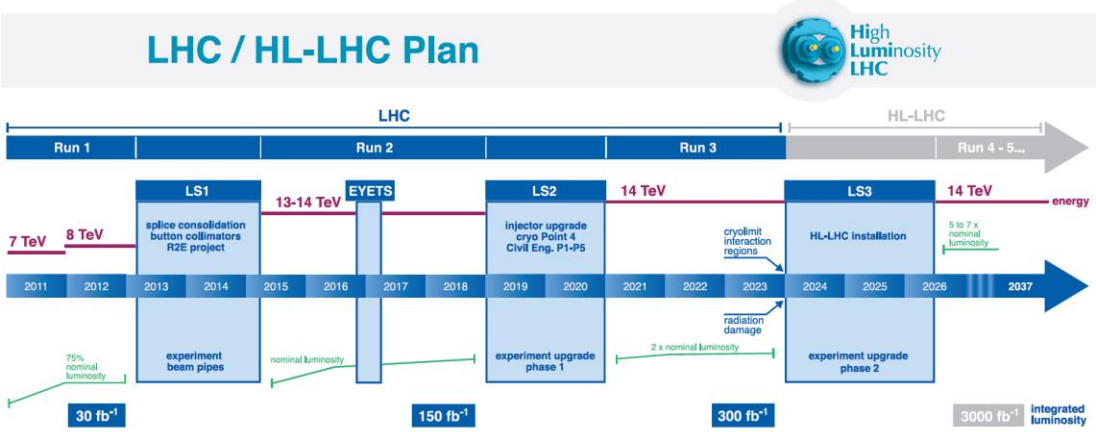
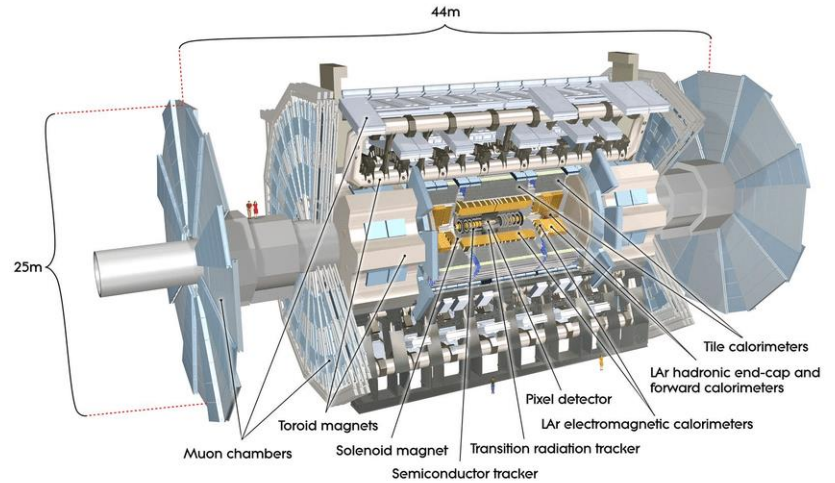
- SM
 - EWSB is poorly explored experimentally.
- Where is the new physics energy scale ?
 - Dark matter, naturalness, SUSY -> TeV scale ?



- Explore the energy and precision frontiers

The energy frontier

Main collider work horse – Large Hadron Collider

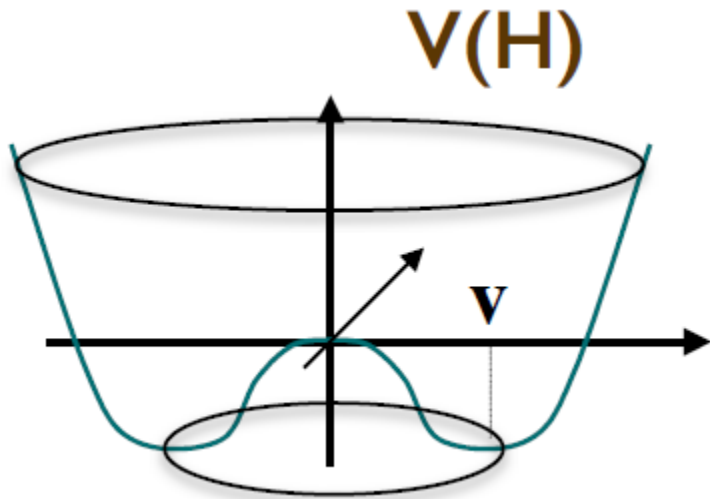


Luminosity and detector upgrades.

Taken ~% of total luminosity

Built to find the Higgs

EWSB and the Higgs



Couplings to massive particles

A Feynman diagram showing a Higgs boson (H) decaying into two fermions (f). The Higgs is represented by a vertical dashed line, and the fermions by solid lines. A black dot at the vertex indicates the interaction.

$$g_{Hf\bar{f}} = \frac{m_f}{v}$$

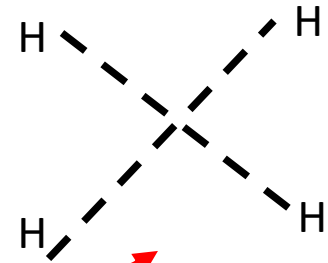
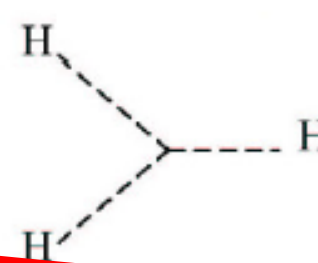
A Feynman diagram showing a Higgs boson (H) decaying into two vector bosons (V). The Higgs is represented by a vertical dashed line, and the vector bosons by wavy lines. A black dot at the vertex indicates the interaction.

$$g_{HVV} = \frac{2m_V^2}{v}$$

$$V(h) = \frac{M_H^2}{2} h^2 + 4\lambda v h^3 + \lambda h^4$$

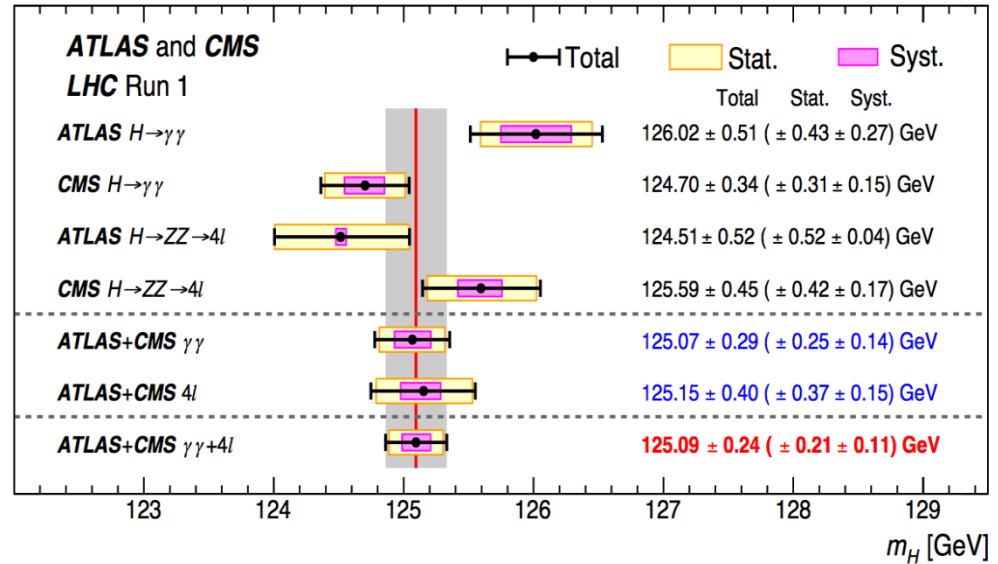
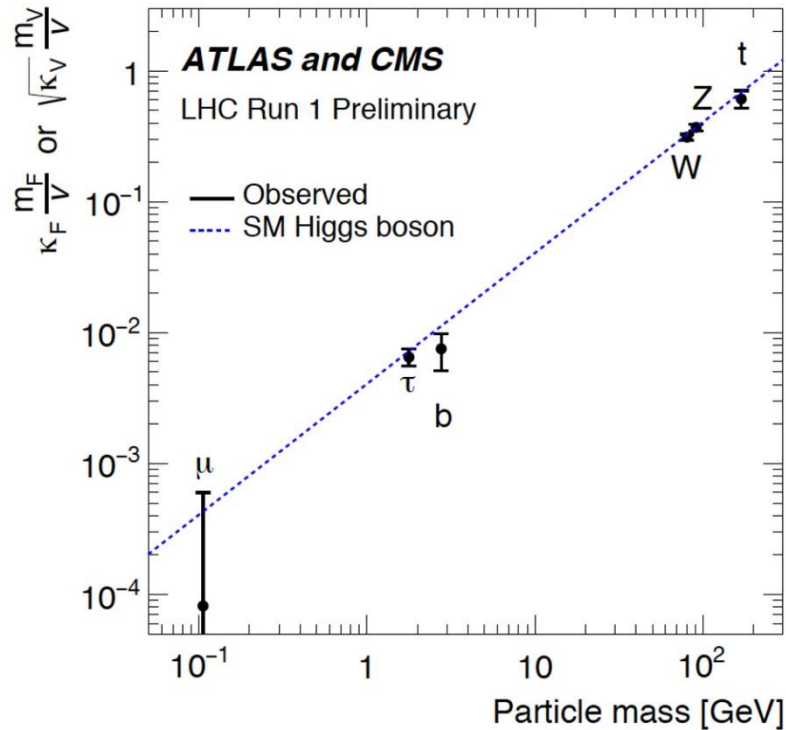
$$v^2 = -\frac{\mu^2}{2\lambda} = (\sqrt{2}G_F)^2 = (246 \text{ GeV})^2$$

$$M_H = 126 \text{ GeV} \Rightarrow \lambda = 0.13$$



Self-couplings

The LHC Higgs



Looks and smells like a SM Higgs with ~ 125 GeV mass
 Imprecise measurements.
 Model-dependent (no absolute normalisation)
 Missing couplings (eg self-interactions).

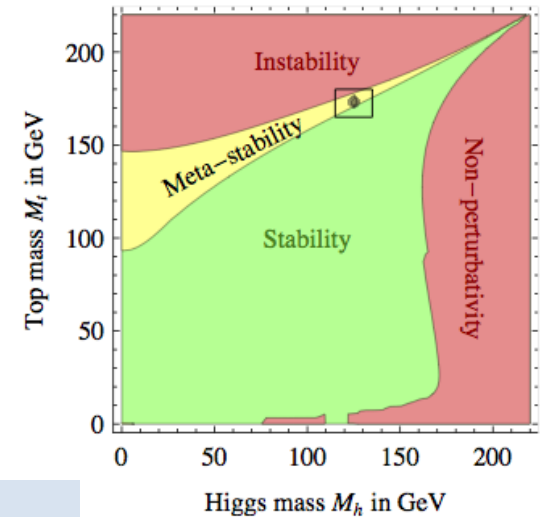
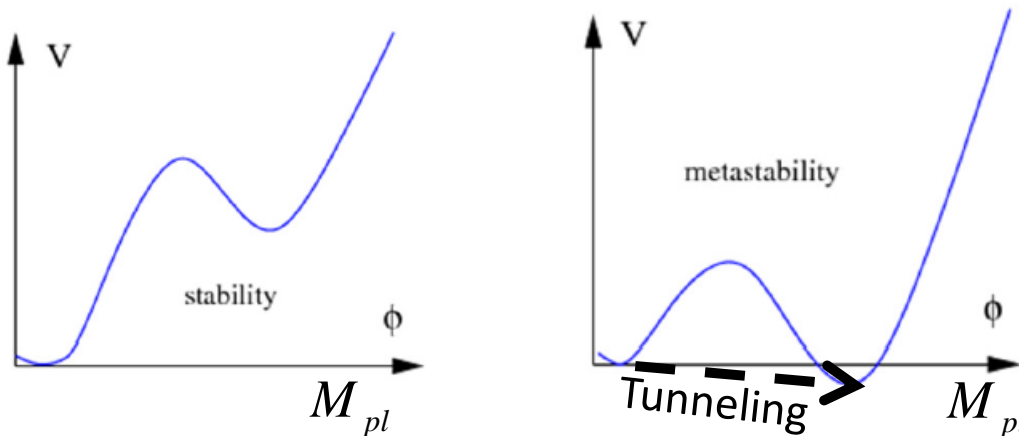
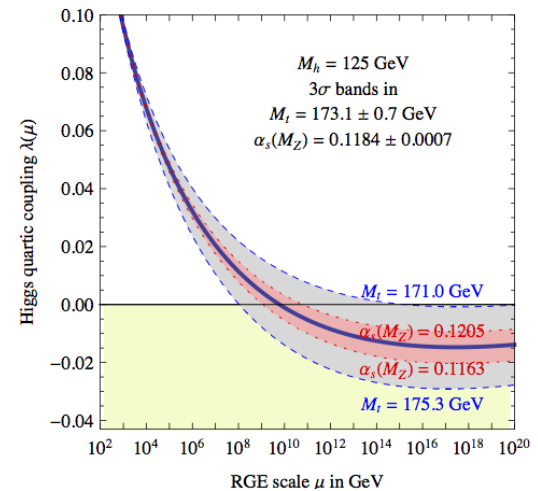
Vacuum (meta)-stability

$$V(h) = \frac{M_H^2}{2} h^2 + 4\lambda v h^3 + \lambda h^4$$

⇒ Large values of the Higgs field: $V(h) \sim \lambda(h) h^4$

Renormalisation: $\lambda(h)$ runs due to top mass.

$\lambda(h)$ can turn negative



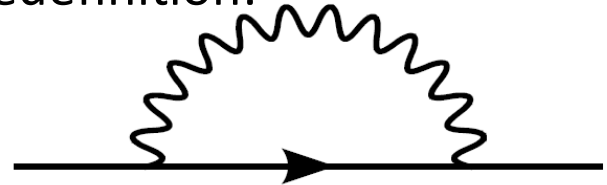
$$P_{\text{tunnel}} \propto e^{-\frac{1}{\lambda}} \Rightarrow \text{no need to panic } \tau_{\text{vacuum}} > \tau_{\text{universe}}$$

Interesting that the SM Higgs sits on edge of stability.

Naturalness and Higgs fine-tuning

Renormalize loop divergences away via mass redefinition:

$$m_{phys} = m_0 + \delta m_e \quad \left[\delta m_e = \int_0^\Lambda f(k) dk \right]$$



\equiv "new theory of short distance physics" + SM "long distance physics"

Λ = arbitrary cut-off scale for SM integral.

Electron: $m_e \sim m_{e0} + \delta m_e, \quad \delta m_e \sim \frac{2\alpha}{\pi} \ln \frac{\Lambda_{pl}}{m_e} \sim 0.25 m_e$

Higgs: $m_H^2 \sim m_{0H}^2 + \delta m_H^2$

$$(100)^2 \sim m_0^2 + \mathcal{O}(1) \Lambda_{pl}^2 \sim (10^{19})^2 - (10^{19})^2$$

δm_H (long-distance) and m_{0H} (short-distance) should be decoupled !

Such fine tuning is implausible/unlikely !

The SM Higgs is unnatural!

\Rightarrow New physics (SUSY, ED) at $\Lambda \sim \text{TeV}$ to suppress fine-tuning.

The fine tuning problem



Canada
9,984,670 km²

—



United States
9,826,675 km²

= 1 Å²

—

= 157,995 km²

tuning of the Higgs mass would correspond to the surface area of Canada and the United States differing by approximately the size of an atom!

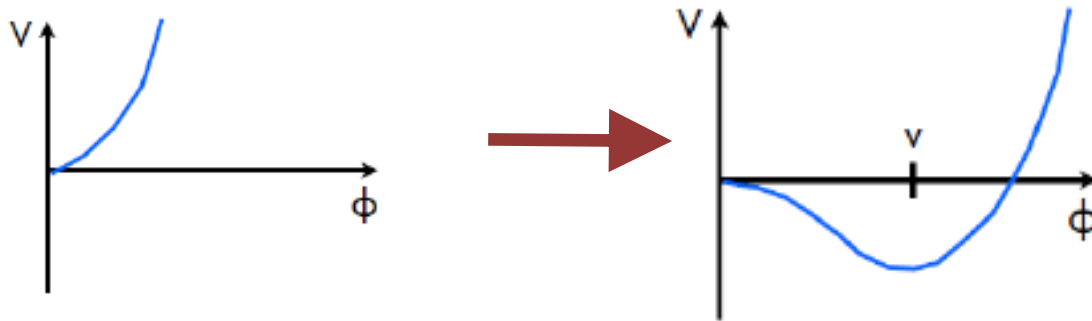
In order to **protect the Higgs mass** from huge quantum corrections and to avoid finetuning, we expect **New Physics at or below the TeV scale** not far above the mass of the Higgs

From unbroken to broken symmetry

No broken symmetry in the early universe

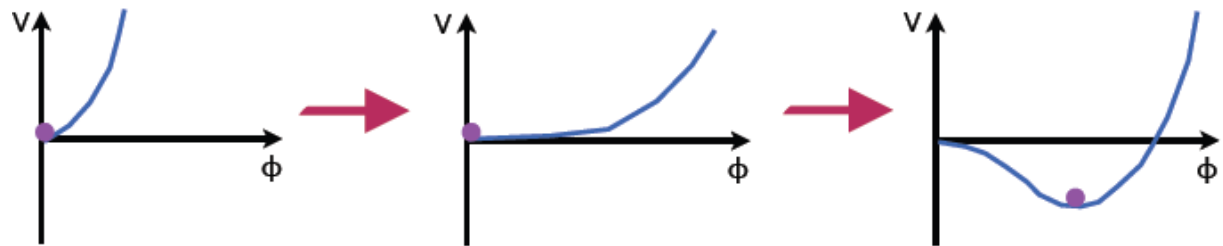
$$T \gg 100 \text{ GeV}$$

$$T = 0$$

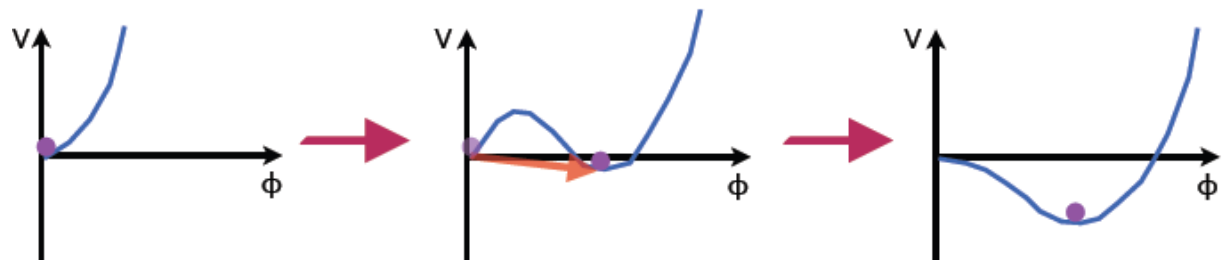


Phase transition

Second order



First order



Why is the phase transition important ?

SM has ingredients for baryogenesis :

baryon number and CP violation

Electroweak baryogenesis if first order phase transition.

⇒ Higgs mass < 70 GeV or new physics (eg SUSY).

⇒ Observable via anomalous cubic and quartic self-couplings.

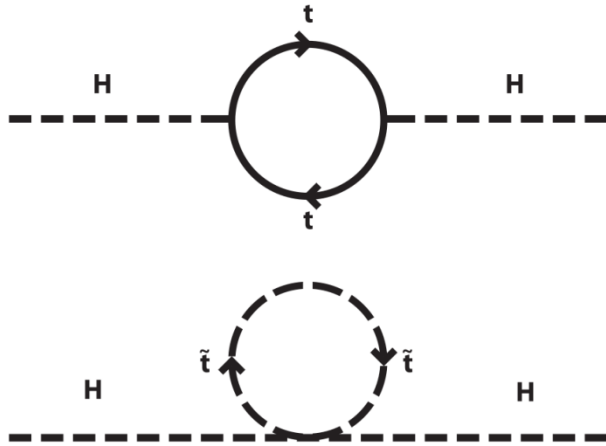
⇒ Use the Higgs to explain baryogenesis !

EWSB is the key issue for the energy frontier

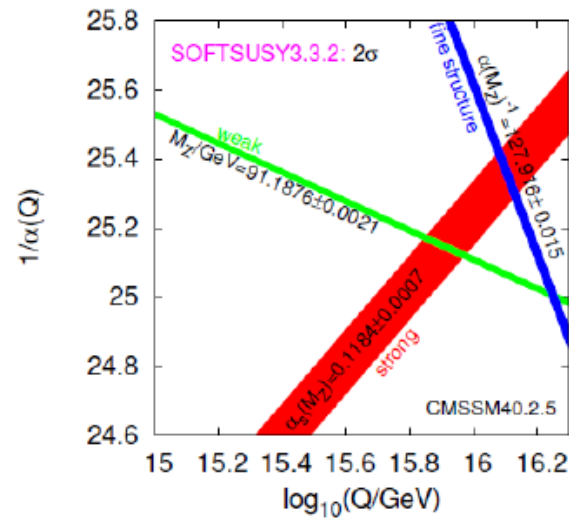
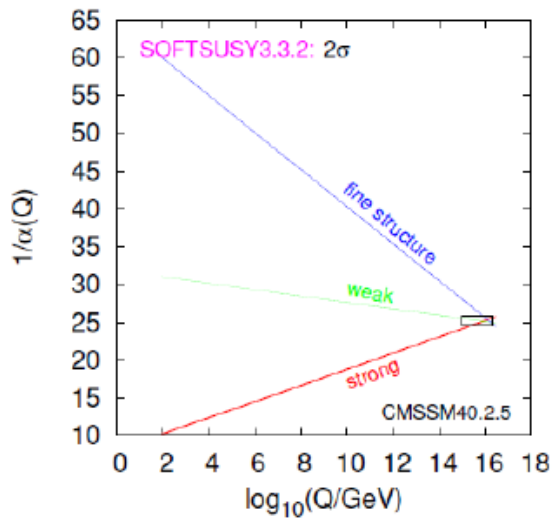
- Experiment
 - The only (apparently) fundamental scalar ever observed.
 - Properties must be measured to high precision (mass, couplings to SM inc. self-couplings)
- Theory
 - Naturalness
 - Vacuum stability
 - Electroweak phase transition (baryogenesis)

Major open questions about the Higgs can be posed.
Do we know the answers ?

Supersymmetry to the rescue ?



Stop and top cancel to suppress fine-tuning for SUSY \sim TeV mass.



+ Approximate unification of gauge couplings

+ First order phase transition

+ WIMP dark matter candidate

ATLAS SUSY Searches* - 95% CL Lower Limits

December 2017

ATLAS Preliminary

$\sqrt{s} = 7, 8, 13$ TeV

Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} d\tau [\text{fb}^{-1}]$	Mass limit	$\sqrt{s} = 7, 8$ TeV	$\sqrt{s} = 13$ TeV	Reference	
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{q}	1.57 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(1^{\text{st}} \text{ gen. } \tilde{q}) = m(2^{\text{nd}} \text{ gen. } \tilde{q})$	1712.02332
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ (compressed)	mono-jet	1-3 jets	Yes	36.1	\tilde{q}	710 GeV	$m(\tilde{q}) - m(\tilde{\chi}_1^0) < 5$ GeV	1711.03301
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.02 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	1712.02332
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0 \rightarrow q\tilde{q}W^{\pm}\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.01 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(\tilde{K}^{\pm}) = 0.5(m(\tilde{\chi}_1^0) + m(\tilde{g}))$	1712.02332
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell)\tilde{\chi}_1^0$	$ee, \mu\mu$	2 jets	Yes	14.7	\tilde{g}	1.7 TeV	$m(\tilde{\chi}_1^0) < 300$ GeV,	1611.05791
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell/\nu\nu)\tilde{\chi}_1^0$	$3e, \mu$	4 jets	-	36.1	\tilde{g}	1.87 TeV	$m(\tilde{\chi}_1^0) = 0$ GeV,	1706.03731
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}WZ\tilde{\chi}_1^0$	0	7-11 jets	Yes	36.1	\tilde{g}	1.8 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	1708.02794
	GMSB ($\tilde{\ell}$ NLSP)	$1-2\tau + 0-1\ell$	0-2 jets	Yes	3.2	\tilde{g}	2.0 TeV	$m(\tilde{\chi}_1^0) < 1700$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu > 0$	1607.05979
	GGM (bino NLSP)	2γ	-	Yes	36.1	\tilde{g}	2.15 TeV	$c\tau(\text{NLSP}) < 0.1$ mm	ATLAS-CONF-2017-080
	GGM (higgsino-bino NLSP)	γ	2 jets	Yes	36.1	\tilde{g}	2.05 TeV	$m(\tilde{\chi}_1^0) = 1700$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu > 0$	ATLAS-CONF-2017-080
Gravitino LSP	0	mono-jet	Yes	20.3	$R^{1/2}$ scale	865 GeV	$m(\tilde{G}) > 1.8 \times 10^{-4}$ eV, $m(\tilde{g}) = m(\tilde{q}) = 1.5$ TeV	1502.01518	
3 rd gen. \tilde{g} medi.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 b	Yes	36.1	\tilde{g}	1.92 TeV	$m(\tilde{\chi}_1^0) < 600$ GeV	1711.01901
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	$0-1 e, \mu$	3 b	Yes	36.1	\tilde{g}	1.97 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	1711.01901
3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 b	Yes	36.1	\tilde{b}_1	950 GeV	$m(\tilde{\chi}_1^0) < 420$ GeV	1708.09266
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^0$	$2e, \mu$ (SS)	1 b	Yes	36.1	\tilde{b}_1	275-700 GeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(\tilde{K}^{\pm}) = m(\tilde{\chi}_1^0) + 100$ GeV	1706.03731
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\chi}_1^0$	$0-2e, \mu$	1-2 b	Yes	4.7/13.3	\tilde{t}_1	117-170 GeV	$m(\tilde{\chi}_1^0) = 2m(\tilde{K}^{\pm}), m(\tilde{K}^{\pm}) = 55$ GeV	1209.2102, ATLAS-CONF-2016-077
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $t\tilde{\chi}_1^0$	$0-2e, \mu$	0-2 jets/1-2 b	Yes	20.3/36.1	\tilde{t}_1	90-198 GeV	$m(\tilde{\chi}_1^0) = 1$ GeV	1506.08616, 1709.04183, 1711.11520
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet	Yes	36.1	\tilde{t}_1	90-430 GeV	$m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 5$ GeV	1711.03301
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	$2e, \mu$ (Z)	1 b	Yes	20.3	\tilde{t}_1	150-600 GeV	$m(\tilde{\chi}_1^0) > 150$ GeV	1403.5222
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	$3e, \mu$ (Z)	1 b	Yes	36.1	\tilde{t}_2	290-790 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	1706.03986
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$	$1-2e, \mu$	4 b	Yes	36.1	\tilde{t}_2	320-880 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	1706.03986	
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	$2e, \mu$	0	Yes	36.1	$\tilde{\ell}$	90-500 GeV	$m(\tilde{\chi}_1^0) = 0$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}, \tilde{\chi}_1^{\pm} \rightarrow \tilde{\ell}\nu(\ell\nu)$	$2e, \mu$	0	Yes	36.1	$\tilde{\chi}_1^{\pm}$	750 GeV	$m(\tilde{\chi}_1^0) = 0, m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}/\tilde{\chi}_2^0, \tilde{\chi}_1^{\pm} \rightarrow \tilde{\tau}\nu(\tau\nu), \tilde{\chi}_2^0 \rightarrow \tilde{\tau}\tau(\nu\nu)$	2τ	-	Yes	36.1	$\tilde{\chi}_1^{\pm}$	760 GeV	$m(\tilde{\chi}_1^0) = 0, m(\tilde{\tau}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_1^0))$	1708.07875
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_1\nu\tilde{\ell}_1\ell(\nu\nu), \tilde{\ell}\tilde{\nu}\tilde{\ell}_1\ell(\nu\nu)$	$3e, \mu$	0	Yes	36.1	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_2^0$	1.13 TeV	$m(\tilde{\chi}_1^0) = m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0) = 0, m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_2^0))$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^{\pm}Z\tilde{\chi}_1^0$	$2-3e, \mu$	0-2 jets	Yes	36.1	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_2^0$	580 GeV	$m(\tilde{\chi}_1^0) = m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0) = 0, \tilde{\ell}$ decoupled	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^{\pm}h\tilde{\chi}_1^0, h \rightarrow b\tilde{b}/WW/\tau\tau/\gamma\gamma$	e, μ, γ	0-2 b	Yes	20.3	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_2^0$	270 GeV	$m(\tilde{\chi}_1^0) = m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0) = 0, \tilde{\ell}$ decoupled	1501.07110
	$\tilde{\chi}_2^0\tilde{\chi}_3^0, \tilde{\chi}_2^0 \rightarrow \tilde{\ell}_R\tilde{\ell}$	$4e, \mu$	0	Yes	20.3	$\tilde{\chi}_2^0, \tilde{\chi}_3^0$	635 GeV	$m(\tilde{\chi}_2^0) = m(\tilde{\chi}_3^0), m(\tilde{\chi}_1^0) = 0, m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_2^0) + m(\tilde{\chi}_1^0))$	1405.5086
GGM (wino NLSP) weak prod., $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$	$1e, \mu + \gamma$	-	Yes	20.3	\tilde{W}	115-370 GeV	$c\tau < 1$ mm	1507.05493	
GGM (bino NLSP) weak prod., $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$	2γ	-	Yes	36.1	\tilde{W}	1.06 TeV	$c\tau < 1$ mm	ATLAS-CONF-2017-080	
Long-lived particles	Direct $\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}$ prod., long-lived $\tilde{\chi}_1^{\pm}$	Disapp. trk	1 jet	Yes	36.1	$\tilde{\chi}_1^{\pm}$	460 GeV	$m(\tilde{\chi}_1^{\pm}) - m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^{\pm}) = 0.2$ ns	1712.02118
	Direct $\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}$ prod., long-lived $\tilde{\chi}_1^{\pm}$	dE/dx trk	-	Yes	18.4	$\tilde{\chi}_1^{\pm}$	495 GeV	$m(\tilde{\chi}_1^{\pm}) - m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^{\pm}) < 15$ ns	1506.05332
	Stable, stopped \tilde{g} R-hadron	0	1-5 jets	Yes	27.9	\tilde{g}	850 GeV	$m(\tilde{\chi}_1^0) = 100$ GeV, $10 \mu\text{s} < c\tau(\tilde{g}) < 1000$ s	1310.6584
	Stable \tilde{g} R-hadron	trk	-	-	3.2	\tilde{g}	1.58 TeV	-	1606.05129
	Metastable \tilde{g} R-hadron	dE/dx trk	-	-	3.2	\tilde{g}	1.57 TeV	$m(\tilde{\chi}_1^0) = 100$ GeV, $\tau > 10$ ns	1604.04520
	Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	displ. vtx	-	Yes	32.8	\tilde{g}	2.37 TeV	$\tau(\tilde{g}) = 0.17$ ns, $m(\tilde{\chi}_1^0) = 100$ GeV	1710.04901
	GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{\ell}, \tilde{\mu}) + \tau(e, \mu)$	$1-2\mu$	-	-	19.1	$\tilde{\chi}_1^0$	537 GeV	$10 < \tan\beta < 50$	1411.6795
	GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$, long-lived $\tilde{\chi}_1^0$	2γ	-	Yes	20.3	$\tilde{\chi}_1^0$	440 GeV	$1 < \tau(\tilde{\chi}_1^0) < 3$ ns, SPS8 model	1409.5542
	$\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow ee/\mu\nu/\mu\mu\nu$	displ. $ee/e\mu/\mu\mu$	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$7 < c\tau(\tilde{\chi}_1^0) < 740$ mm, $m(\tilde{g}) = 1.3$ TeV	1504.05162
	RPV	LFV $pp \rightarrow \tilde{\nu}_e + X, \tilde{\nu}_e \rightarrow e\mu/\tau\mu$	$e\mu, e\tau, \mu\tau$	-	-	3.2	$\tilde{\nu}_e$	1.9 TeV	$\lambda'_{111} = 0.11, \lambda'_{122}/\lambda'_{133}/\lambda'_{233} = 0.07$
Bilinear RPV CMSSM		$2e, \mu$ (SS)	0-3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.45 TeV	$m(\tilde{q}) = m(\tilde{g}), c\tau_{LSP} < 1$ mm	1404.2500
$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}, \tilde{\chi}_1^{\pm} \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow ee\nu, e\mu\nu, \mu\nu$		$4e, \mu$	-	Yes	13.3	$\tilde{\chi}_1^{\pm}$	1.14 TeV	$m(\tilde{\chi}_1^0) > 400$ GeV, $\lambda'_{12k} \neq 0$ ($k = 1, 2$)	ATLAS-CONF-2016-075
$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}, \tilde{\chi}_1^{\pm} \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tau\nu_e, e\tau\nu_e$		$3e, \mu + \tau$	-	Yes	20.3	$\tilde{\chi}_1^{\pm}$	450 GeV	$m(\tilde{\chi}_1^0) > 0.2x m(\tilde{\chi}_1^{\pm}), \lambda'_{133} \neq 0$	1405.5086
$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow q\tilde{q}$		0	4-5 large-R jets	-	36.1	\tilde{g}	1.875 TeV	$m(\tilde{\chi}_1^0) = 1075$ GeV	SUSY-2016-22
$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow q\tilde{q}$		$1e, \mu$	8-10 jets/0-4 b	-	36.1	\tilde{g}	2.1 TeV	$m(\tilde{\chi}_1^0) = 1$ TeV, $\lambda'_{112} \neq 0$	1704.08493
$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow b\tilde{s}$		$1e, \mu$	8-10 jets/0-4 b	-	36.1	\tilde{g}	1.65 TeV	$m(\tilde{t}_1) = 1$ TeV, $\lambda'_{323} \neq 0$	1704.08493
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{s}$		0	2 jets + 2 b	-	36.7	\tilde{t}_1	100-470 GeV	$BR(\tilde{t}_1 \rightarrow b\ell/\mu) > 20\%$	1710.07171
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\ell}$		$2e, \mu$	2 b	-	36.1	\tilde{t}_1	480-610 GeV	-	1710.05544
Other		Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$	0	2 c	Yes	20.3	\tilde{c}	510 GeV	$m(\tilde{\chi}_1^0) < 200$ GeV

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹ 1 Mass scale [TeV]

No new physics (so far).
Exclusions up to ~2 TeV.

The naturalness score card

Observable	Theory	Successful prediction of new physics
Kaon mass difference	Effective	Yes – charm mass
Pion mass difference	Effective	Postdiction - ρ^0
Electron mass	Classical-quantum	Postdiction – e^+
Higgs mass	SM	No (SUSY, ED)
Strong CP	SM	No (Axions)
Cosmological constant	SM/GR	No

A mixed set of results for the main motivation for new TeV-scale physics

Is new physics "around the corner" or hidden/invisible at the LHC ?

Is the basic reasoning correct ?

Quantum gravity may allow short/long distance mixing.

Baryogenesis

Electroweak baryogenesis is not the only model on the market....

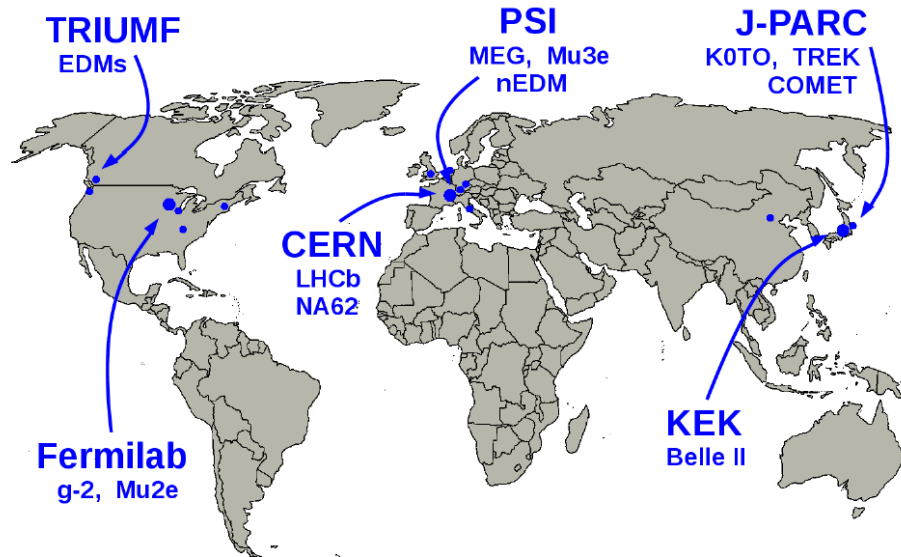
It is, however, broadly falsifiable.

Many models predict high mass states and can't be tested.

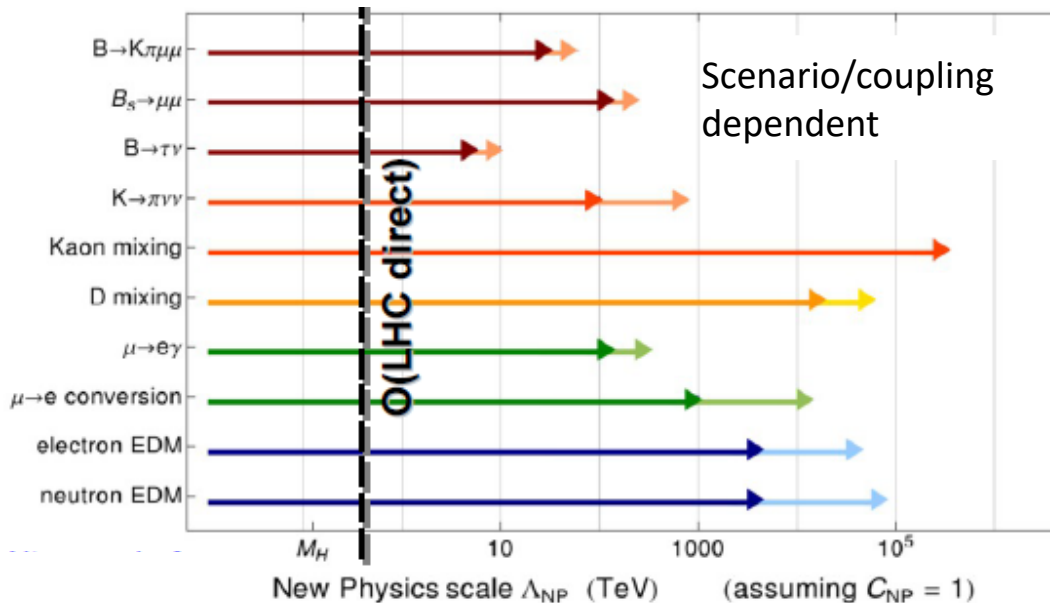
- Affleck-Dine baryogenesis
- GUT baryogenesis
- Leptogenesis
- Spontaneous baryogenesis
- Electroweak baryogenesis
- Dissipative baryogenesis
- Warm baryogenesis
- Cloistered baryogenesis
- Cold baryogenesis
- Planck baryogenesis
- WIMPy baryogenesis
- Dirac Leptogenesis
- Post-sphaleron baryogenesis
- Non-local electroweak baryogenesis
- Magnet assisted baryogenesis
- Singlet-assisted baryogenesis

The precision frontier

Precision frontier experiments



Broad and diverse program and international and national labs and institutions.



NP scale sensitivity beyond the LHC for range of couplings and new particles.

Electric dipole moment= new physics

Electric dipole moment along spin axis: $\bar{d} = d\hat{S}$

T -transformation:

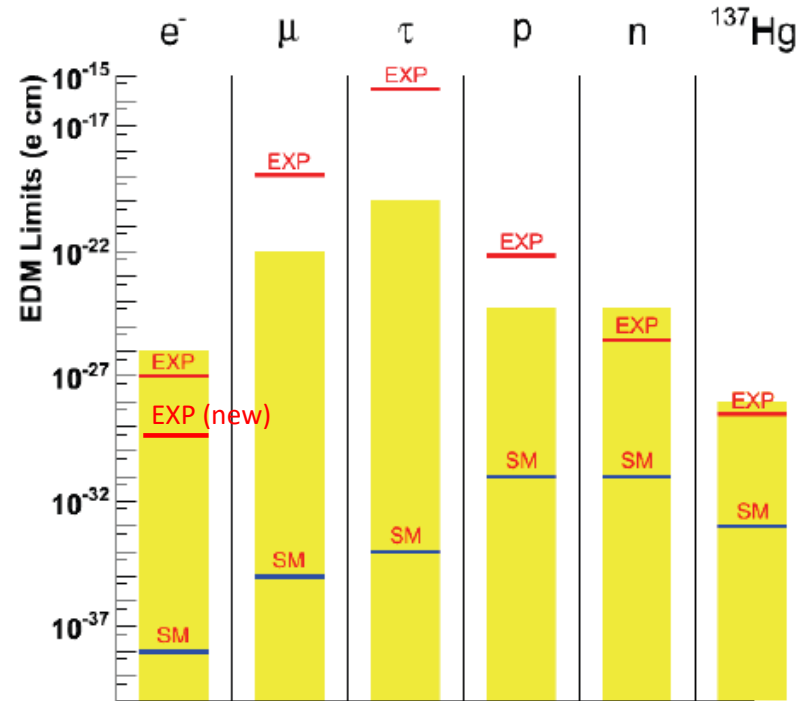
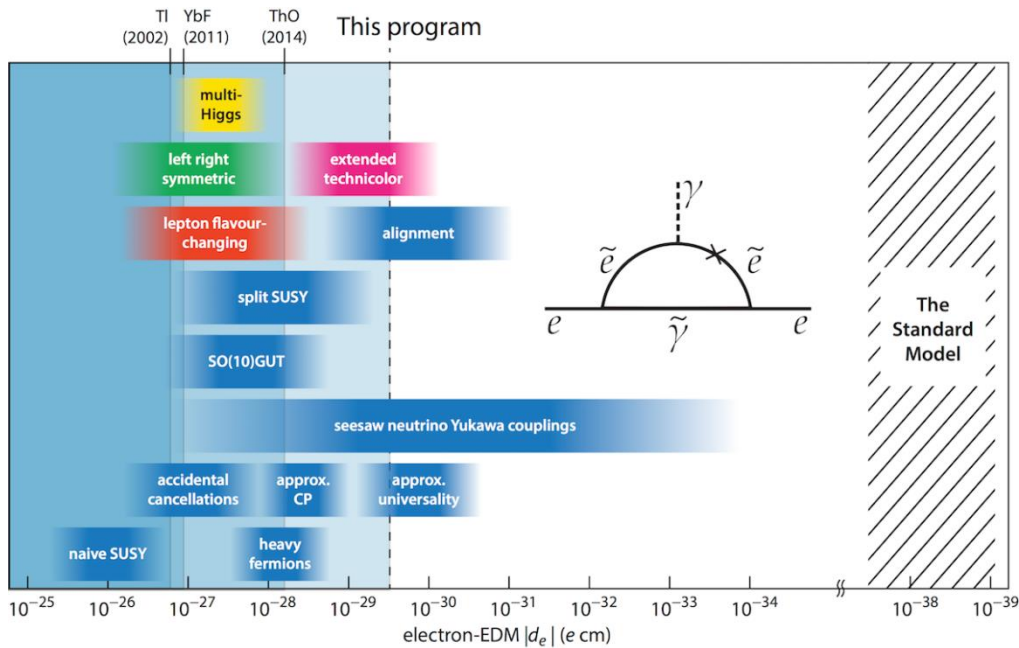
$$d \rightarrow d \text{ (even)}, \hat{S} = \frac{\bar{S}}{|S|} \rightarrow -\hat{S} = \frac{\bar{S}}{|S|} \quad [S \equiv ML^2T^{-1} \text{ odd}]$$

$$\Rightarrow \bar{d} \rightarrow -\bar{d}$$

T -violation and CPT symmetry $\Rightarrow CP$ violation.

Measurement of a non-zero EDM \Rightarrow measurement of CPV
SM EDMs tiny \Rightarrow any observation = new physics.

Worldwide program of EDM searches



More constraining than LHC for certain scenarios.

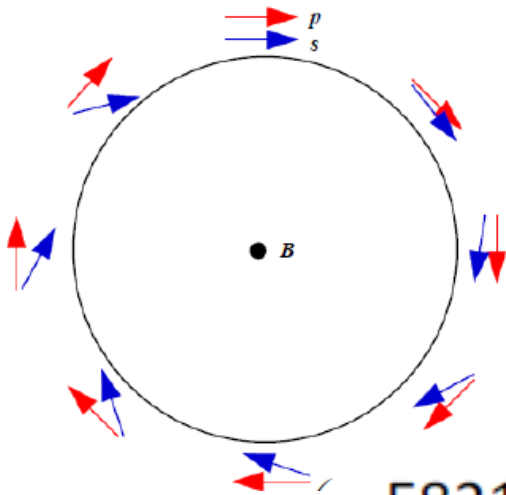
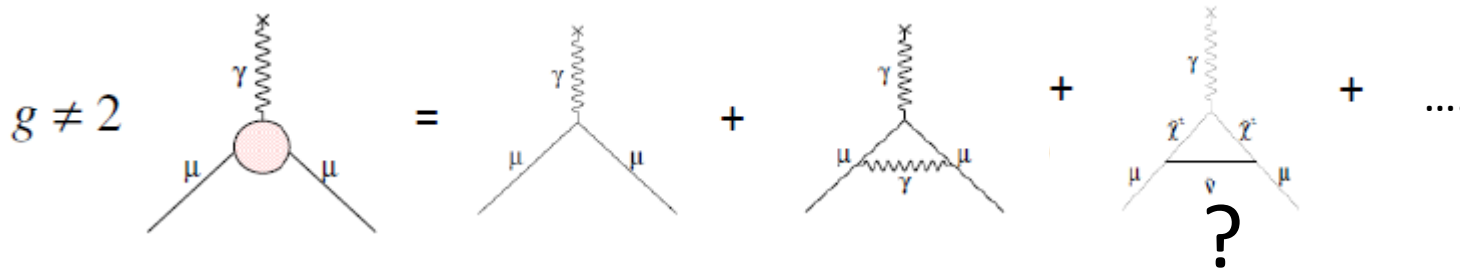
Eg e -EDM search (ACME)

$$\frac{|d_e|}{e} \sim \kappa \hbar c \left(\frac{\alpha_{eff}}{4\pi} \right)^n \left(\frac{m_e c^2}{\Lambda^2} \right) \sin \phi_{CP}$$

\Rightarrow new physics scale > 3 TeV (1 loop)

$g-2$ muon – a crack in the SM ?

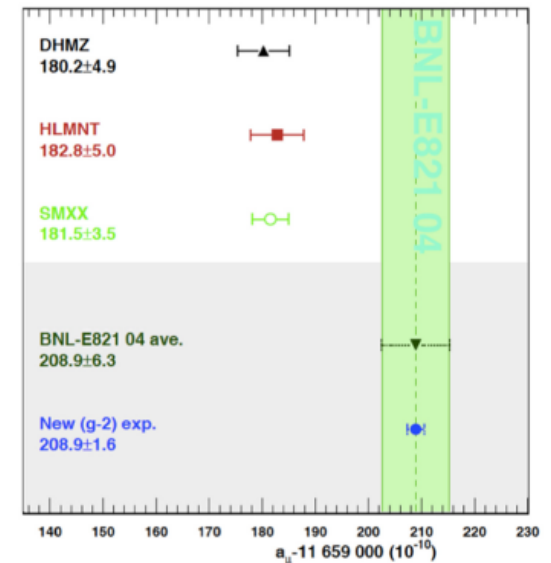
Magnetic moment for muon : $\mu = \frac{eg}{2m} S$



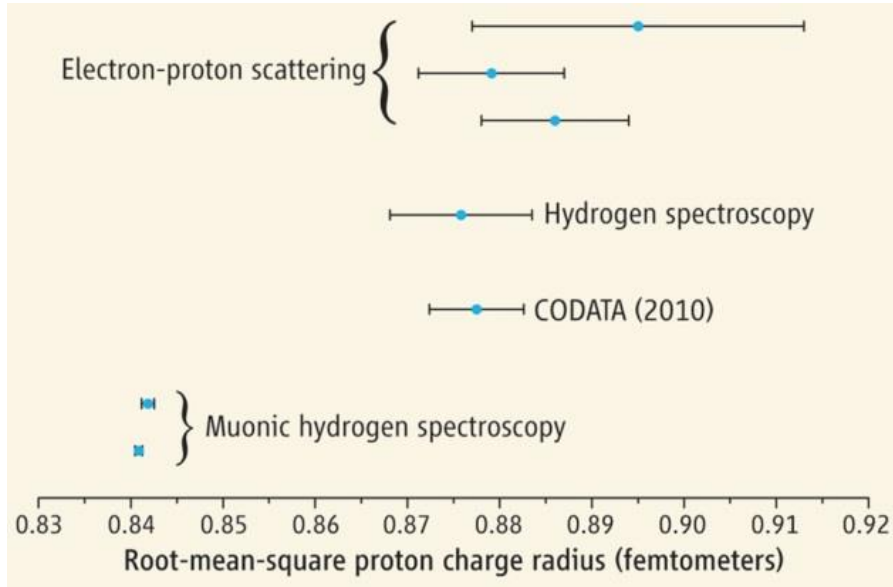
E821 Experiment

$$a_\mu = 11\,659\,208(6) \times 10^{-10} \quad (0.5\text{ppm})$$

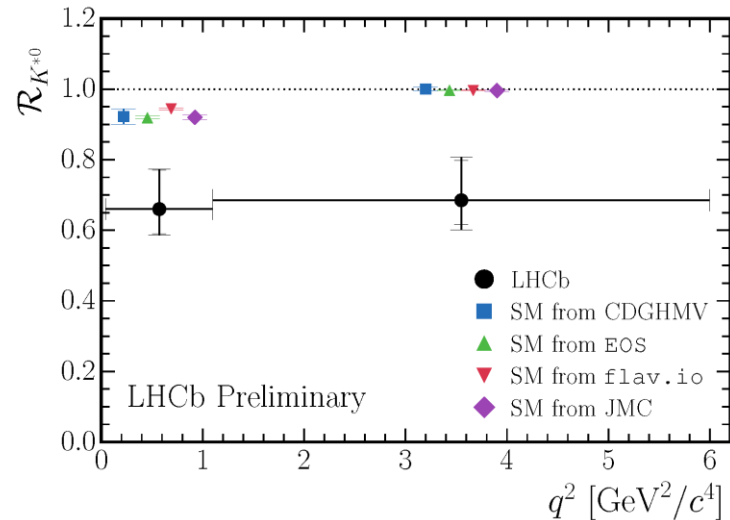
$\sim 3.5\sigma$ from theory!



Other cracks in the SM at the intensity frontier ?



Different proton radius measurements from muonic hydrogen and scattering+hydrogen spectroscopy.



Flavour anomalies

$$R(J/\psi) = \frac{\text{Br}(B_c^+ \rightarrow J/\psi \tau^+ \nu_\tau)}{\text{Br}(B_c^+ \rightarrow J/\psi \mu^+ \nu_\mu)} = 0.71 \pm 0.17 \pm 0.18$$

$$\text{SM } R(J/\psi) = 0.25$$

Deviations in $B \rightarrow K^* \mu^+ \mu^-$, $B \rightarrow K^* e^+ e^-$

Where are we ?

- A complete but poorly constrained SM in the top/EWSB sector.
 - We've discovered a very special particle and field
 - Precision needed.
- New physics processes excluded
 - masses up to 1-2 TeV from collider experiments.
 - >2 TeV from non-collider experiments (more scenario dependent)
- Some possible hints for new physics at 2-3 sigma level.
- The naturalness paradigm for new TeV-scale physics has (so far) not worked out.

Some interesting reactions

The Dawn of the Post-Naturalness Era

There are many indications that, following the recursive pattern of scientific revolutions, we are now witnessing the beginning of the phase of crisis. The lack of new physics in the initial stages of the LHC project is putting into question the logic of naturalness when applied to the Higgs; the absence of a positive detection in dark matter experiments is casting doubts about nature taking advantage of the WIMP miracle. We are not simply confronted with
Arxiv:1710.07663 (Giudice)

Saturday, August 06, 2016

The LHC “nightmare scenario” has come true.

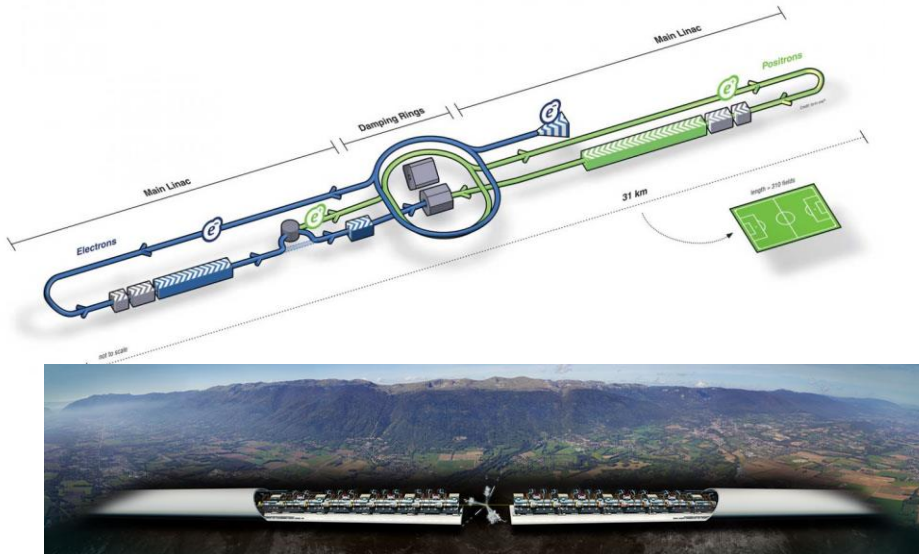
Bizarre : the LHC has delivered ~% of its data.

Where are we going ?

What will (almost certainly) happen ?

- The LHC will deliver 3000fb^{-1} by 2035
 - Sensitivity up to ~ 3 TeV in new physics
 - Greater precision on SM quantities
- New measurements of $g-2$ muon
 - x 5 improvement in sensitivity
- New searches for electron, neutron EDMs + other non-collider experiments (SHIP, nnbar are possibilities).
 - \sim order of magnitude improvement expected

What can happen ?



e^+e^- collider at

$$\sqrt{s} = 200 - 3000 \text{ GeV}$$

for precision Higgs/EWSB.

Precision frontier - new physics
from deviations wrt SM.



hadron-hadron collider

$$\text{at } \sqrt{s} \sim 100 \text{ TeV}$$

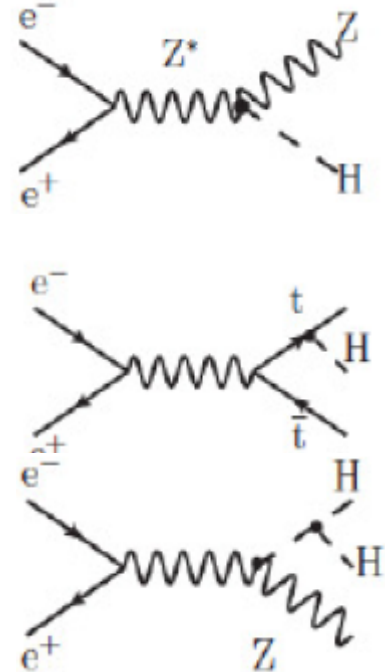
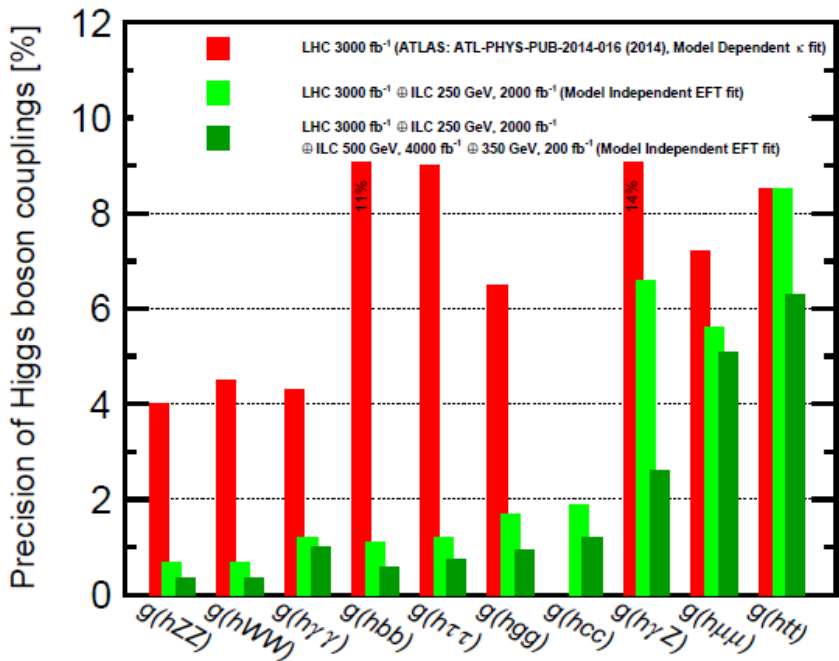
Precision Higgs/EWSB +
new physics searches
at energy frontier.

The collider market place

Collider	\sqrt{s} (TeV)	$\int \mathcal{L}$ (ab^{-1})	Location	Type	Date	Mature
ILC	0.25 0.5,1.0 ?	2	Japan	Linear e+e-	>2030	Shovel-ready TDR (2013)
CLIC	0.38 1.5,3	0.5 1.5,2	CERN	Linear e+e-	>2035	CDR (2012)
FCC-ee	0.25/0.36	10/2.6	CERN	Circular e+e-	>2035	No
LEP3	0.25	1	CERN	Circular e+e-	>2035	No
CEPC	0.25	5	China	Circular e+e-	>2035	CDR (2017)
HE-LHC	28	10	CERN	Circular hh	>2040	No
FCC-hh	100	10	CERN	Circular hh	>2045	No

$e^+ + e^- \Leftrightarrow$ high precision Higgs factory

$$\frac{\sigma(pp \rightarrow H + X)}{\sigma(pp \rightarrow X)} \sim 10^{-9} \quad \frac{\sigma(e^+e^- \rightarrow H + X)}{\sigma(e^+e^- \rightarrow X)} \sim 0.01$$



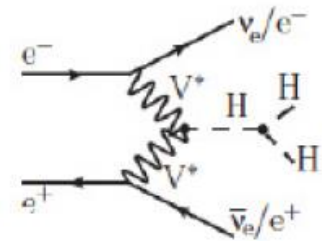
$\sqrt{s} = 250$ GeV
 g_Z , BR's,
 (LHC)-invisible

+

$\sqrt{s} = 500$ GeV
 g_t, g_{HHH}

+

$\sqrt{s} = 1000$ GeV
 g_{HHH}



ILC~CLIC~LEP3 sensitivity, FCCee (x 2-5 better)

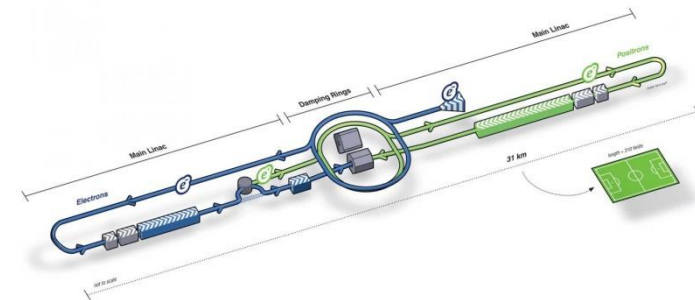
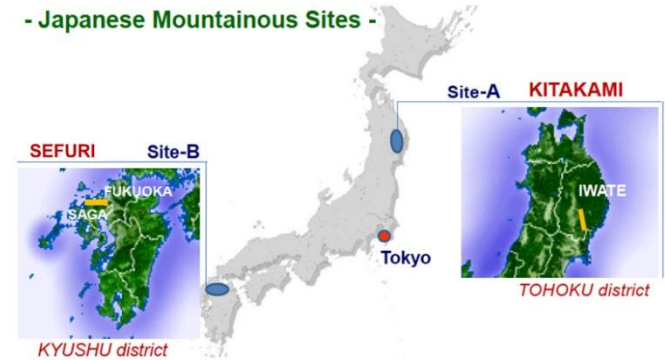
Tree-level sensitivity to new physics:

$\Delta g < 1\%$ for $\Lambda > 1$ TeV

Sensitivity to new physics in loops

International Linear collider

<i>ILC Physics Goals</i>	<i>500 GeV</i>	<i>350 GeV</i>	<i>250 GeV</i>
• precision Higgs couplings	✓	✓	✓
• gHWW and overall normalization of Higgs couplings	✓	✓	
• search for invisible and exotic Higgs decay modes	✓	✓	✓
• Higgs couplings to top	✓		
• Higgs self-coupling	✓		
• search for extended Higgs states	✓		
• precision electroweak couplings of the top quark	✓		
• precision W couplings	✓	✓	
• precision search for Z'	✓		
• search for supersymmetry	✓		
• search for Dark Matter	✓		
• top quark mass from threshold scan		✓	
• precision Higgs mass			✓



Original plan $\sqrt{s}=500$ GeV \rightarrow 1 TeV .

Now $\sqrt{s}=250$ GeV , maybe 380 GeV, 500 GeV

Most (not all) core measurements are kept but a lot is lost.

Project started in the 1990's. Decision (hoped for) in 2018.

Hadron-hadron colliders



Future circular collider

pp collider at $\sqrt{s} = 100$ TeV

New tunnel (80-100 km).

16 T magnets

FCC-ee first ($\sqrt{s} = 380$ GeV) ?



High Energy LHC

pp collider at $\sqrt{s} = 28$ TeV

Existing tunnel

16 T magnets

LEP3 first ($\sqrt{s} = 240$ GeV) ?

hh physics at 100 TeV

- An order of magnitude lower in probed distance scale
 - Factors $\sim 5-10$ in mass scale for new physics
 - Complete exploration of the Higgs sector (quartic coupling)
 - Test electroweak baryogenesis
 - High precision Higgs + SM measurements + DM searches
-
- No “no lose” theorem unlike Tevatron and LHC.
 - No consensus in community.
 - Magnet builders and theorists like it.
 - For the first time in ~ 50 years its not clear that a new insight will come from pushing the high energy frontier.



The cards are on the table



How do we play them ?

(Some) questions for the PP strategy ?

- What benefits Swedish activities and interests ?
- Higgs/EWSB
 - Is a reduced ILC still attractive ?
 - Can we afford a complete Higgs sector exploration ?
 - Is a FCC affordable/justified if the null hypothesis persists ?
- Is it better for CERN/the community to invest in smaller scale experiments instead of a FCC if the null hypothesis persists ?
- How should CERN spend our money in the next 5-10 years ?
- What is the role of national labs ?
- How does CERN maintain a leading role ?
-

- Swedish contribution to European Strategy Update
- Document for end of 2018
- Meeting at UU, later at Lund
- UU bid for community symposium in May 2019
- Take part !

Swedish PP strategy meeting
 Tuesday 13 Mar 2018, 12:30 → 17:05 Europe/Zurich
 Room 80101, Ångströmlaboratoriet, Lägerhyddsvägen 1, 752 37 Uppsala (Other Institutes)

Videokonferens
Rooms Swedish_PP_strategy_meeting

12:30 → 12:50 Updated PP strategy

12:30 Intro
 Speaker: David Anthony Milstead (Stockholm University (SE))
 Strategy-Update to ... Strategy-Update to ...

12:40 Hosting strategy meetings
 Speaker: Richard Brenner (Uppsala University (SE))
 Strategy_Symposiu...

12:50 → 14:30 Energy frontier and collider experiments

12:50 ATLAS experiment
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 slides.pdf

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13:20 Swedish competence and involvement in accelerator development of future machines
 Speaker: Roger Ruber (Uppsala University (SE))
 20180313-Future_A... 20180313-Future_A...

13:35 Future colliders
 Speakers: Caterina Doglioni (Lund University (SE)), David Anthony Milstead (Stockholm University (SE))
 Doglioni_FutureHad... strategy-e+e--summ... strategy-e+e--summ...

13:50 Discussion

14:30 → 15:15 Intensity frontier

14:30 SHIP
 Speaker: Richard Brenner (Uppsala University (SE))
 Particle_Strategy_S...

14:40 HIBEAM/nंबर
 Speaker: David Anthony Milstead (Stockholm University (SE))
 HIBEAM-strategy.pdf HIBEAM-strategy.pp...

14:50 Search for Light Dark Matter using a primary electron beam
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 LDM_eBeam_PPstr...

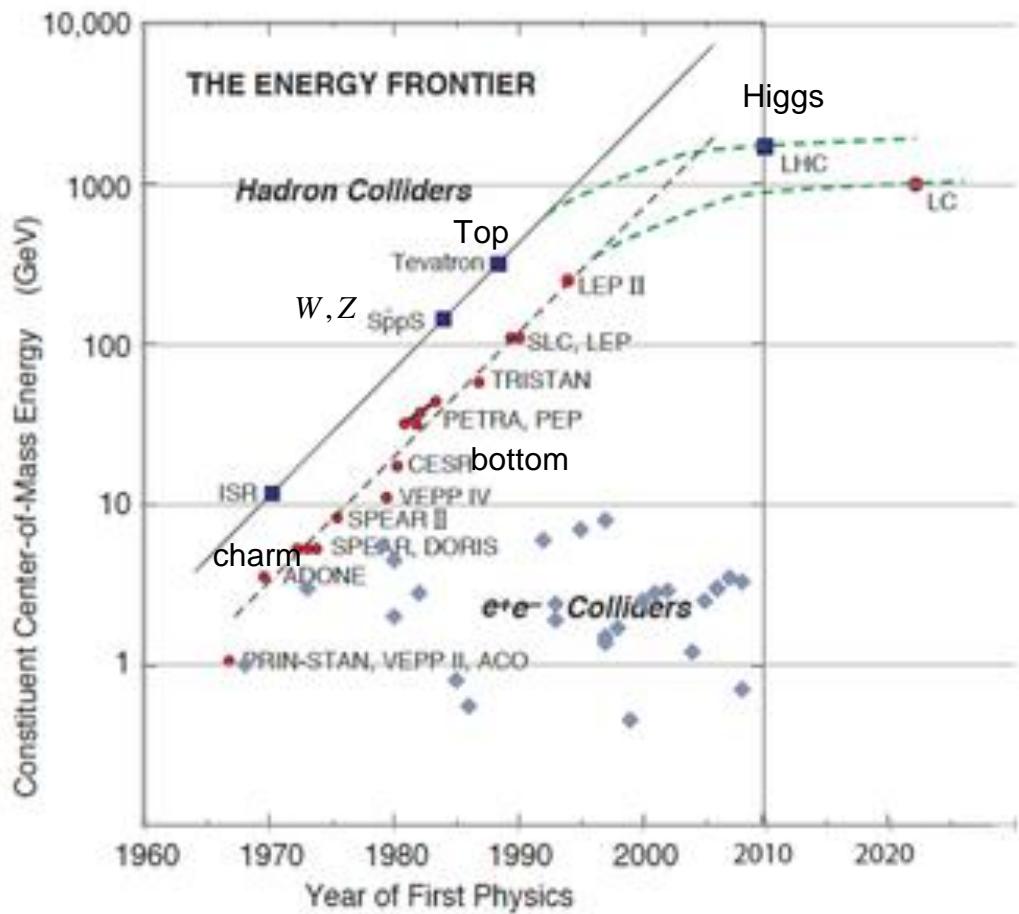
Swedish process ongoing
 Submission at the end of 2018

Summary

- The SM test paradigm is entering its final phase (EWSB)
- The BSM search paradigm has (so far) failed, as has its main motivator, naturalness.
- Lots of interesting projects and interesting unanswered questions.
- An e^+e^- collider is essential for progress.
- Its not obvious how we play the cards.
 - Tectonic plates will shift after Japan's ILC decision.
- Take part in the strategy update.

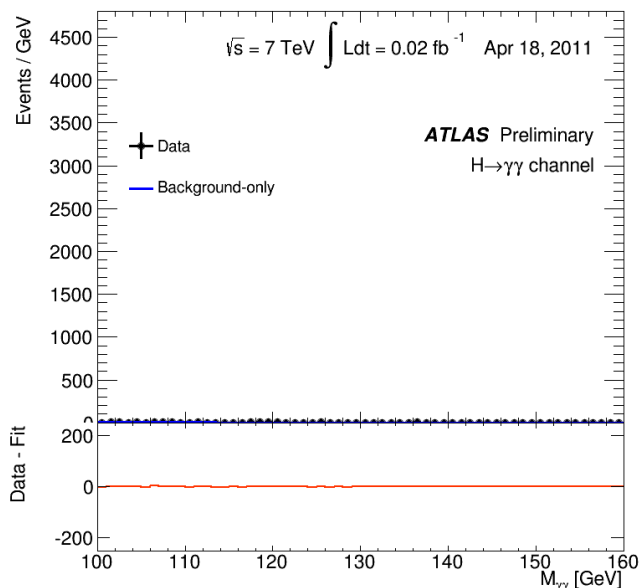
The European Particle Physics Strategy

- 2006
 - First strategy
 - http://council.web.cern.ch/sites/council.web.cern.ch/files/European_Strategy/ESStatement.pdf
- 2013
 - Strategy update
 - <https://cds.cern.ch/record/1567258/files/esc-e-106.pdf>
- 2018
 - Strategy update process starts (submissions at end of 2018)
- 2019
 - Community Symposium (UU have bid to host)
- 2020
 - Strategy update

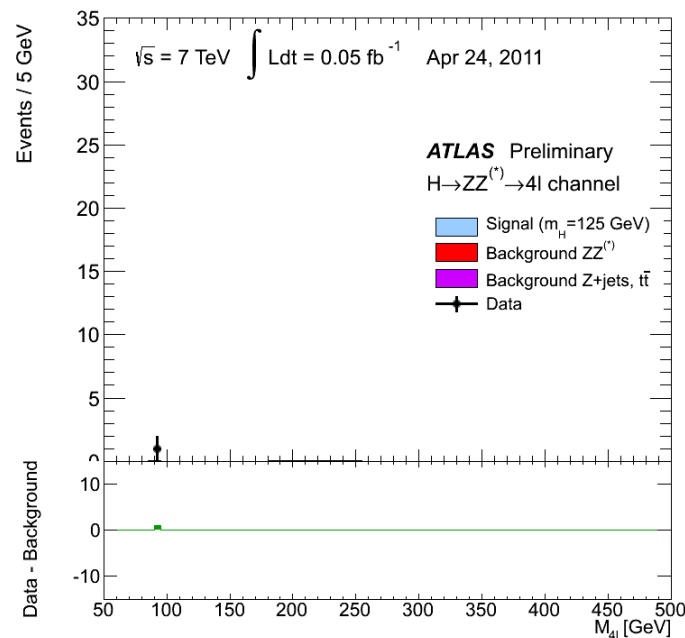


Observation of a Higgs \rightarrow boson pairs

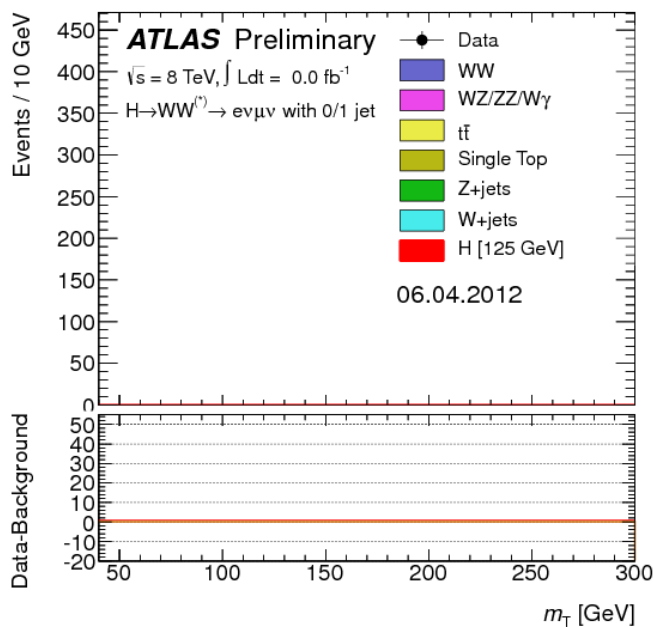
$$H \rightarrow \gamma\gamma$$



$$H \rightarrow ZZ \rightarrow 4\ell$$



$$H \rightarrow WW \rightarrow e\nu\mu\nu$$



And (equally importantly) - no observation in any unexpected channel (yet).

Mixing scales

Collide protons at c.o.m. energy E and measure photons from collision.

Distance scale for collision $\sim \frac{1}{E}$ reduces with E .

\Rightarrow Low $E \Leftrightarrow$ long distance and low photon energy $\langle E_\gamma \rangle$

$E >$ Planck mass \rightarrow Black hole !

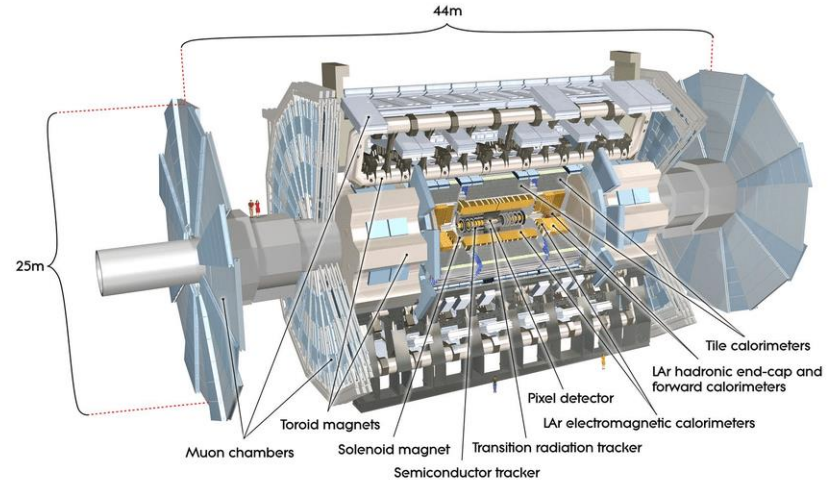
\rightarrow Distance scale, Schwarzschild radius: $R_s = 2GE$, increases with E .

\rightarrow Hawking radiation photons $\langle E_\gamma \rangle \sim \frac{k}{E}$ decreases with E .

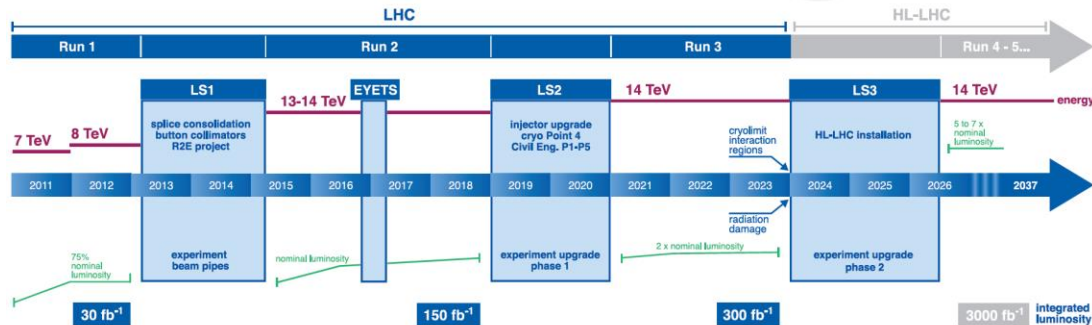
\Rightarrow High $E \Leftrightarrow$ long distance and low photon energy $\langle E_\gamma \rangle$

\Rightarrow Nature may not care about our prejudice for decoupled scales.

Large Hadron Collider



LHC / HL-LHC Plan

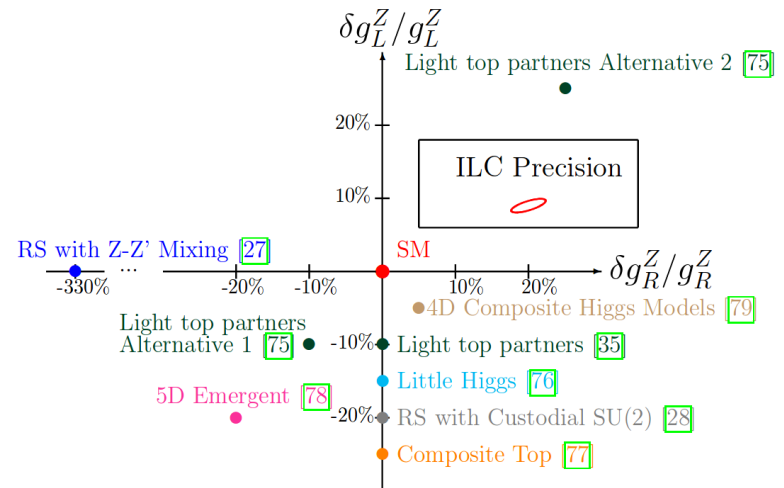
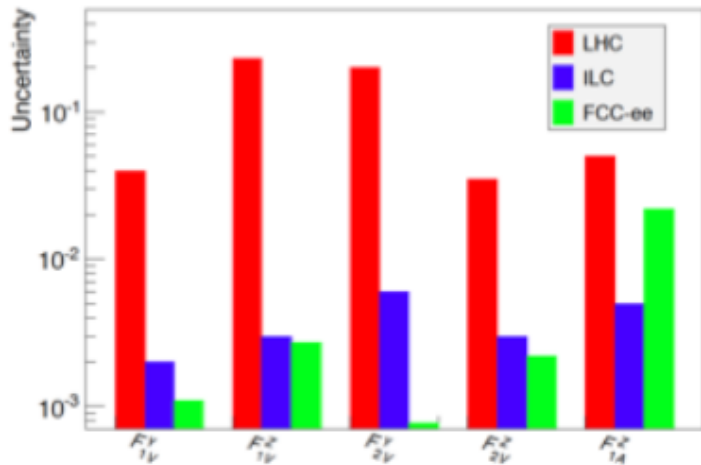
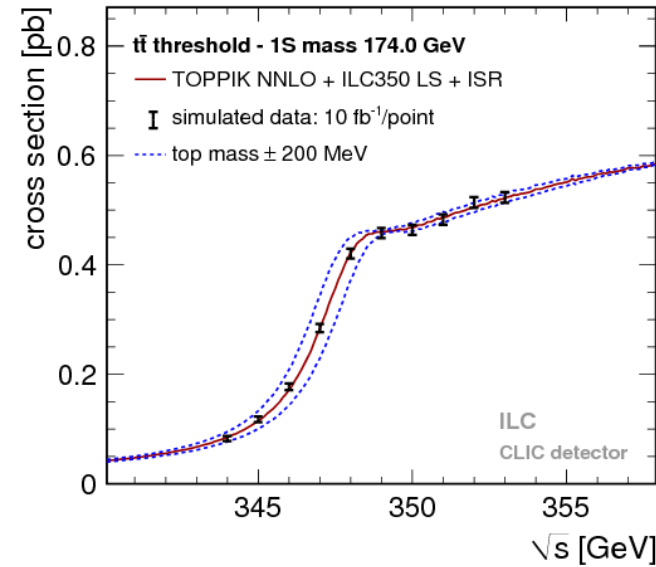


Luminosity and detector upgrades.

Taken ~% of total luminosity

Top Physics ($\sqrt{s} > 340$ GeV)

High precision threshold scan



Boson couplings

Organisational issues

l) Future major facilities in Europe and elsewhere require collaboration on a global scale. **Don't even think of building a European collider outside of CERN!**

m) A Memorandum of Understanding has been signed by CERN and the European Commission, and various cooperative activities are under way. Communication with the European Strategy Forum on Research Infrastructures (ESFRI) has led to agreement on the participation of CERN in the **EU strategy** Working Group. The particle physics community has been actively involved in European Union framework programmes. *CERN and the particle physics community should strengthen their relations with the European Commission in order to participate further in the development of the European Research Area.*

Wider impact of particle physics

n) **Outreach** Sharing the excitement of scientific discoveries with the public is part of our duty as researchers. Many groups work enthusiastically in public engagement. They are assisted by a network of community groups (EPPCN) and an international outreach group (IPPOG). For example, they helped attract tremendous public attention and interest around the world at the start of the LHC and the discovery of the Higgs boson. *Outreach and communication in particle physics should receive adequate funding and be recognised as a central component of the scientific activity. EPPCN and IPPOG should both report regularly to the Council.*

o) **Spin-offs** Knowledge and technology developed for particle physics research have made a lasting impact on society. These technologies are also being advanced by others leading to mutual benefits. Knowledge and technology transfer is strongly promoted in most countries. The HEPTech network has been created to coordinate and promote this activity, and to provide benefit to the European industries. *HEPTech should pursue and amplify its efforts and continue reporting regularly to the Council.*

p) **Training** Particle physics research requires a wide range of skills and knowledge. Many young physicists, engineers and teachers are trained at CERN, in national laboratories and universities. They subsequently transfer their expertise to society and industry. Education and training in key technologies are also crucial for the needs of the field. *CERN, together with national funding agencies, institutes, laboratories and universities, should continue supporting and further develop coordinated programmes for education and training.*

Concluding recommendations

q) This is the first update of the European Strategy for Particle Physics. It was prepared by the European Strategy Group based on the scientific input from the Preparatory Group with the participation of representatives of the Candidate for Accession to Membership, the Associate Member States, the Observer States and other organisations. Such periodic updates at intervals of about five years are essential. *Updates should continue to be undertaken according to the principles applied on the present occasion. The organisational framework for the Council Sessions dealing with European Strategy matters and the mechanism for implementation and follow-up of the Strategy should be revisited in the light of the experience gained since 2006.*

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Videconferencerooms Swedish_PP_strategy_meeting

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HIBEAM-strategy.pdf HIBEAM-strategy.pp...

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LDM_eBeam_PPstr...

Swedish process ongoing
Submission at the end of 2018

Standard Model

Strong

Strong CP problem

Non-perturbative QCD:
confinement

Yet to fail precision tests
at 10^{-18} m in the
perturbative sector

Electroweak

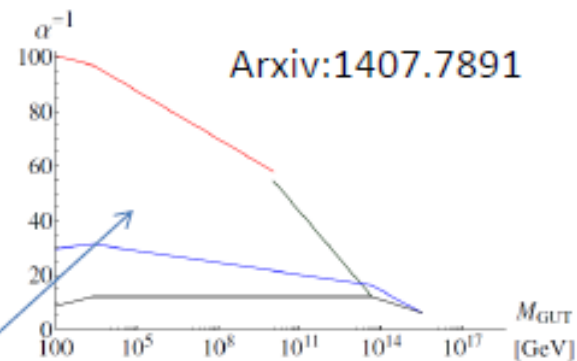
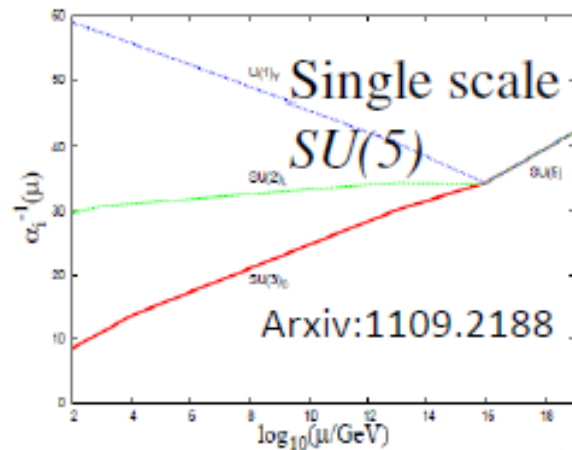
Symmetry breaking:
Higgs properties still to be
precisely determined,
Naturalness
Top quark properties

Non-perturbative EW
effects still to be observed.

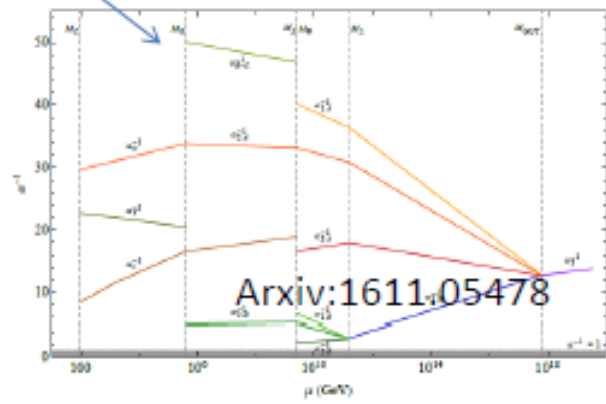
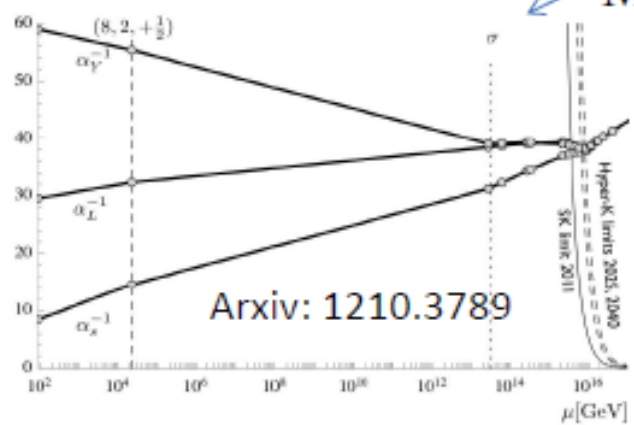
+ neutrino mass, unification of strong+EW, no dark matter candidate,
baryogenesis, electric charge quantisation ...

Single and multi-scale Unification

Multiscale assume new physics at scale(s) between EW and GUT scale.



Multiscale



Another fine-tuning problem

CPV term in QCD Lagrangian:

$$\mathcal{L} = \mathcal{L}_0 + \mathcal{L}_{\text{CPV-strong}} \quad ; \quad \mathcal{L}_{\text{CPV-strong}} = \bar{\theta} \frac{\alpha_s}{8\pi} G_{a\mu\nu} \tilde{G}_a^{\mu\nu}$$

$$\bar{\theta} = \theta_{\text{QCD}} + \arg(\det M_q) \quad 0 \leq \bar{\theta} \leq 2\pi$$

$\theta_{\text{QCD}} \equiv$ QCD vacuum term, $M_q =$ quark mass matrix (EW)

For $\bar{\theta} \sim \mathcal{O}(1)$ EDM neutron $d \sim 10^{-16}$ ecm.

Experimental limit on neutron EDM $d < 10^{-26}$ ecm

$$\Rightarrow \bar{\theta} < \mathcal{O}(10^{-10})$$

\Rightarrow Fine tuning between θ_{QCD} and $\arg(\det M_q)$ (\equiv strong and EW) !

\Rightarrow This is not "natural" \Rightarrow Strong CP problem.

