

# OFF THE BEATEN TRACK

## LONG-LIVED PARTICLES AT THE LHC



TOVA HOLMES  
NUCLEAR AND PARTICLE PHYSICS SEMINAR  
UPPSALA UNIVERSITY  
OCTOBER 22, 2020



my goal:

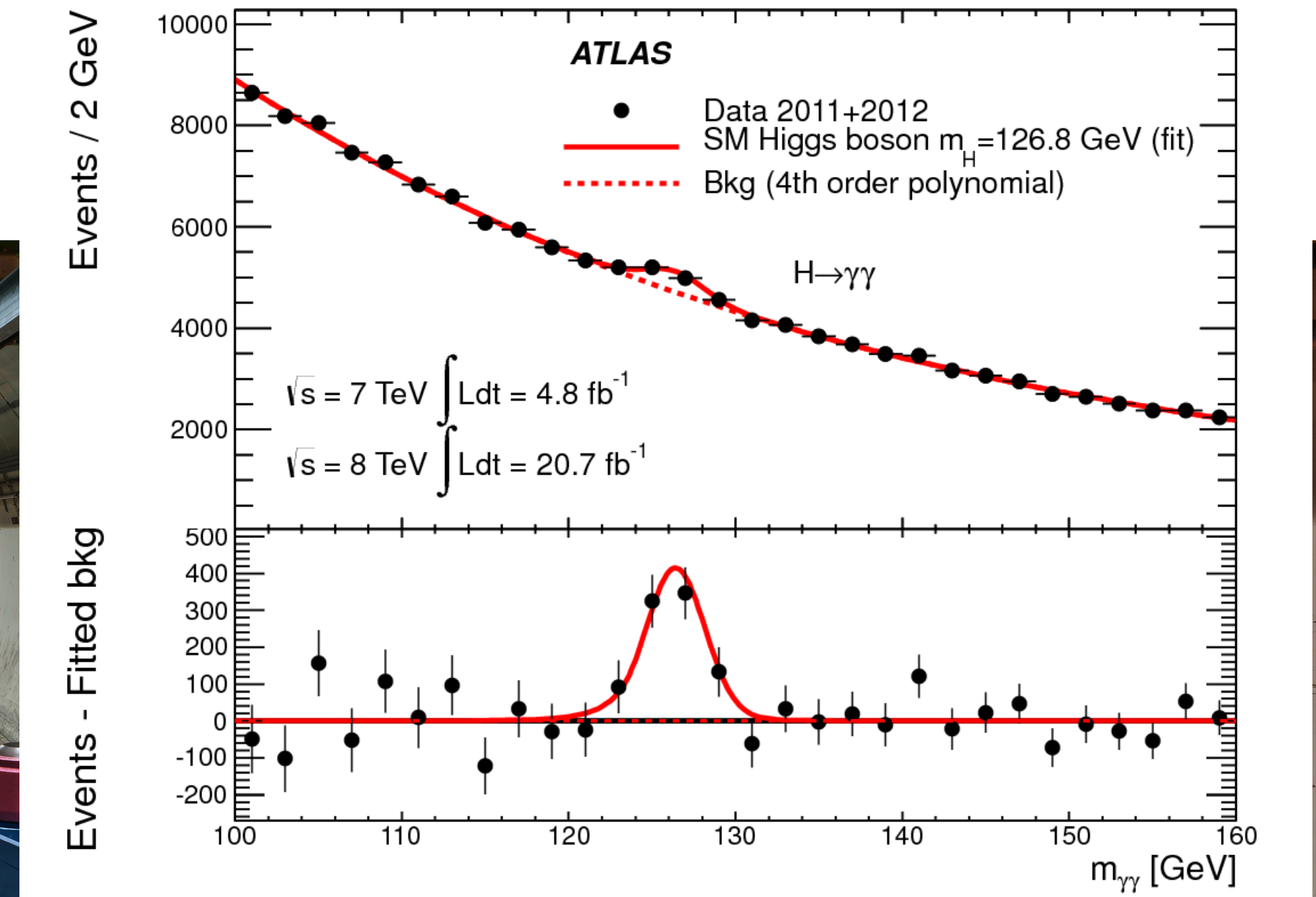
find new  
fundamental particles





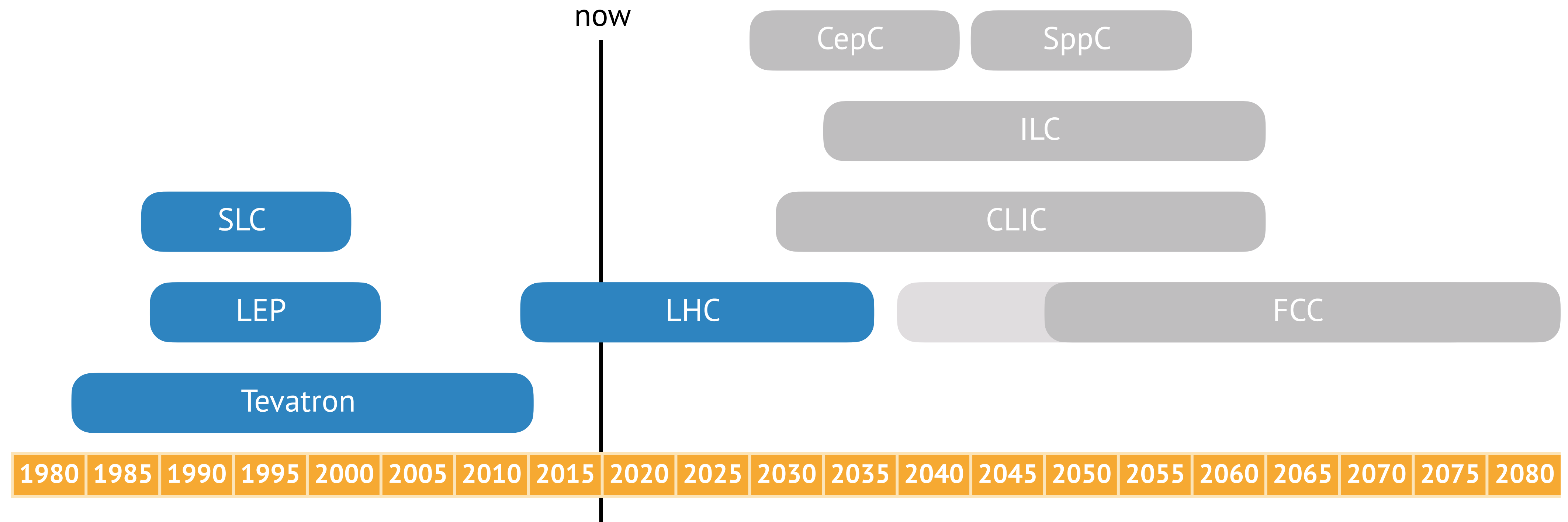
Most recent machine to  
accomplish that goal:  
the Large Hadron Collider

## Higgs discovery, 2012



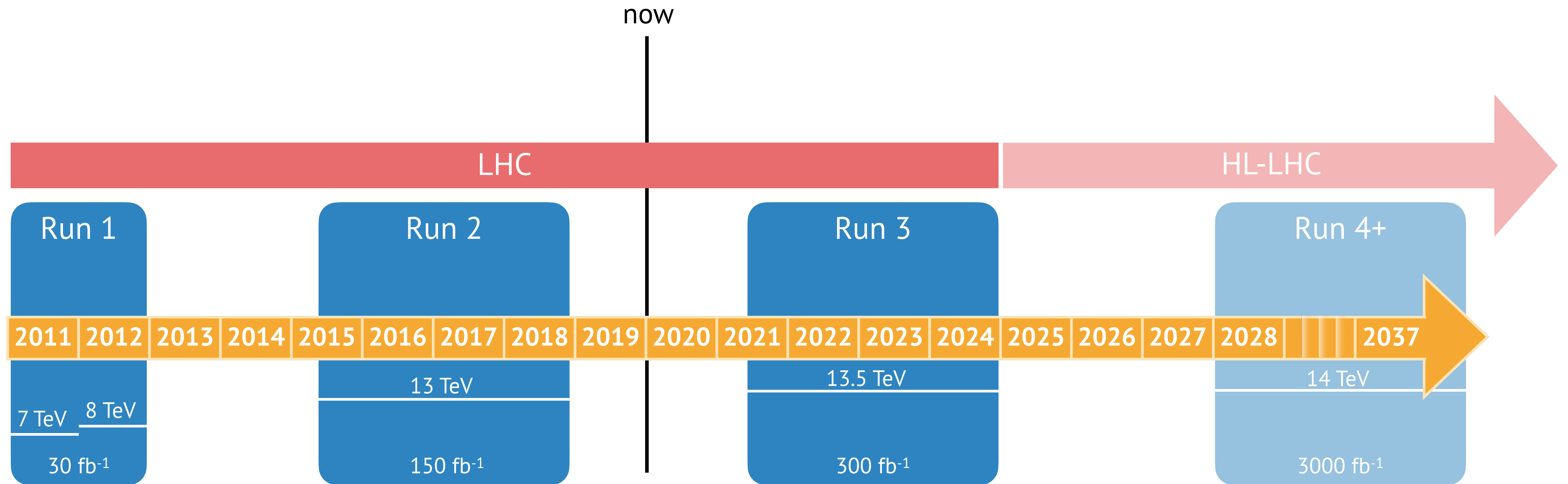


# The Energy Frontier





# The LHC





# The LHC

Run 1 → Run 2

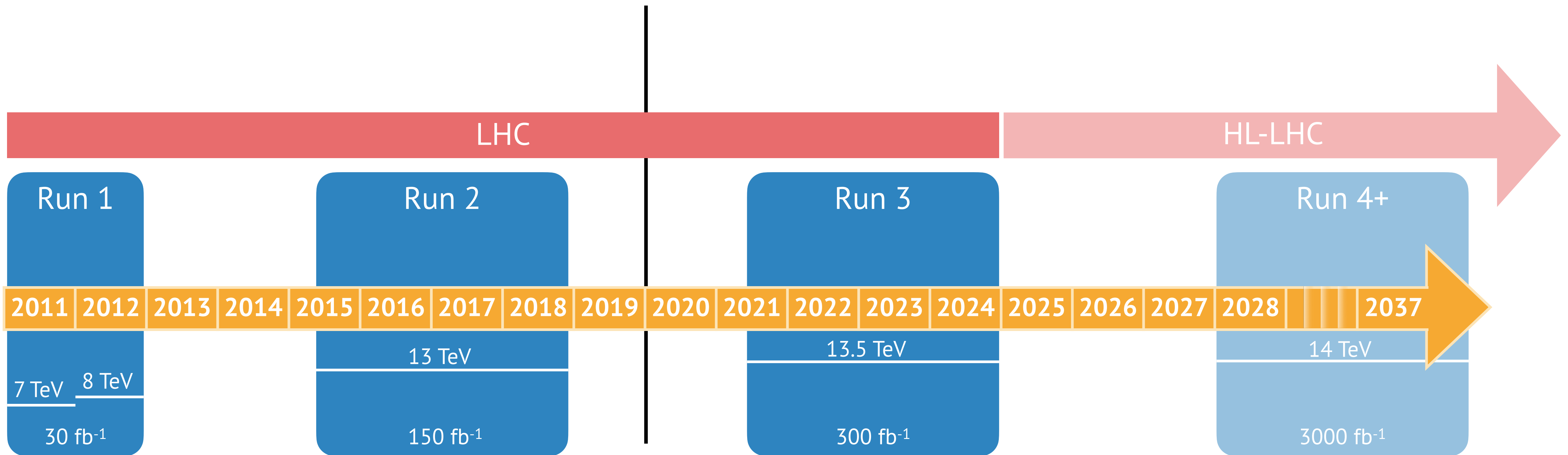
Luminosity x5  
Energy +60%

Run 2 → Run 3

Luminosity x1  
Energy +4%

Run 3 → Run 4+

Luminosity x10  
Energy +4%



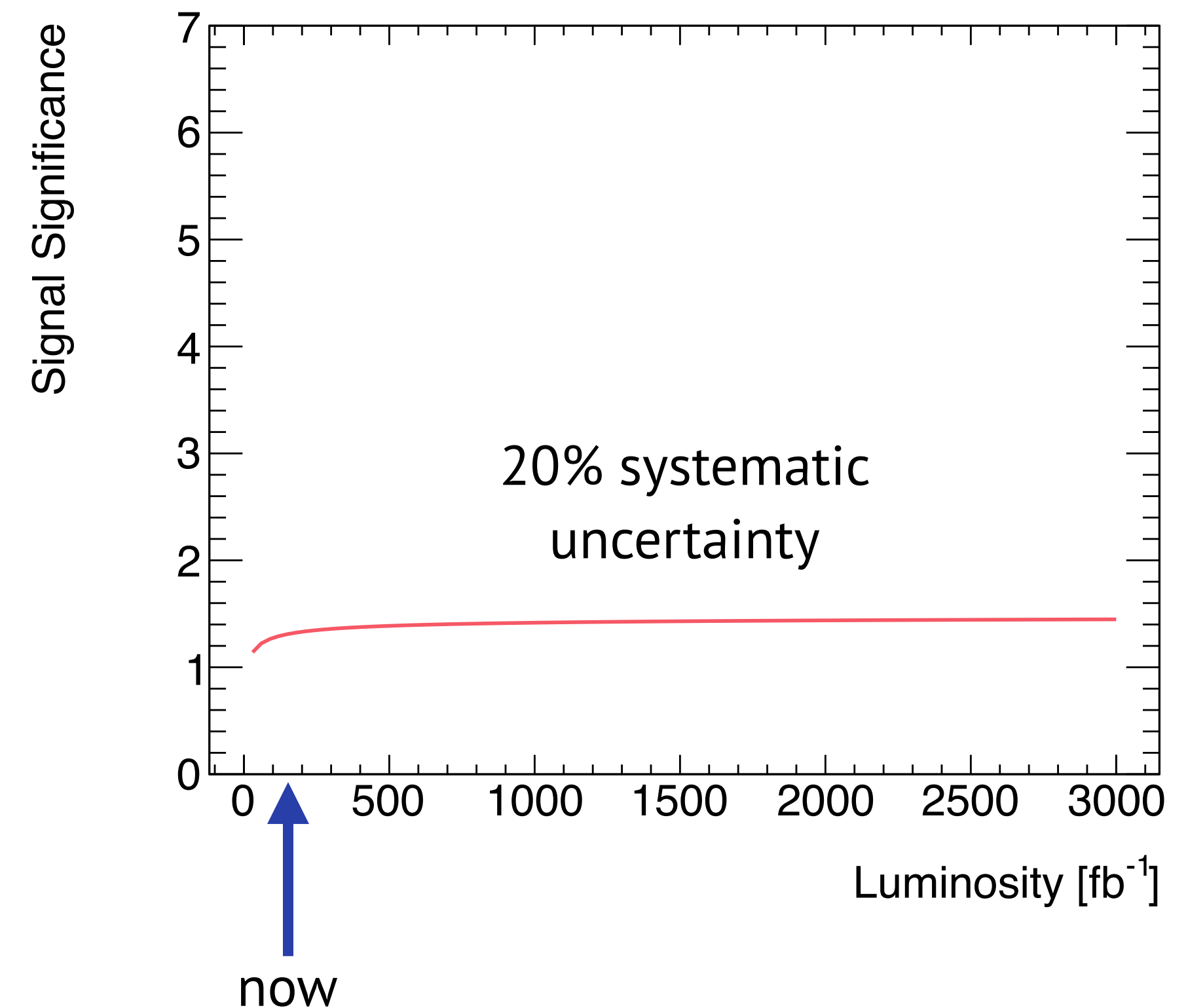


# The Worst Case

$$S \sim \frac{n_{sig}}{\sigma_{bkg}}$$

if  $\sigma_{bkg}$  is **driven by systematics** that are a fixed fraction of the total background

$\sigma_{bkg} \propto$  **luminosity**, and more data doesn't give more sensitivity



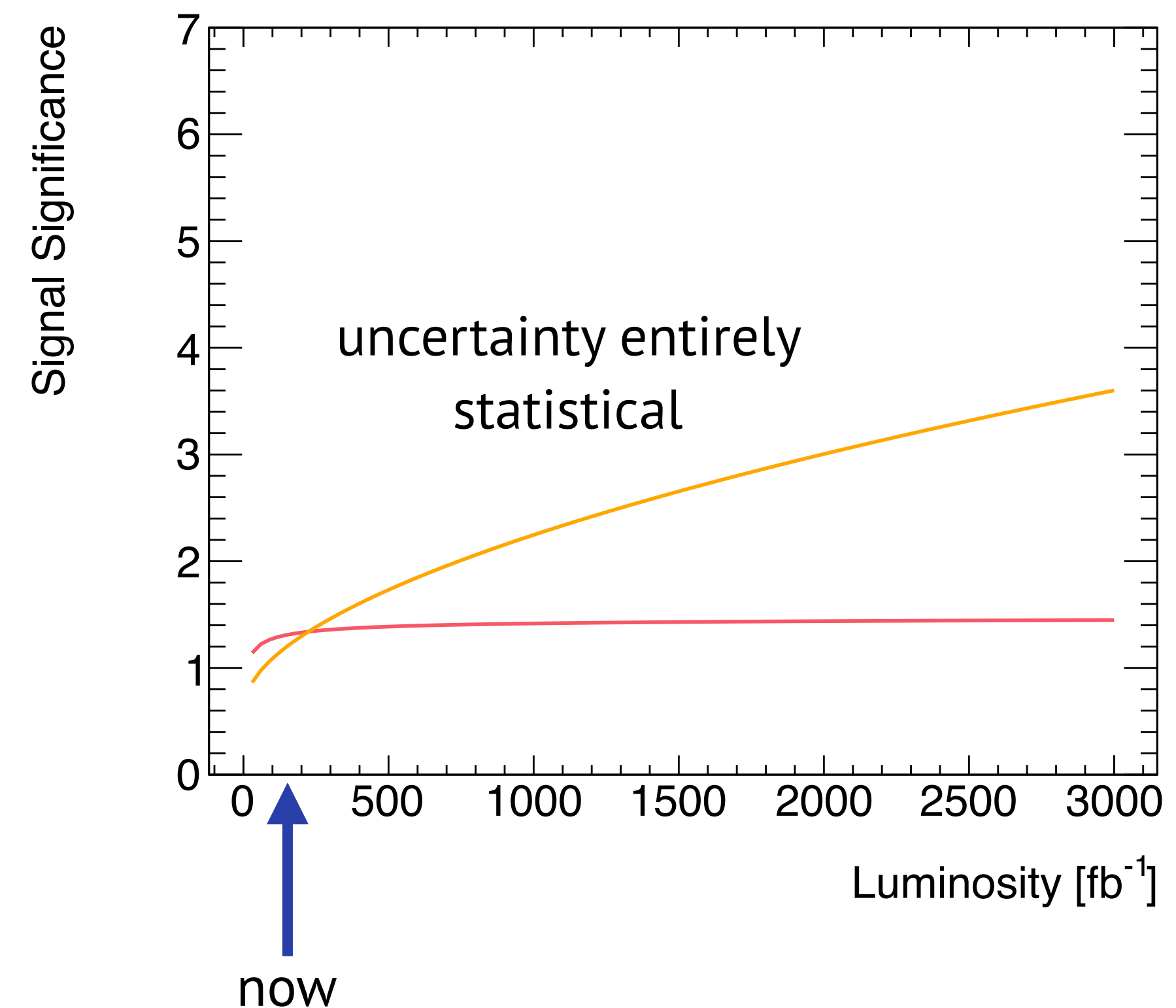


# The Medium Case

$$S \sim \frac{n_{sig}}{\sigma_{bkg}}$$

if  $\sigma_{bkg}$  is **driven by statistical uncertainty**

$\sigma_{bkg} \propto \sqrt{\text{(\textbf{luminosity})}}$ , and sensitivity increases as roughly the square root of luminosity



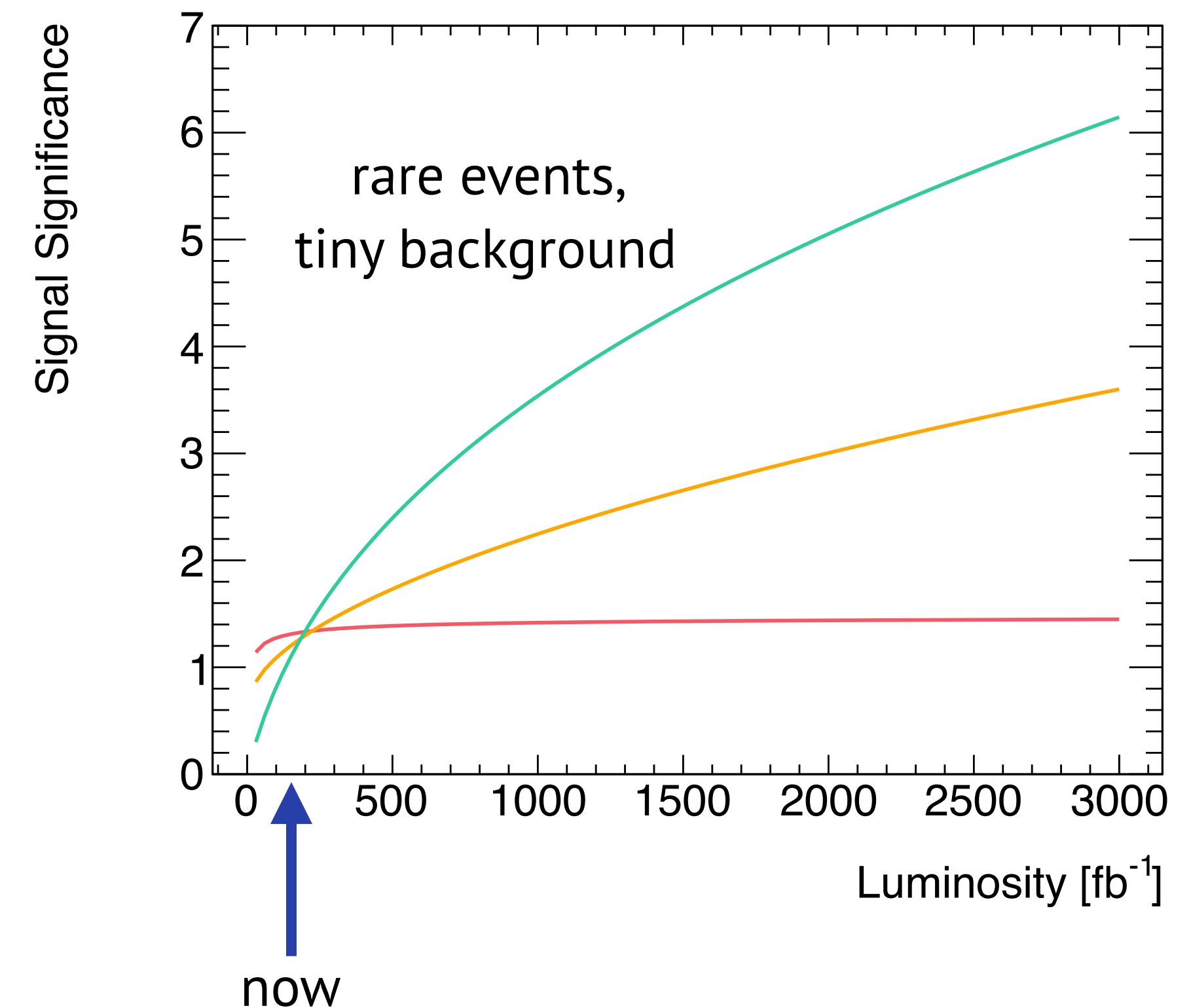


# The Best Case

$$S \sim \frac{n_{sig}}{\sigma_{bkg}}$$

if the **standard model background is tiny**

then it's only a question of producing and reconstructing your signal, and **sensitivity is proportional to luminosity**



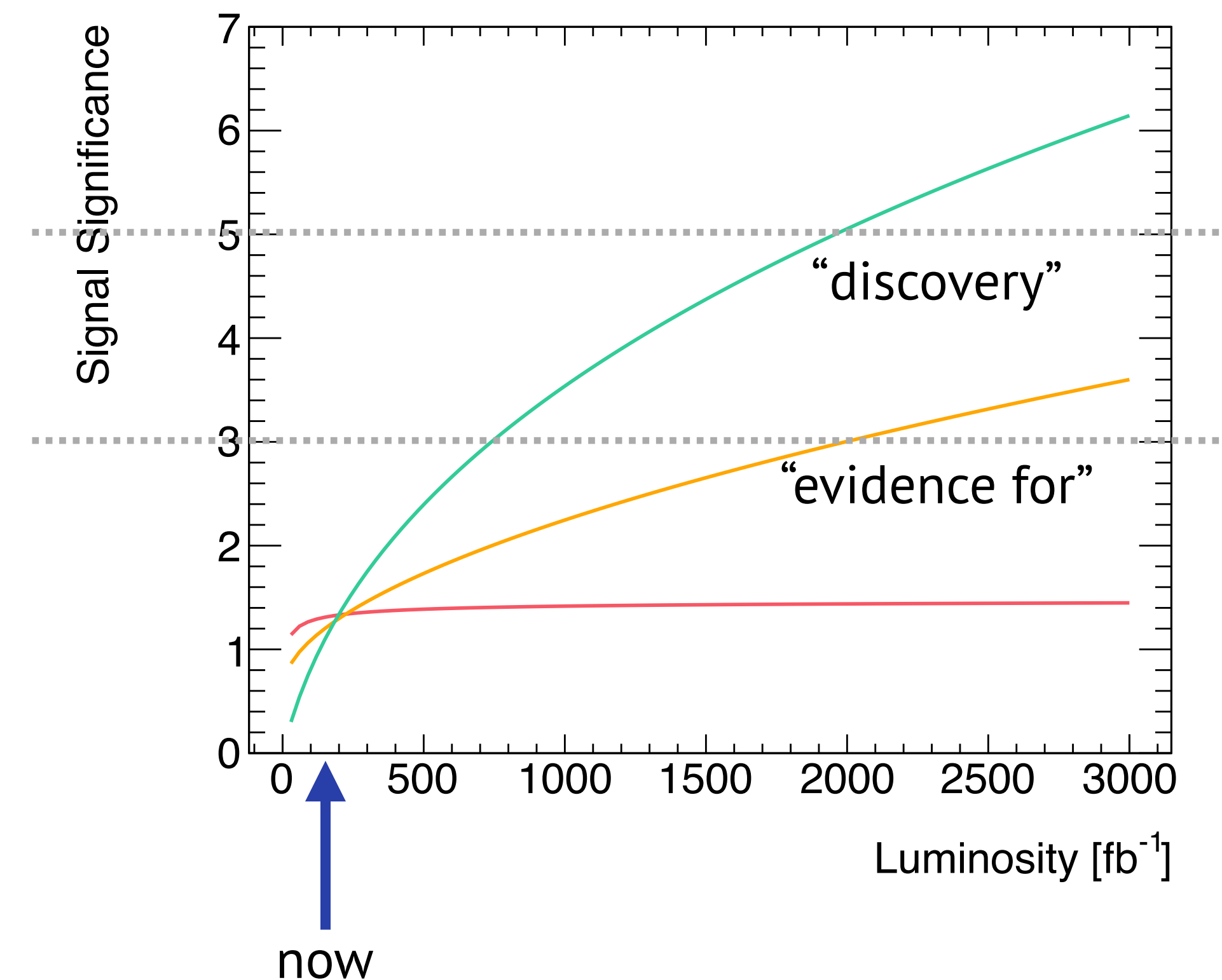


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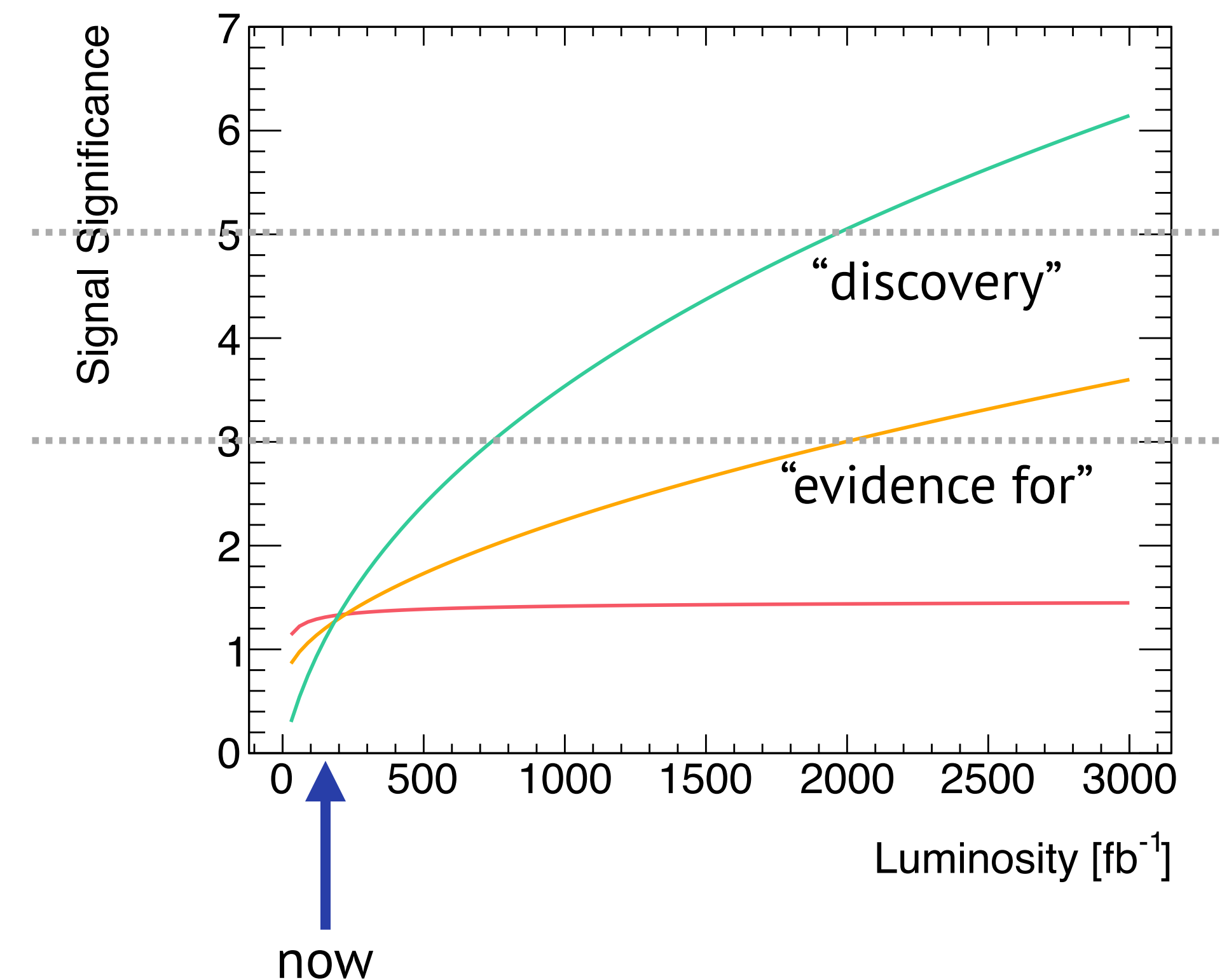
then it's only a question of producing and reconstructing your signal, and **sensitivity is proportional to luminosity**





# The Best Case

As we move forward with the LHC  
the physics that continues to be most  
exciting is rare stuff with small  
backgrounds

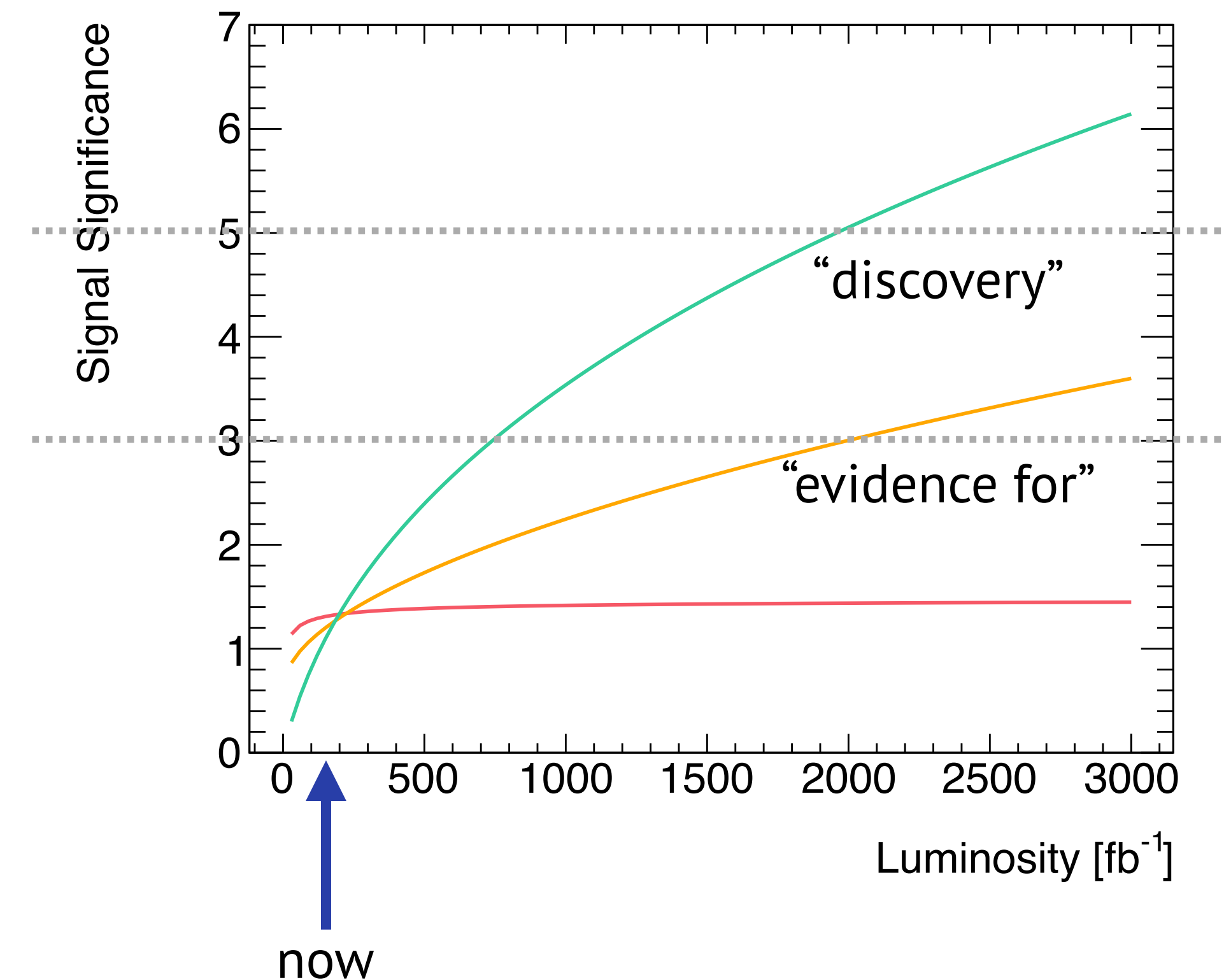




# The Best Case

As we move forward with the LHC  
the physics that continues to be sought  
exci

OR  
stuff that's really hard to  
reconstruct  
→ effectively “rare”





# The Best Case

As we move forward with the LHC  
the physics that continues to be sought  
excited

OR

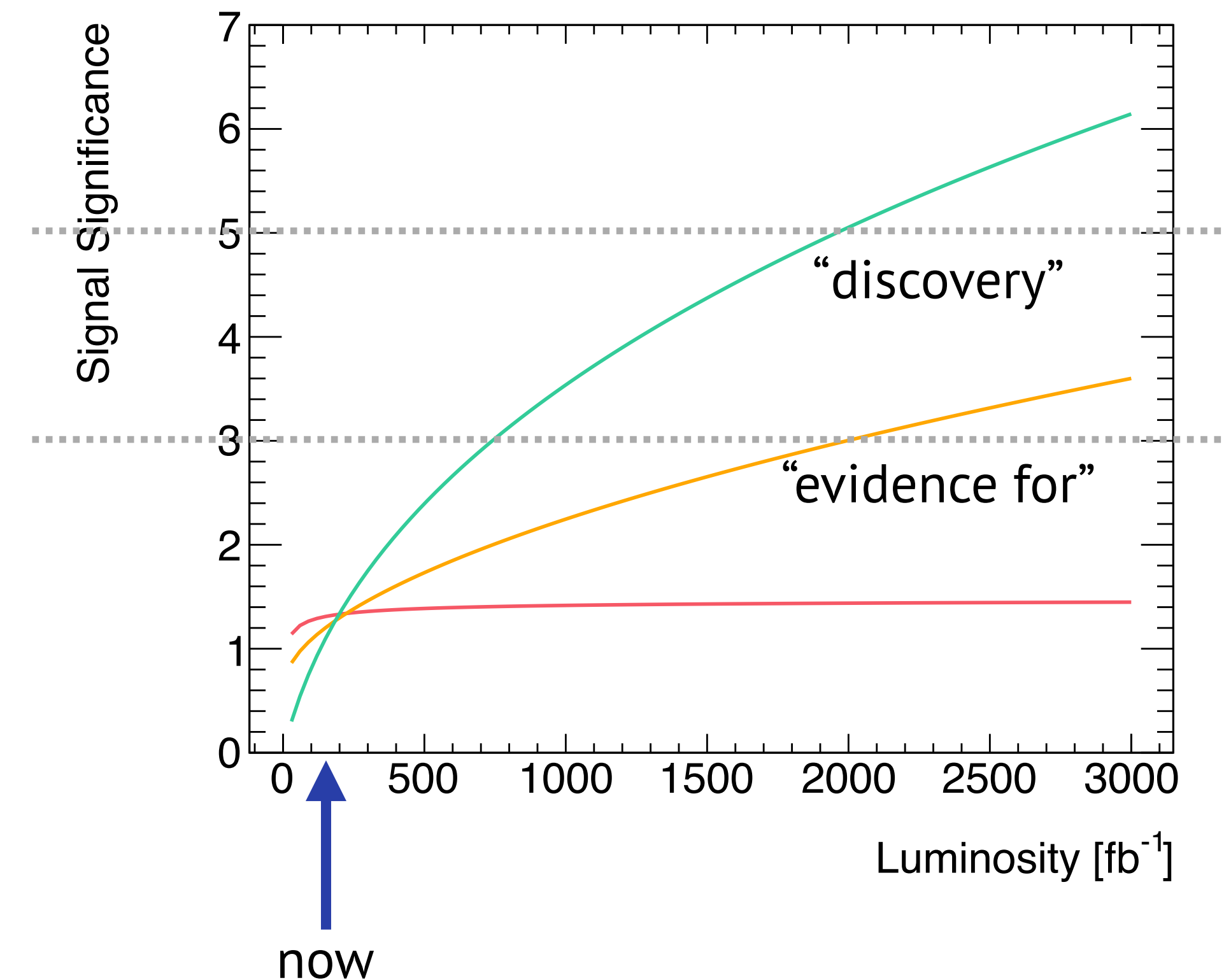
stuff that's really hard to

OR

stuff we haven't  
looked for at all yet

or  
structure

very "rare"





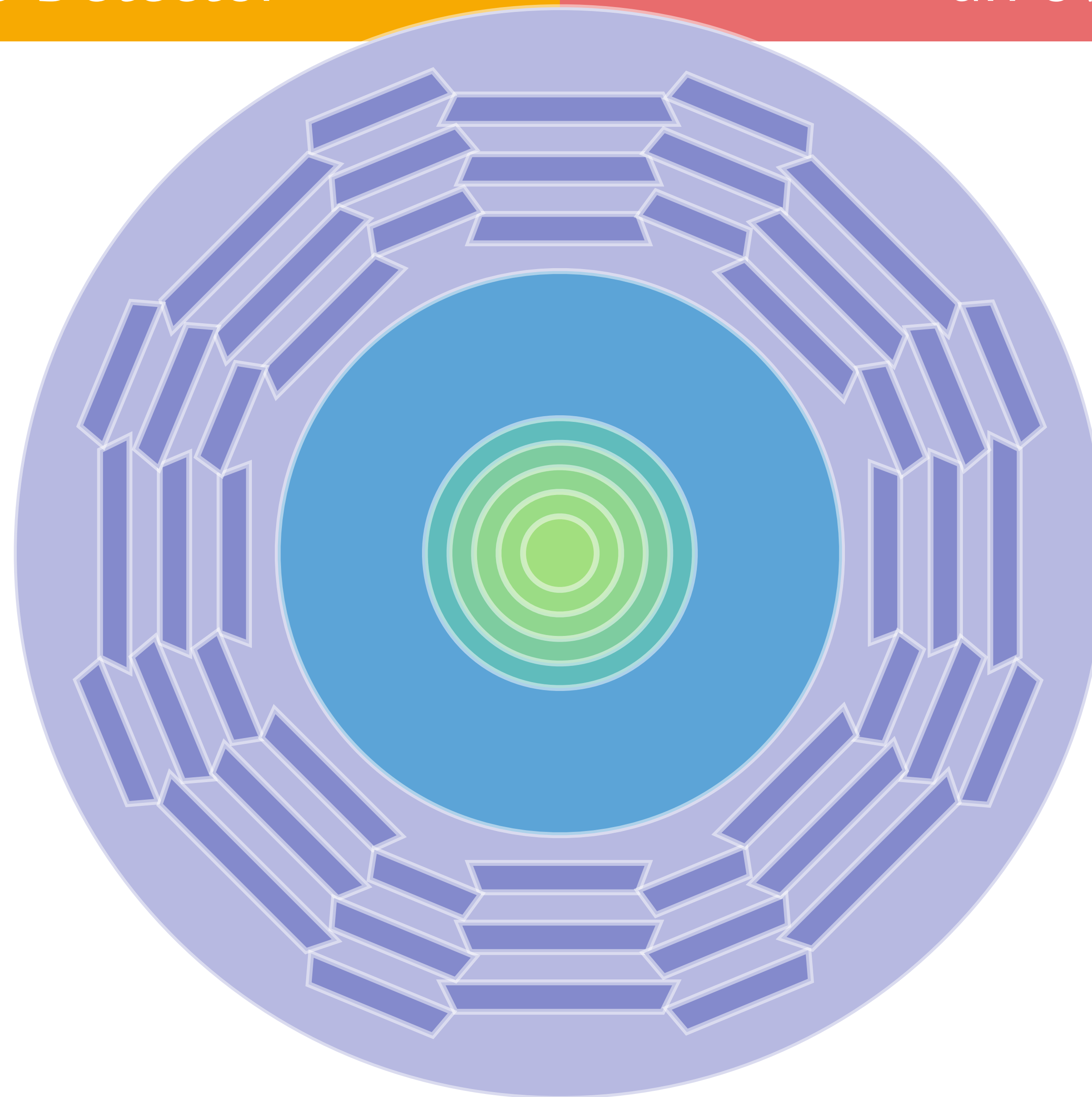
# LHC experiments are huge

searched for many things  
only things left are...

not well motivated

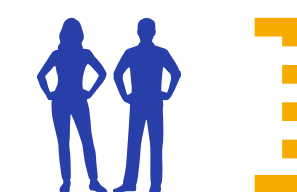
or really hard

Nearly **10K** people working  
on **ATLAS** and **CMS**



24 meters

1.7 meters

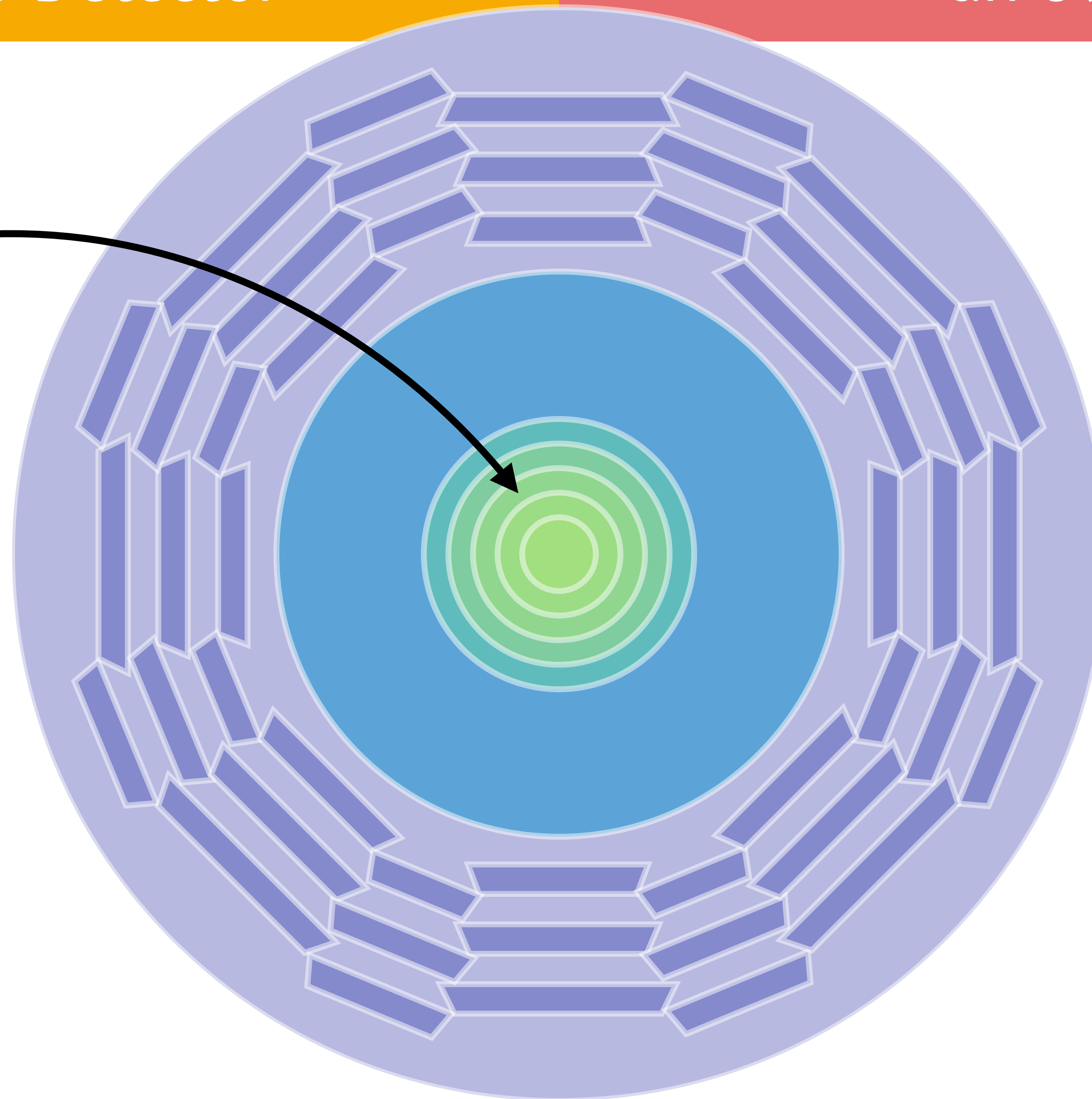


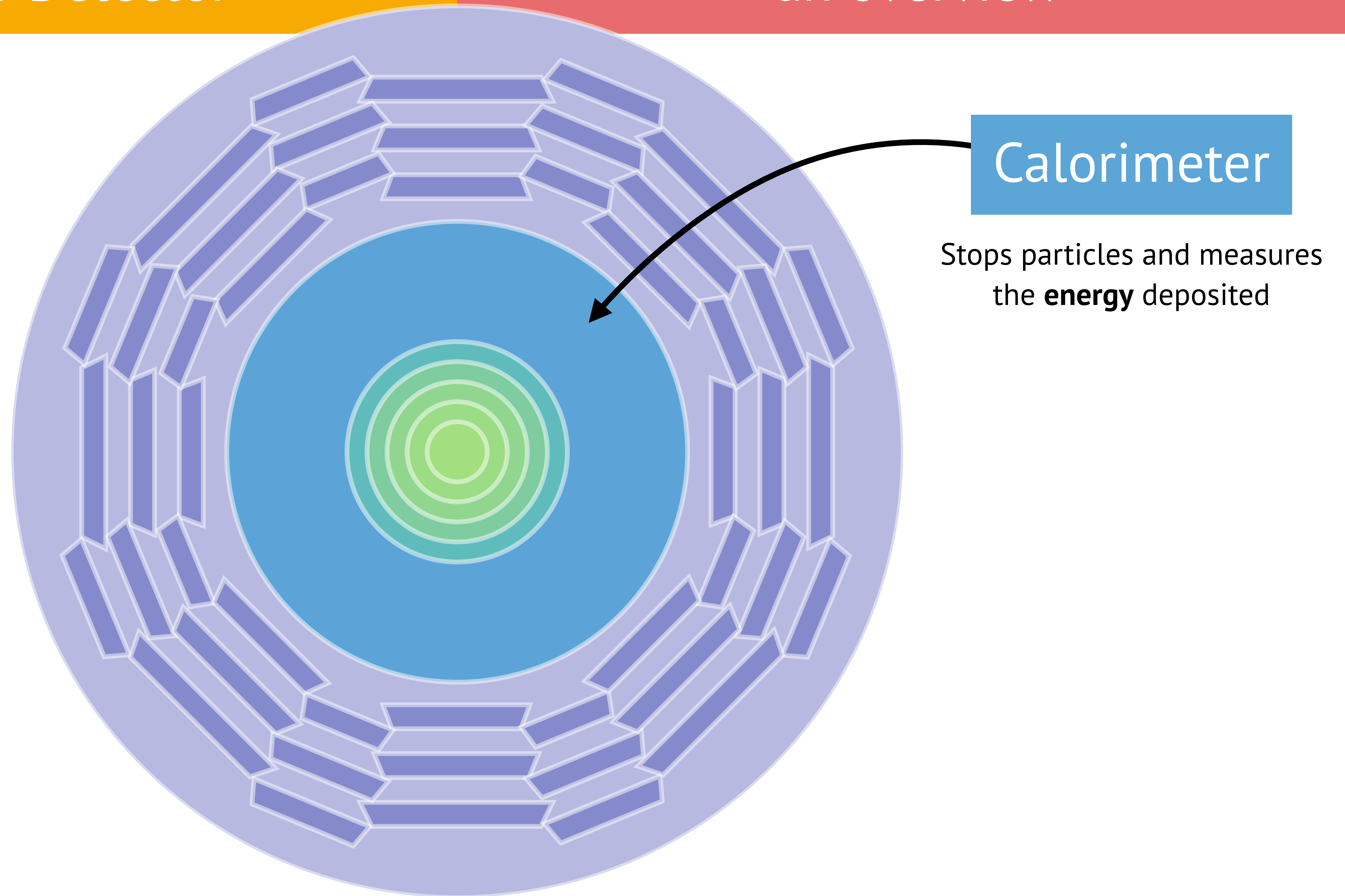


### Tracker

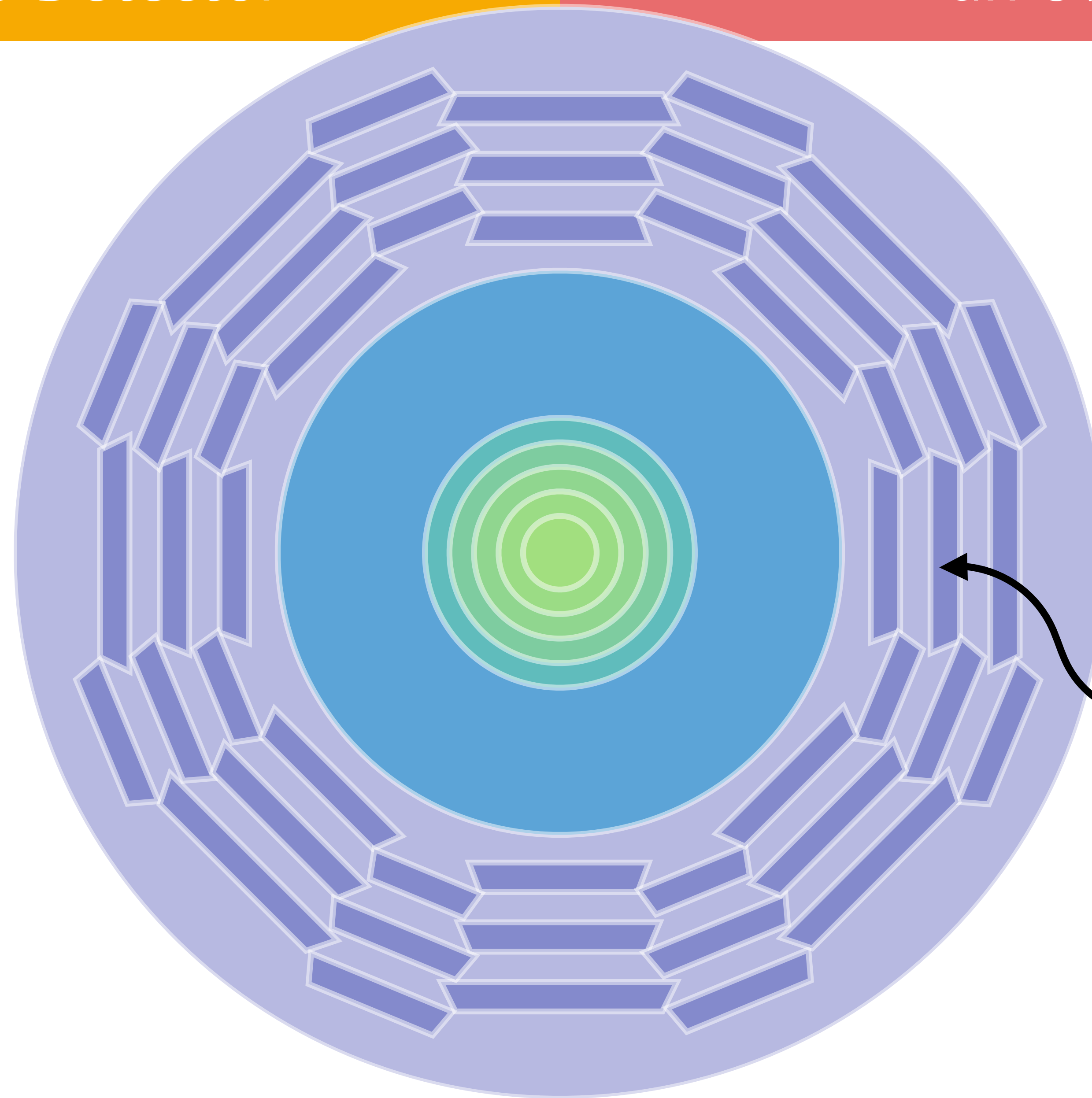
Charged particles leave tracks  
and we can get precise  
**position** measurements

Magnetic field means we  
can also measure **momentum**







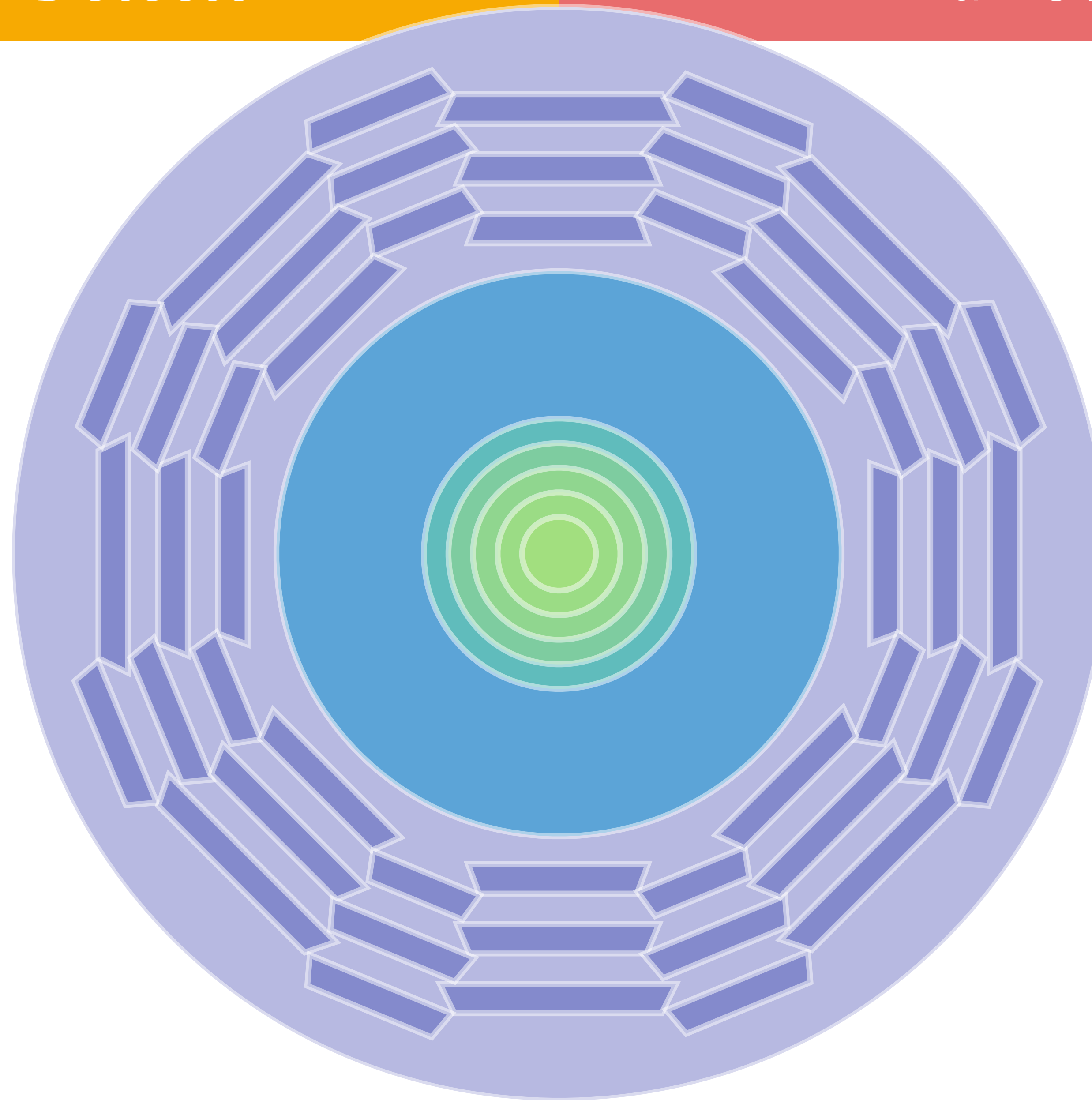
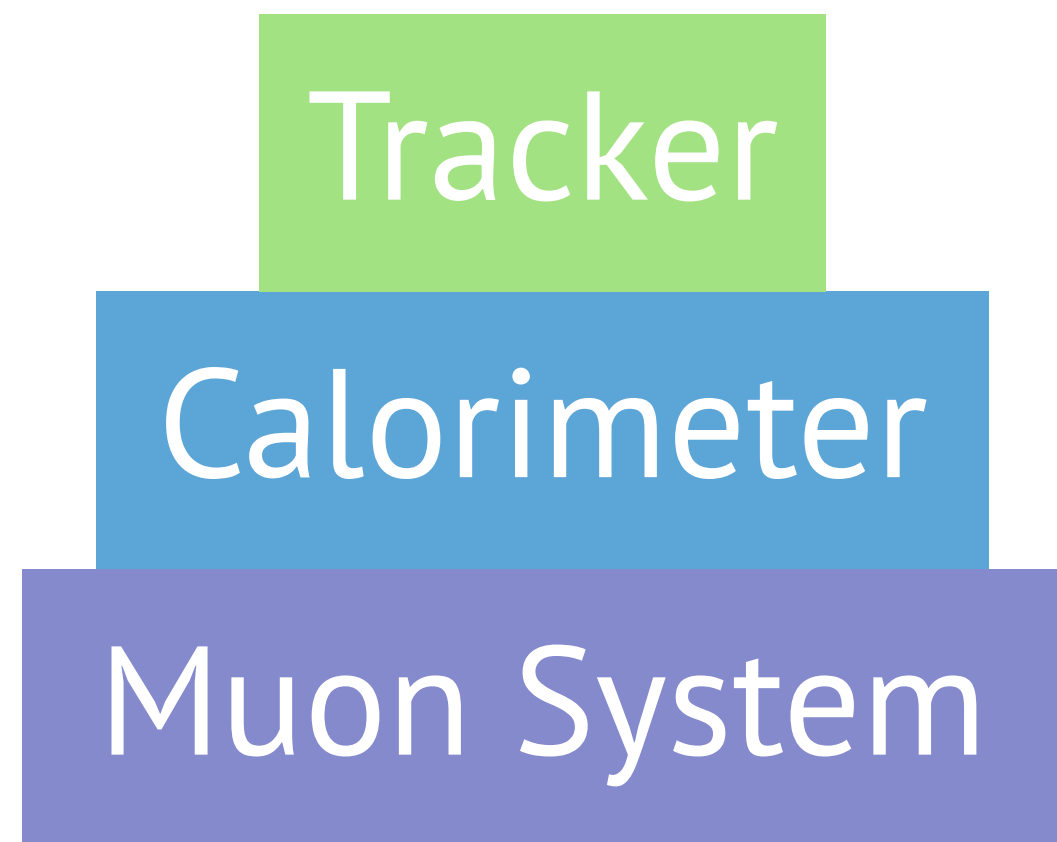


### Muon System

Muons aren't stopped by the calorimeter

Muon system provides signatures similar to the tracker, but with **very little background**

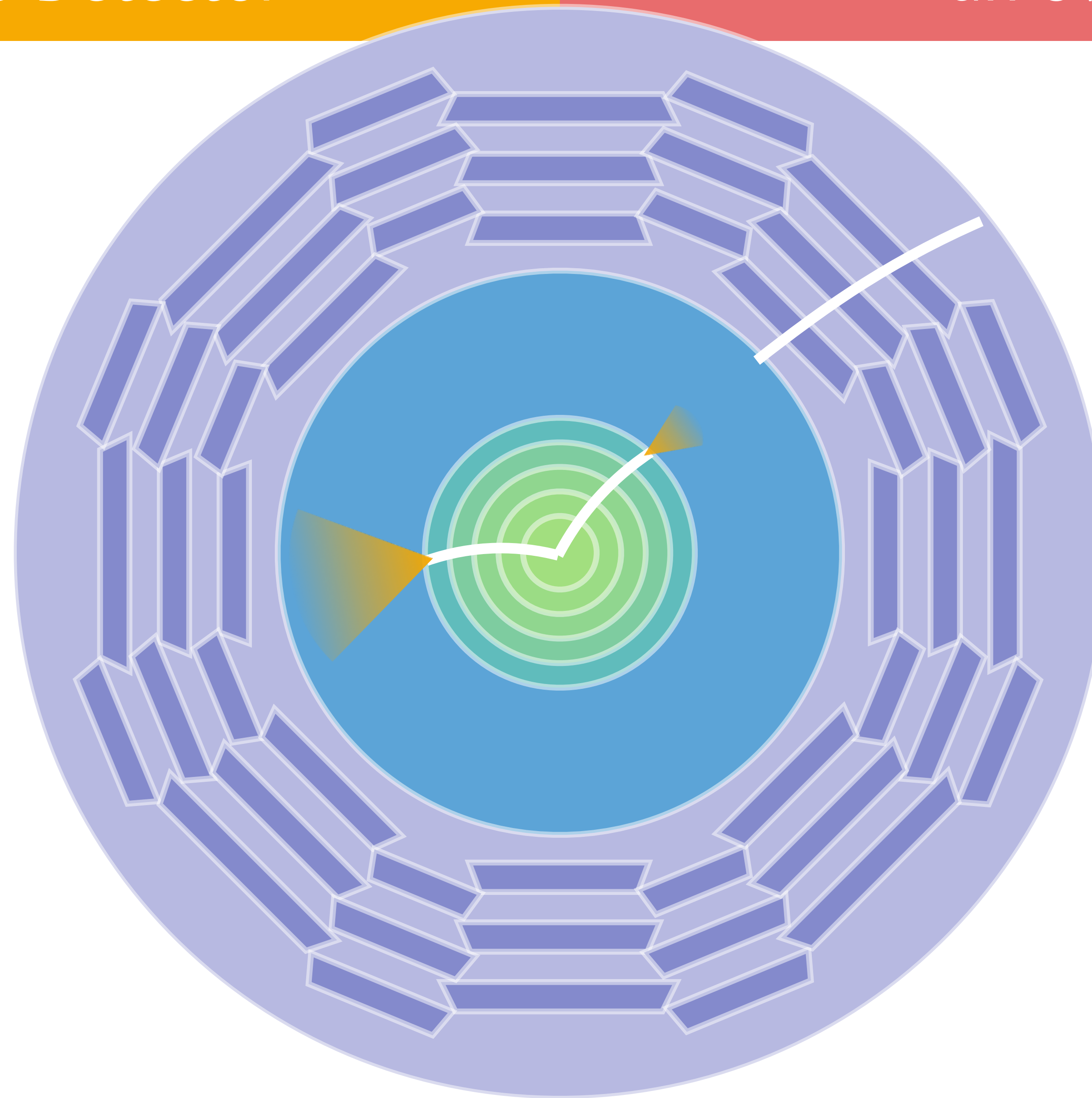




Layers of complimentary  
particle detection

Designed so we can  
**identify** a particle  
traveling through all of them

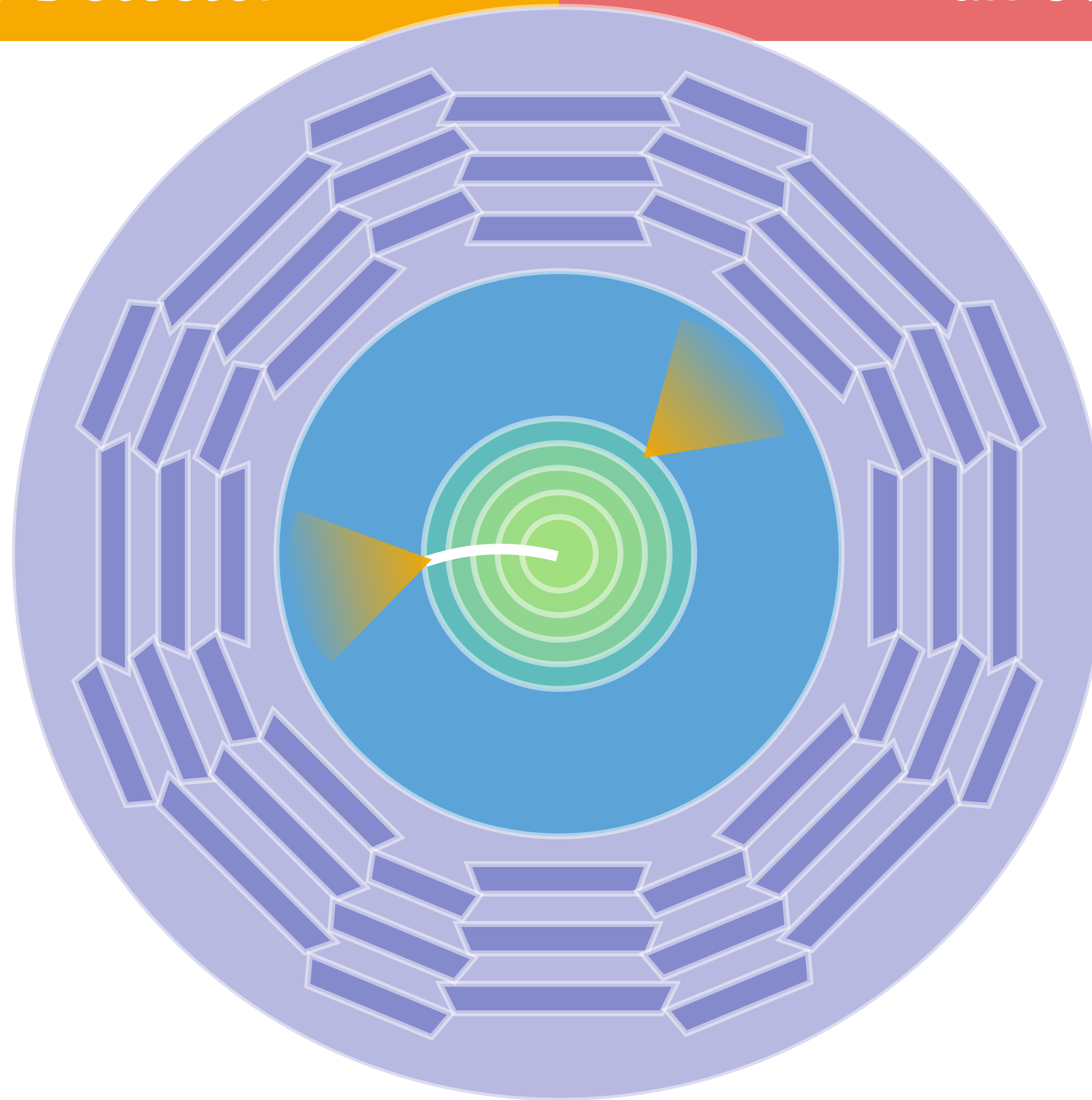




So we can tell an  
**electron** from a **muon**

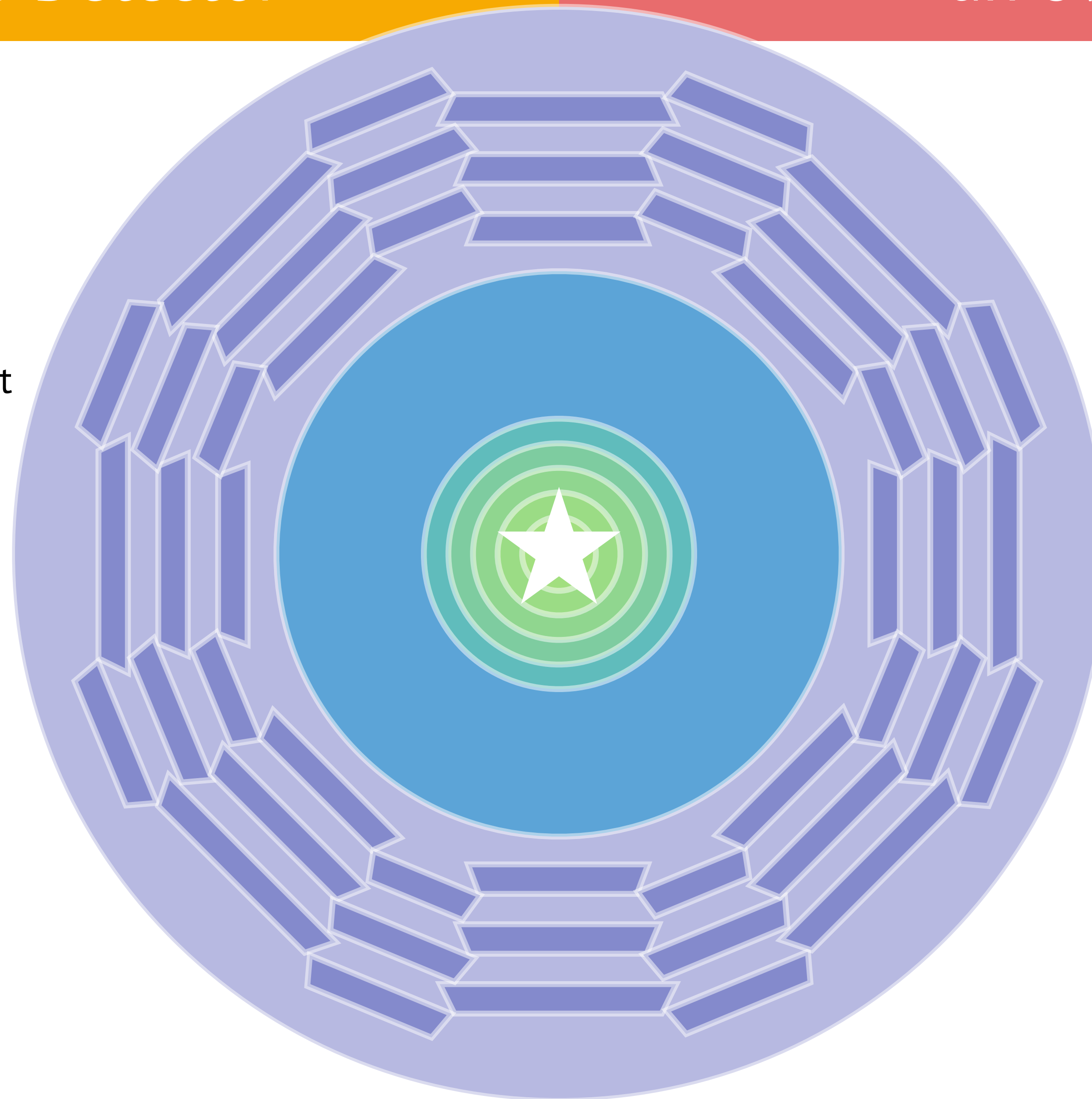
same inner detector tracks  
**electron** leaves a large deposit  
in the calorimeter  
**muon** comes with a track  
in the muon system





An **electron** and a **photon** look the same in the calorimeter, but only the electron leaves a track in the **tracker**



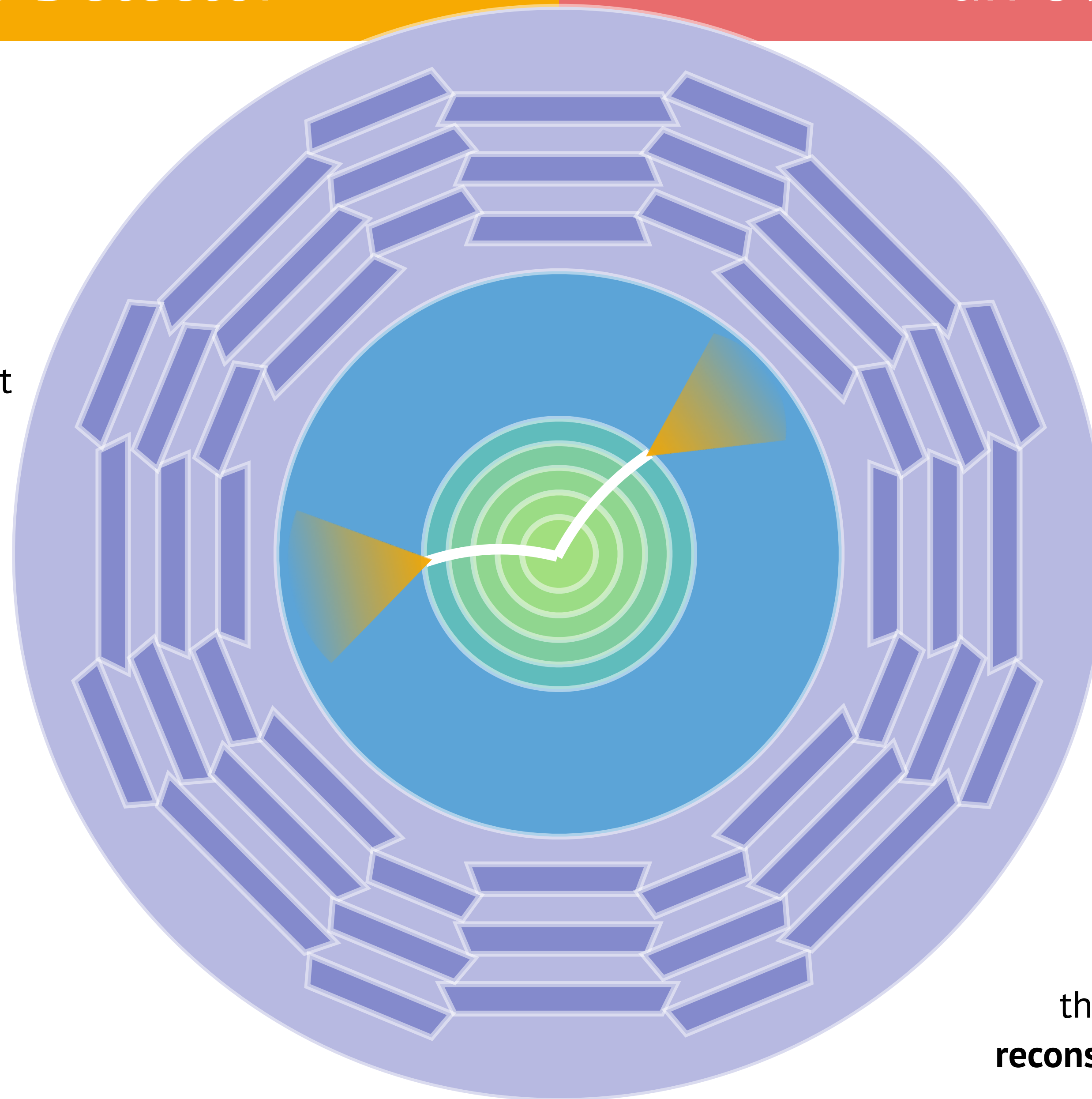


Success of particle ID is dependent  
on two assumptions

your particles **start at the center**  
of the detector and travel out

the particles you detect are  
**standard model** particles



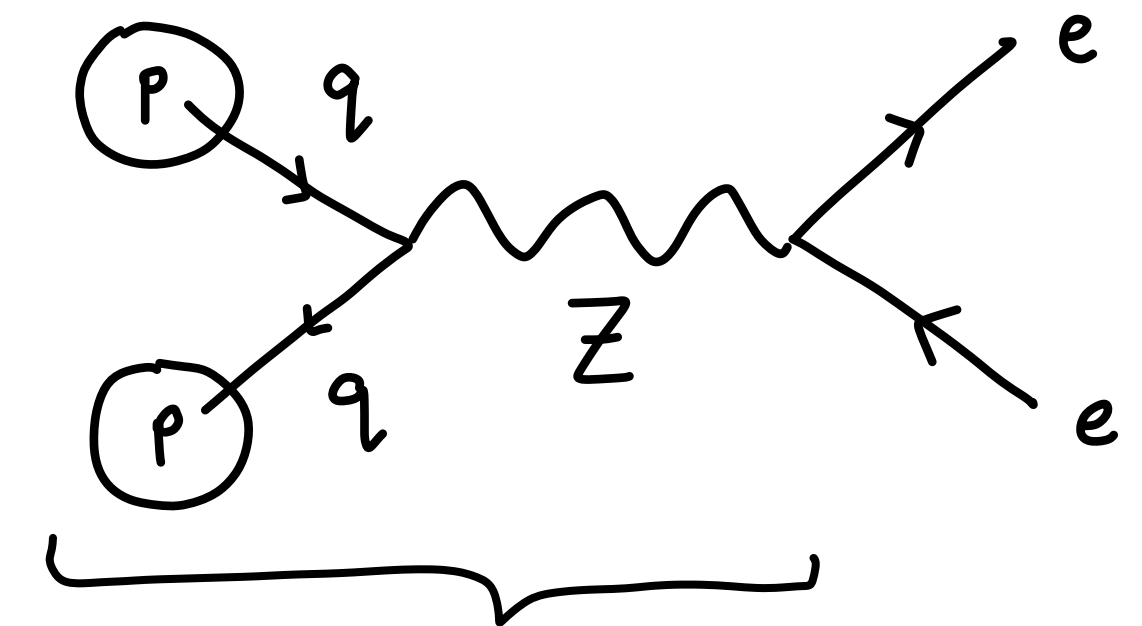


Success of particle ID is dependent  
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Typically true!

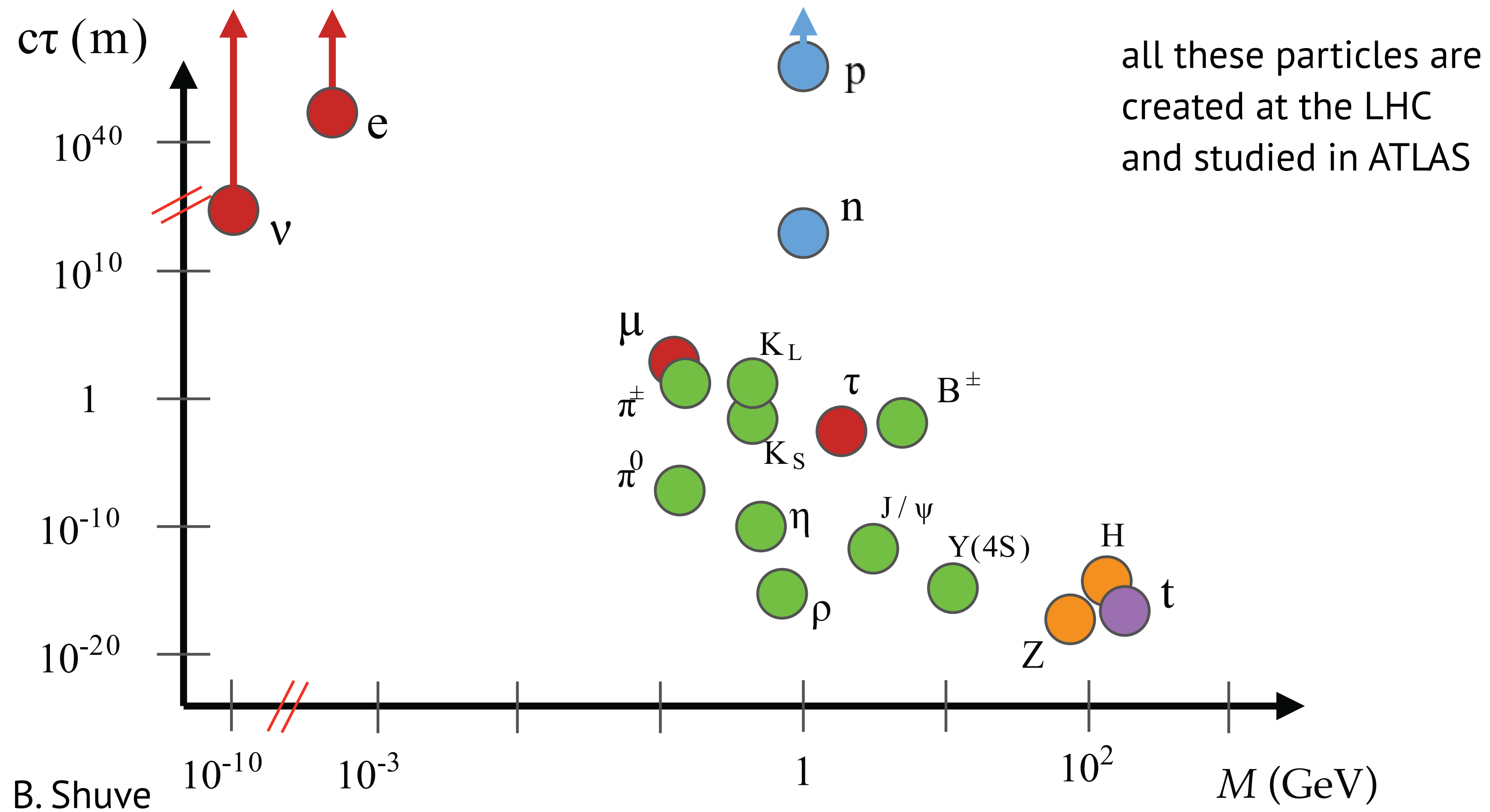


all this happens in  
 $\sim 10^{-25}$  seconds

detector only ever sees  
the **electrons** from the Z decay  
**reconstruct** the Z from their properties

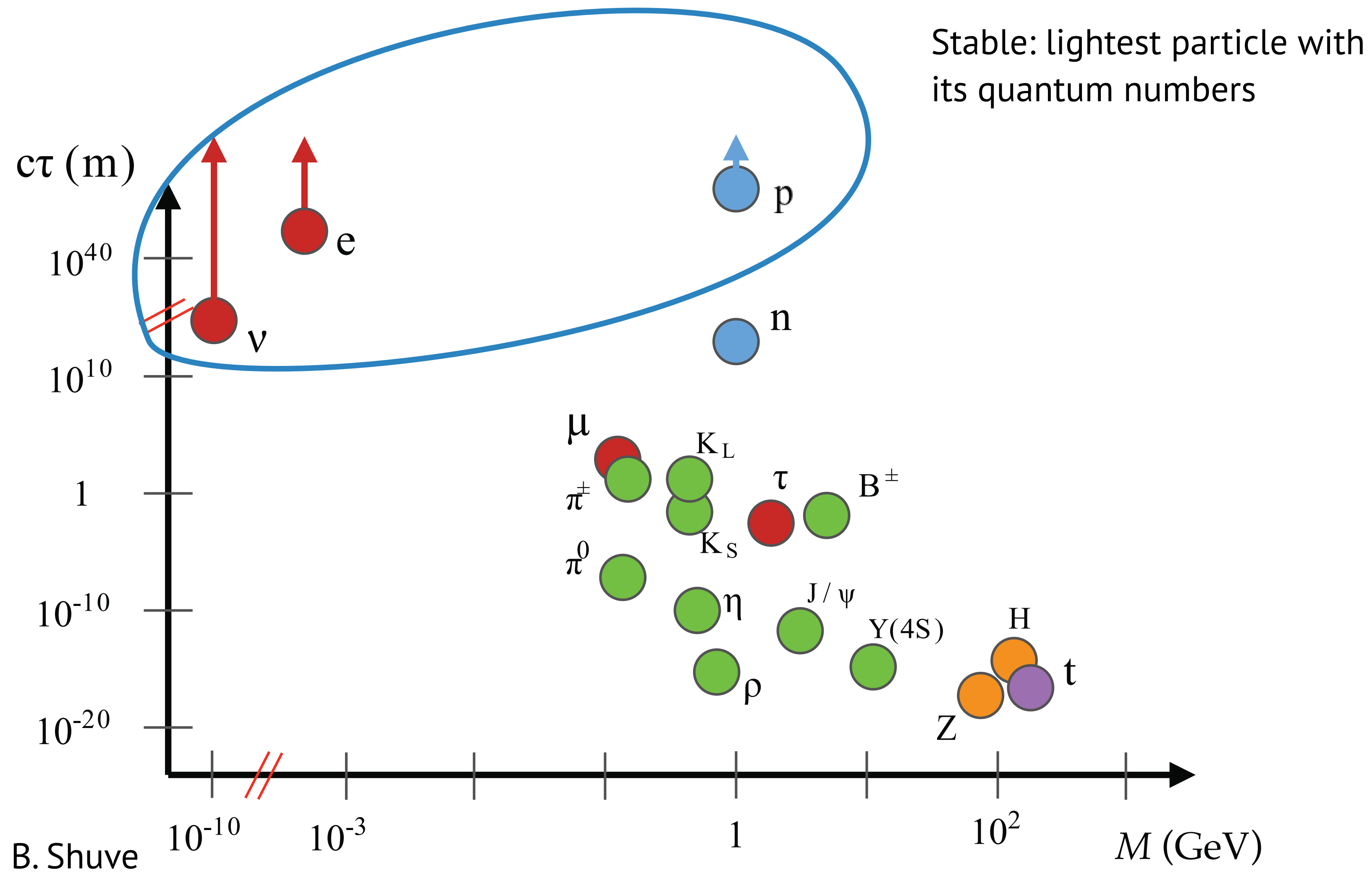


# The standard model is full of “long-lived particles”

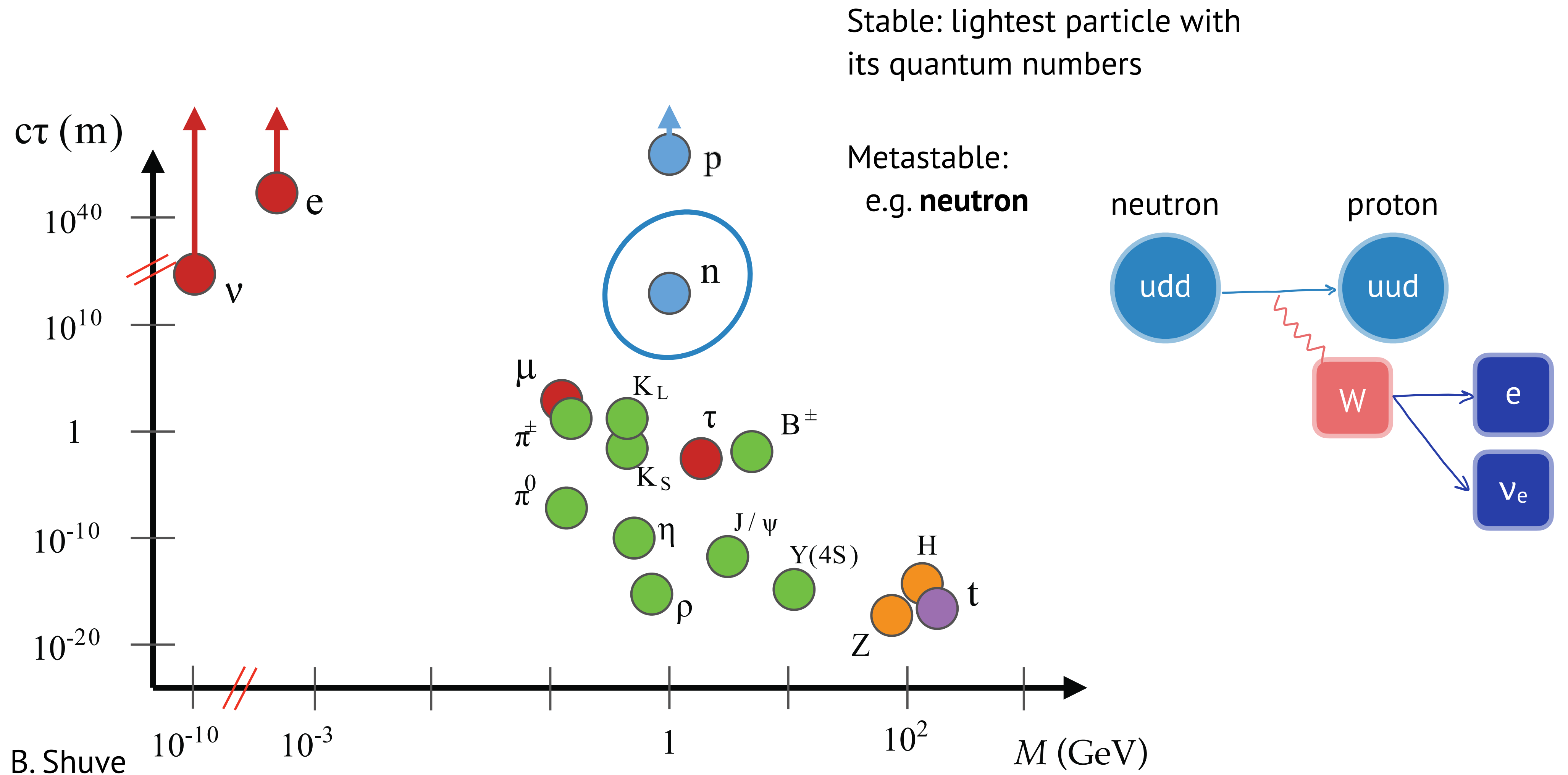




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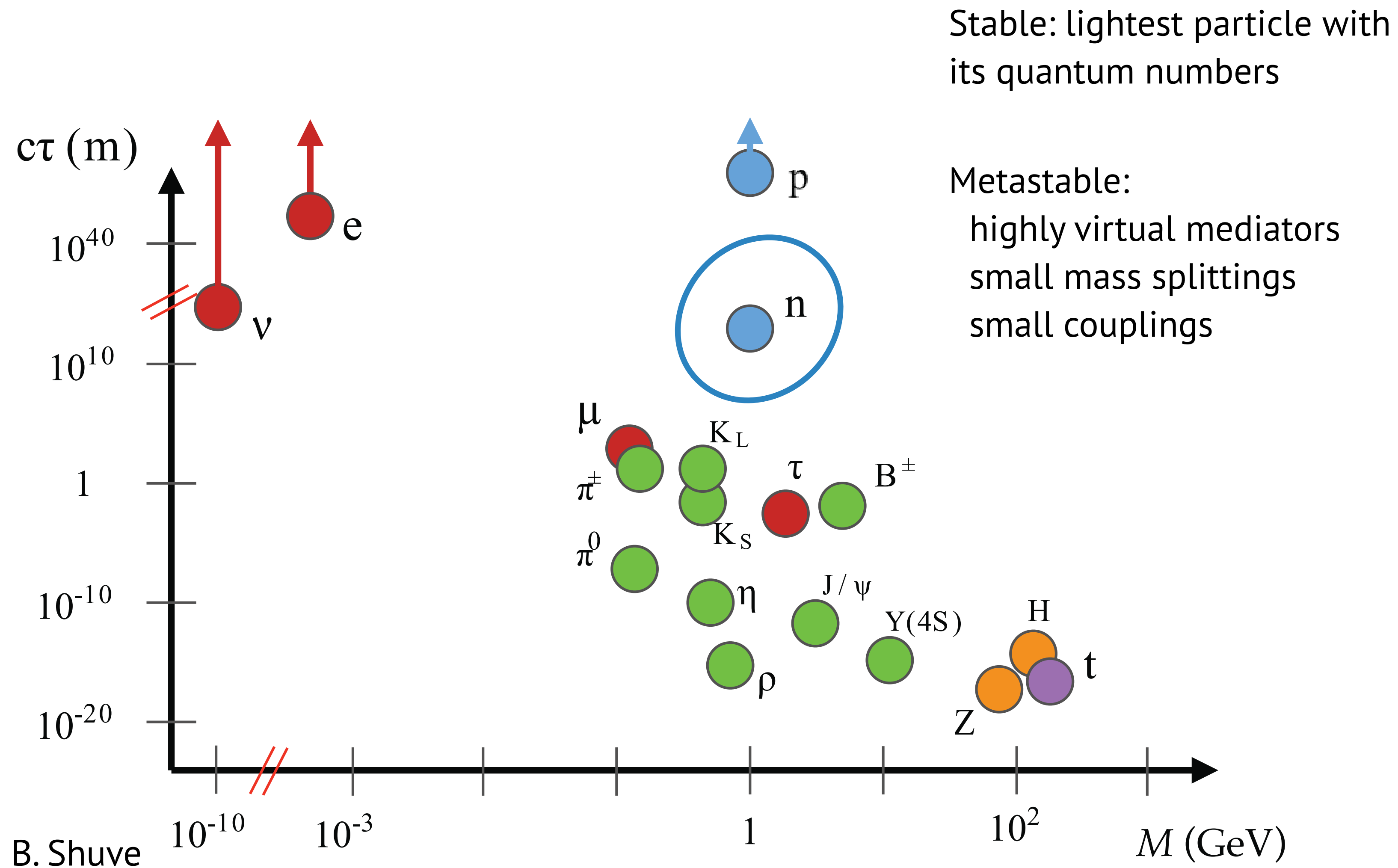


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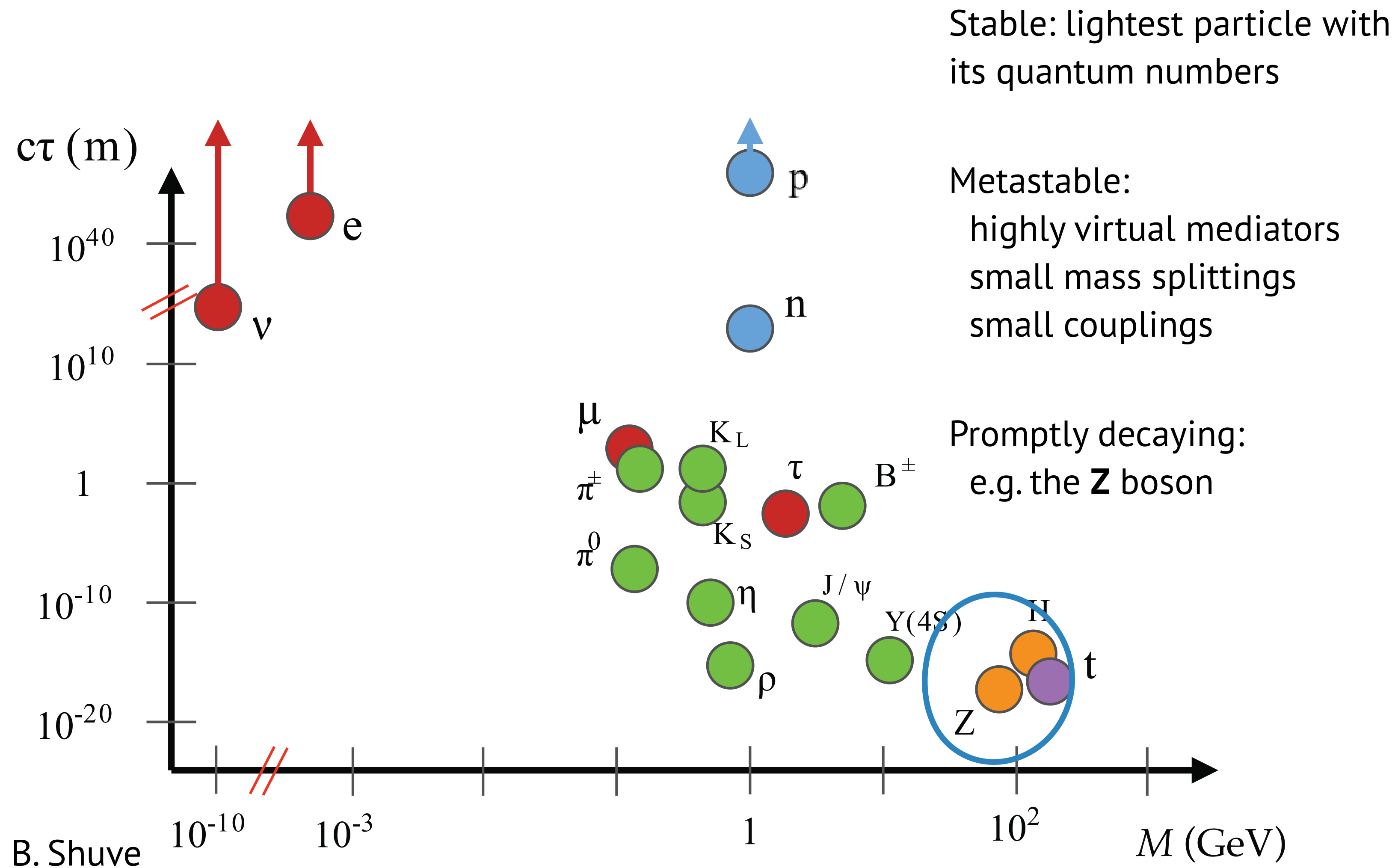




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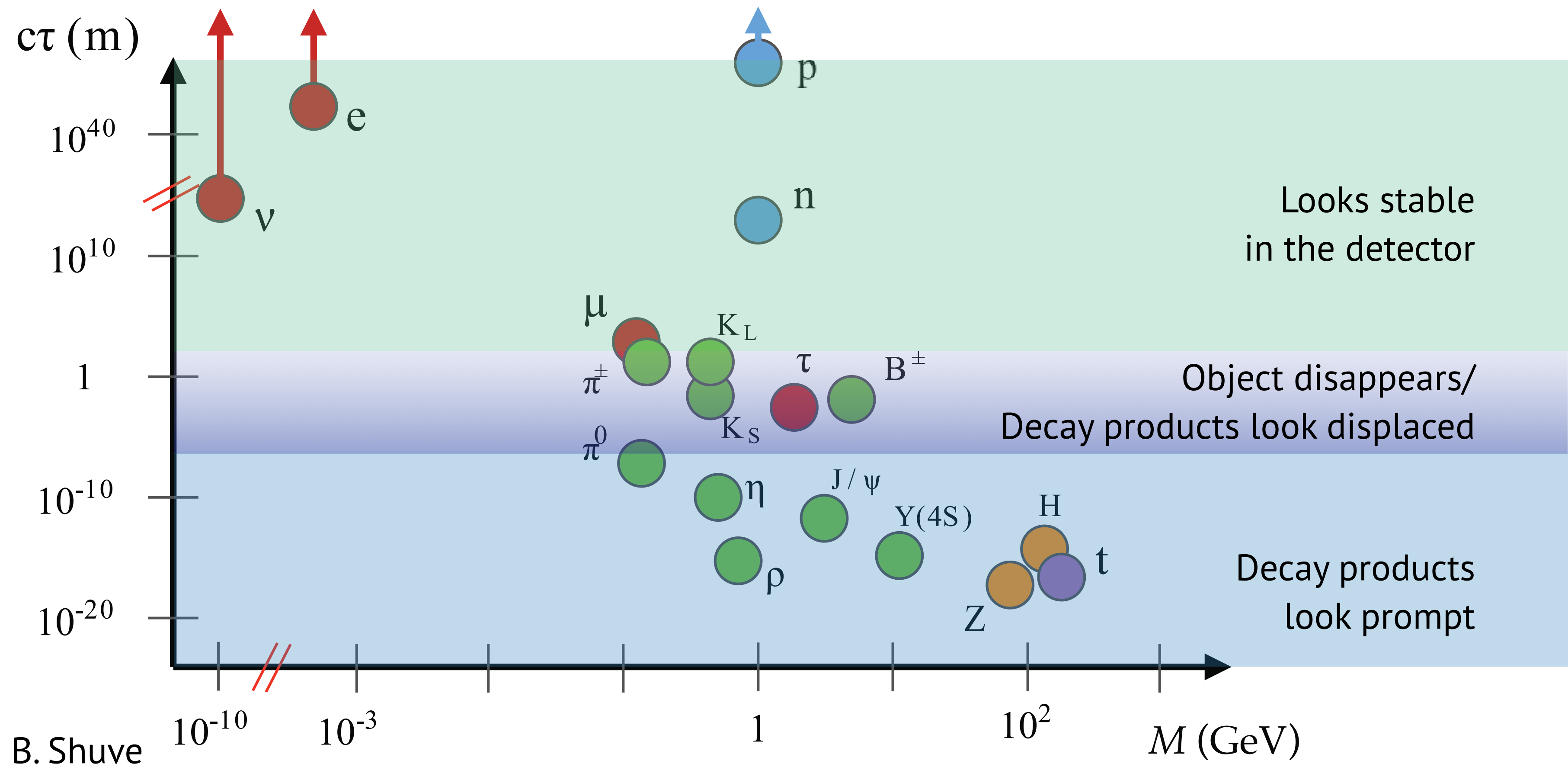


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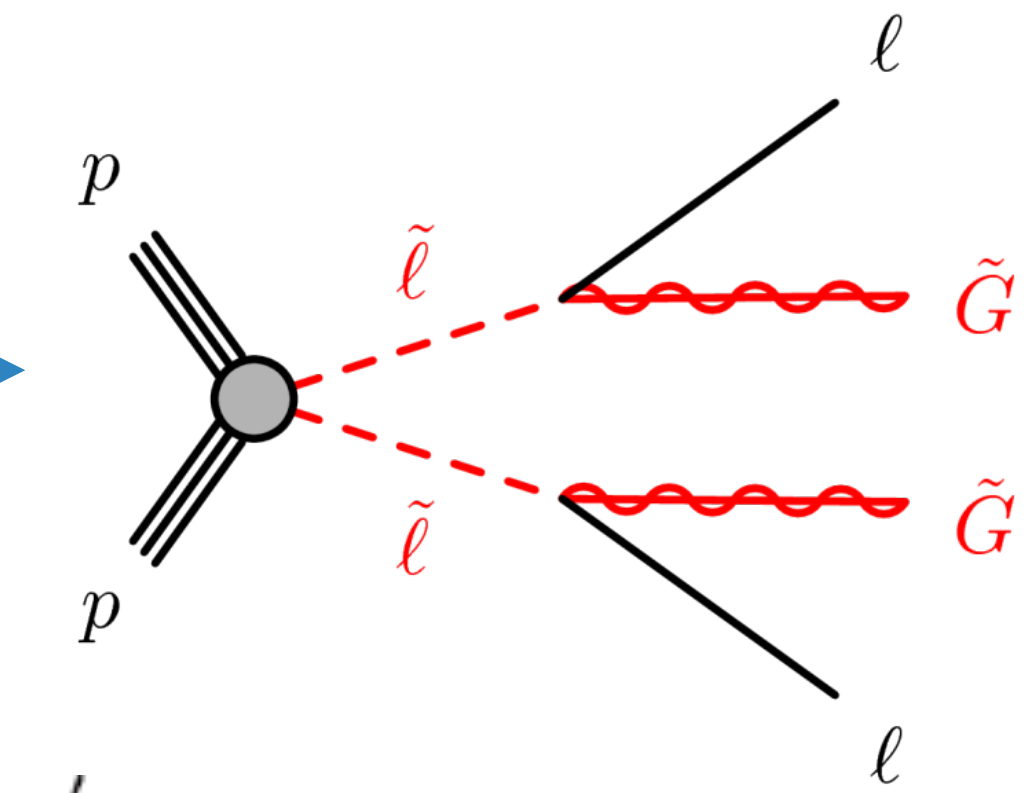
# From the experimental point of view...



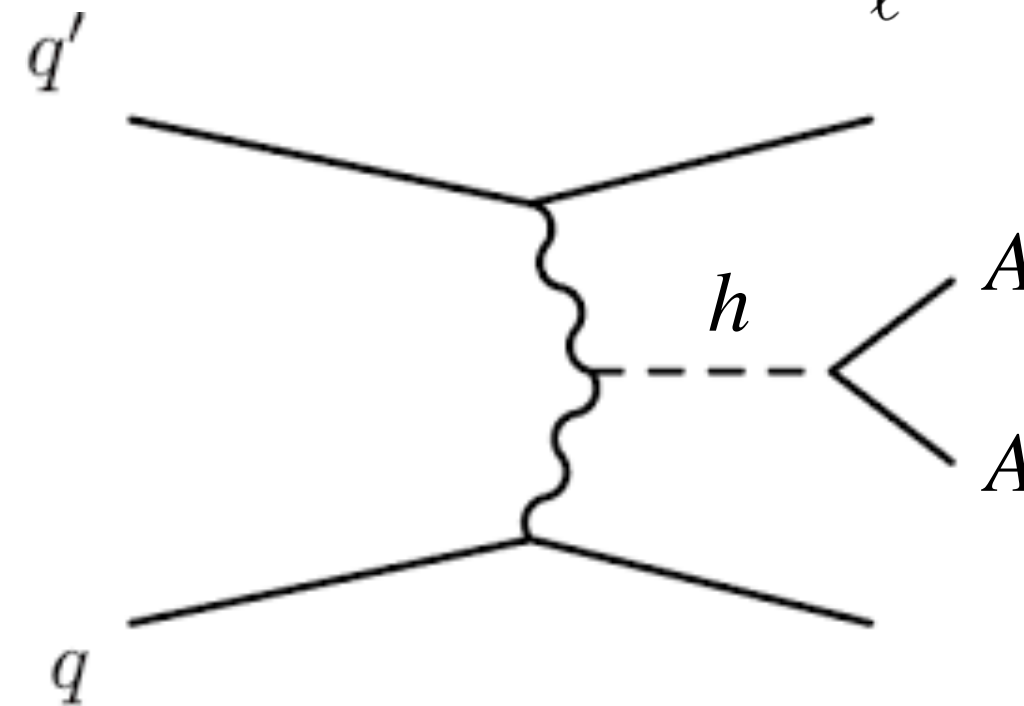
No reason to think there are only Standard Model long-lived particles...

Beyond-the-Standard-Model particles can be long-lived for the same reasons (small mass splittings, small couplings, virtual mediators)

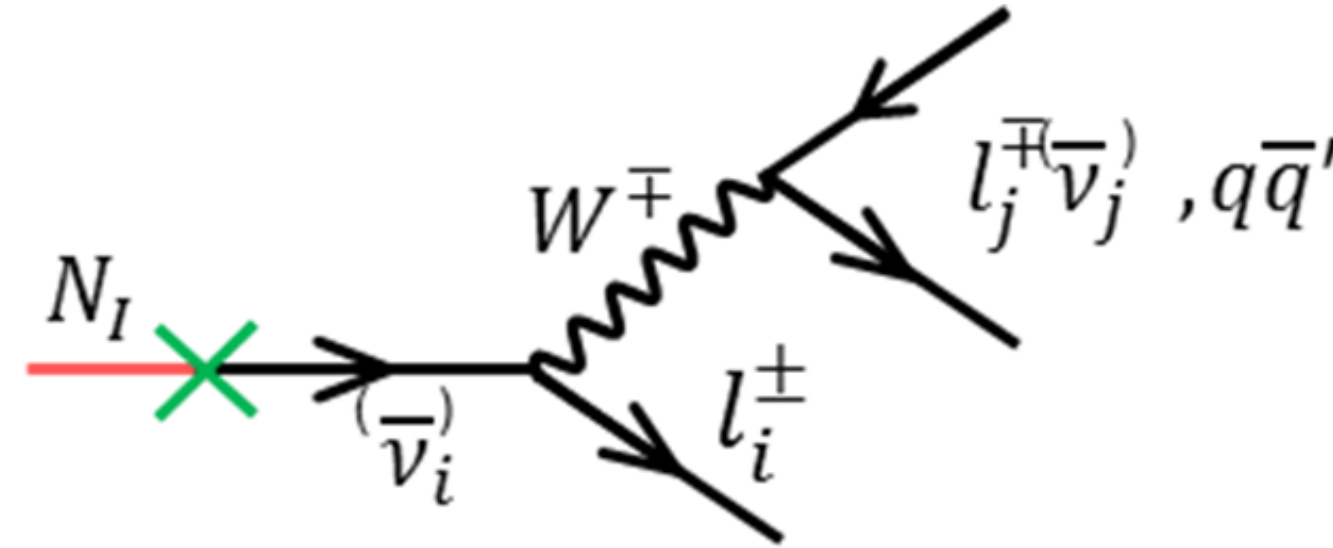
In Supersymmetry...



From hidden sectors...



Right-handed neutrinos...



Any **dark matter candidate** has to be metastable or stable!

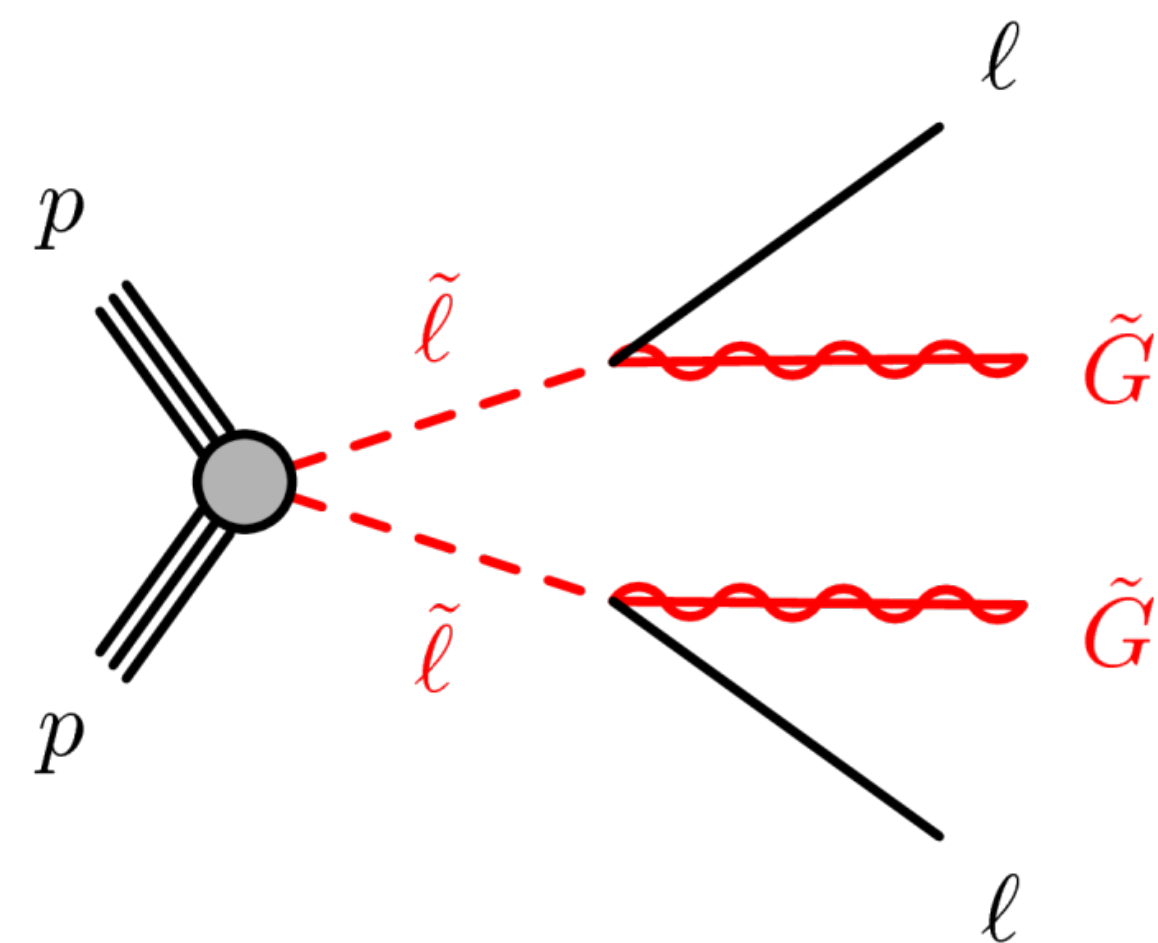


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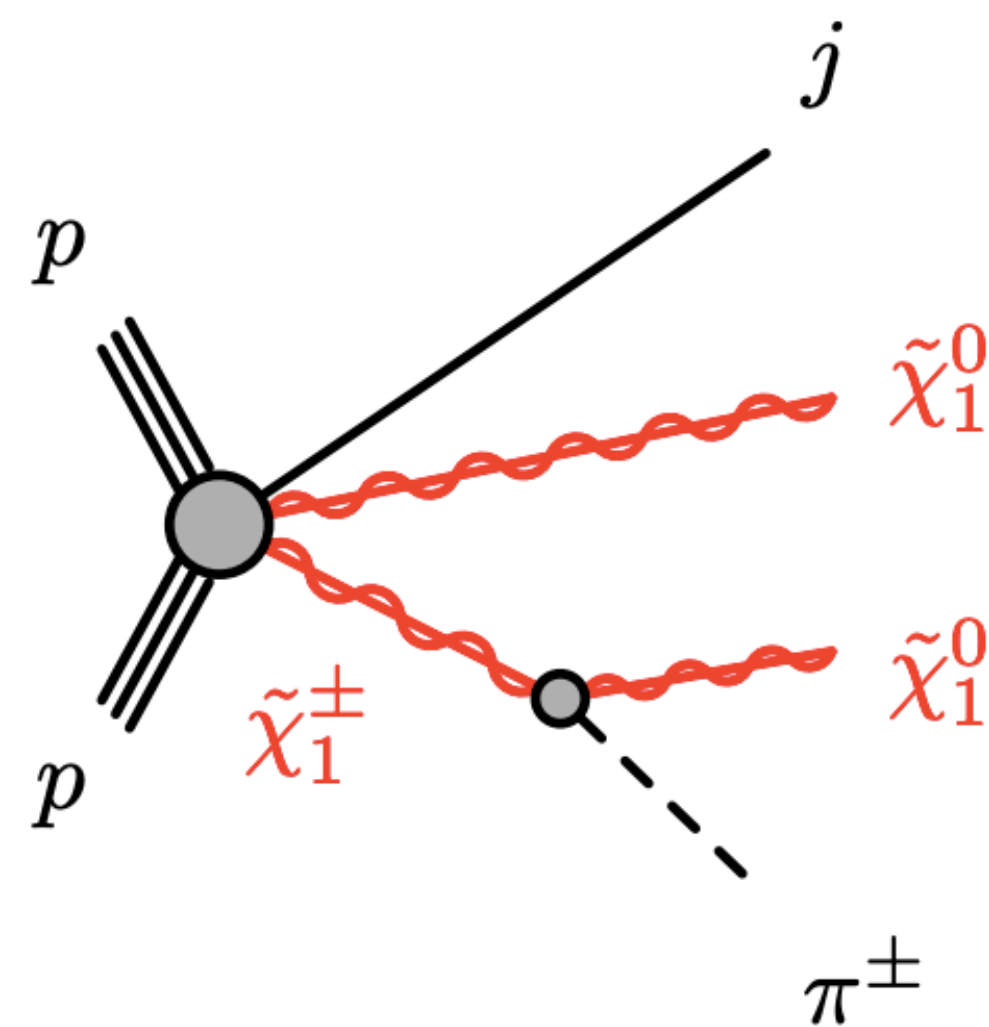
Happens in all kinds of theories of new physics, but I'll focus on one: **SUPERSYMMETRY**

Tiny Couplings



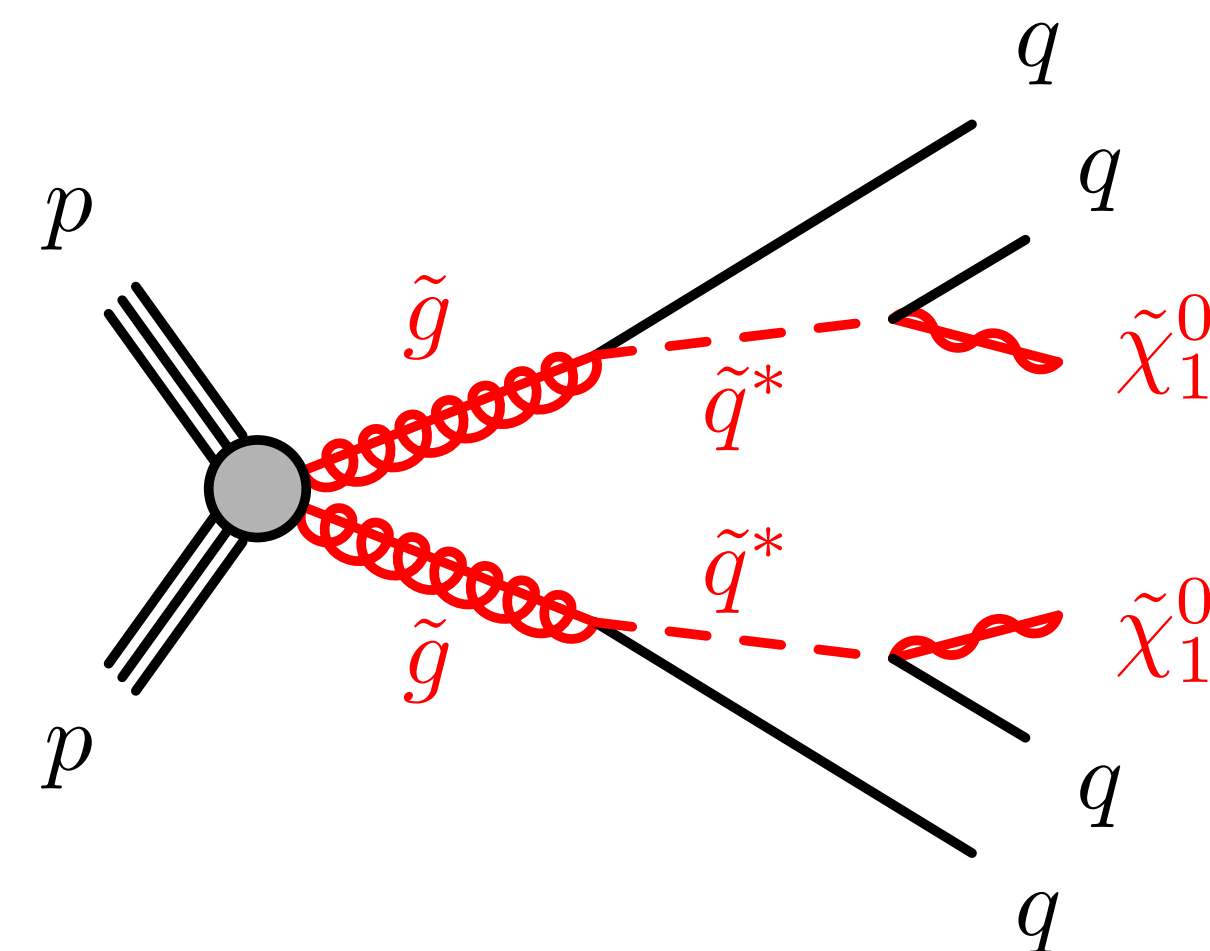
Gauge Mediated Symmetry Breaking

Small Mass Splittings



Anomaly Mediated Symmetry Breaking

Virtual Mediators



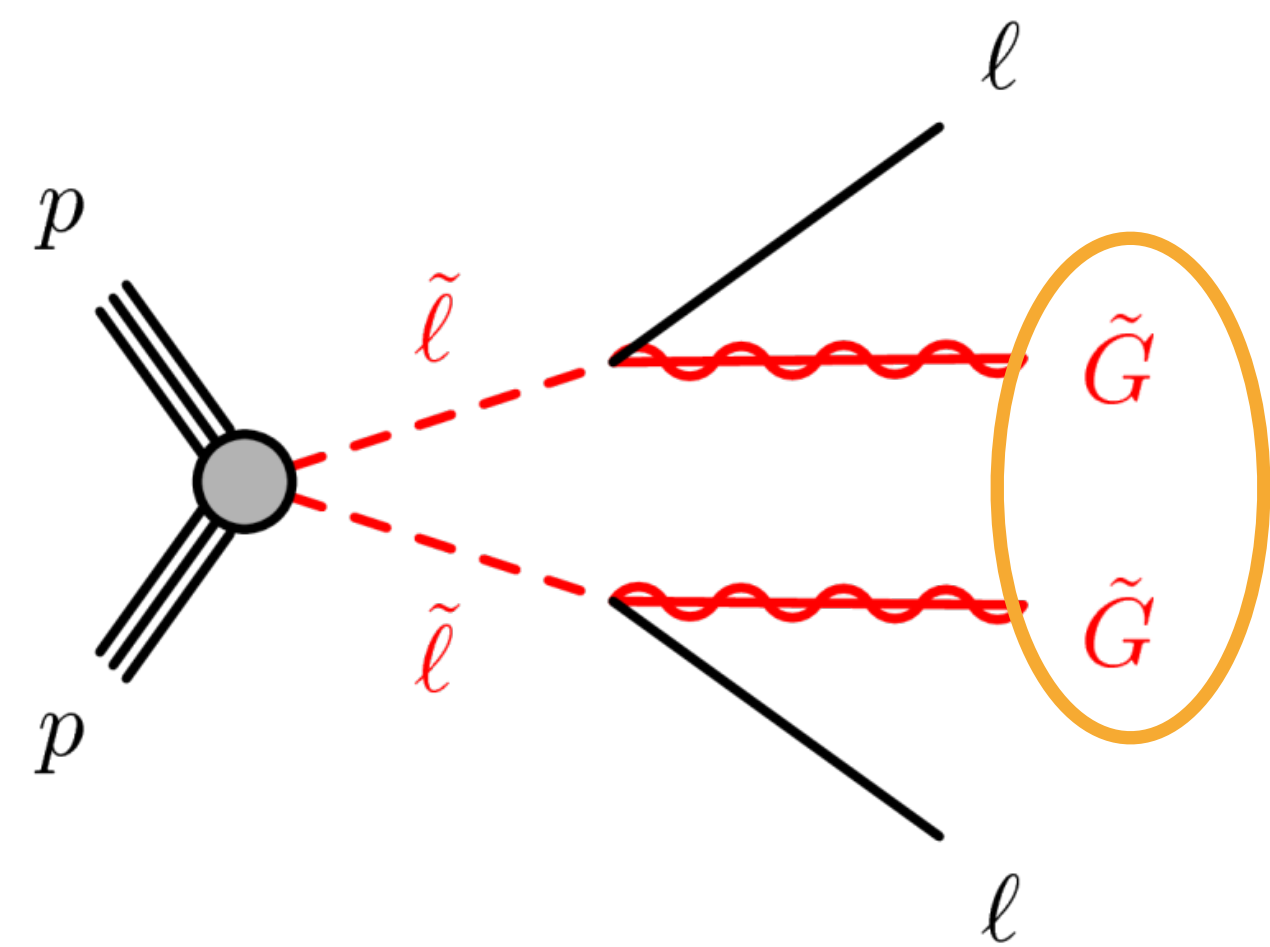
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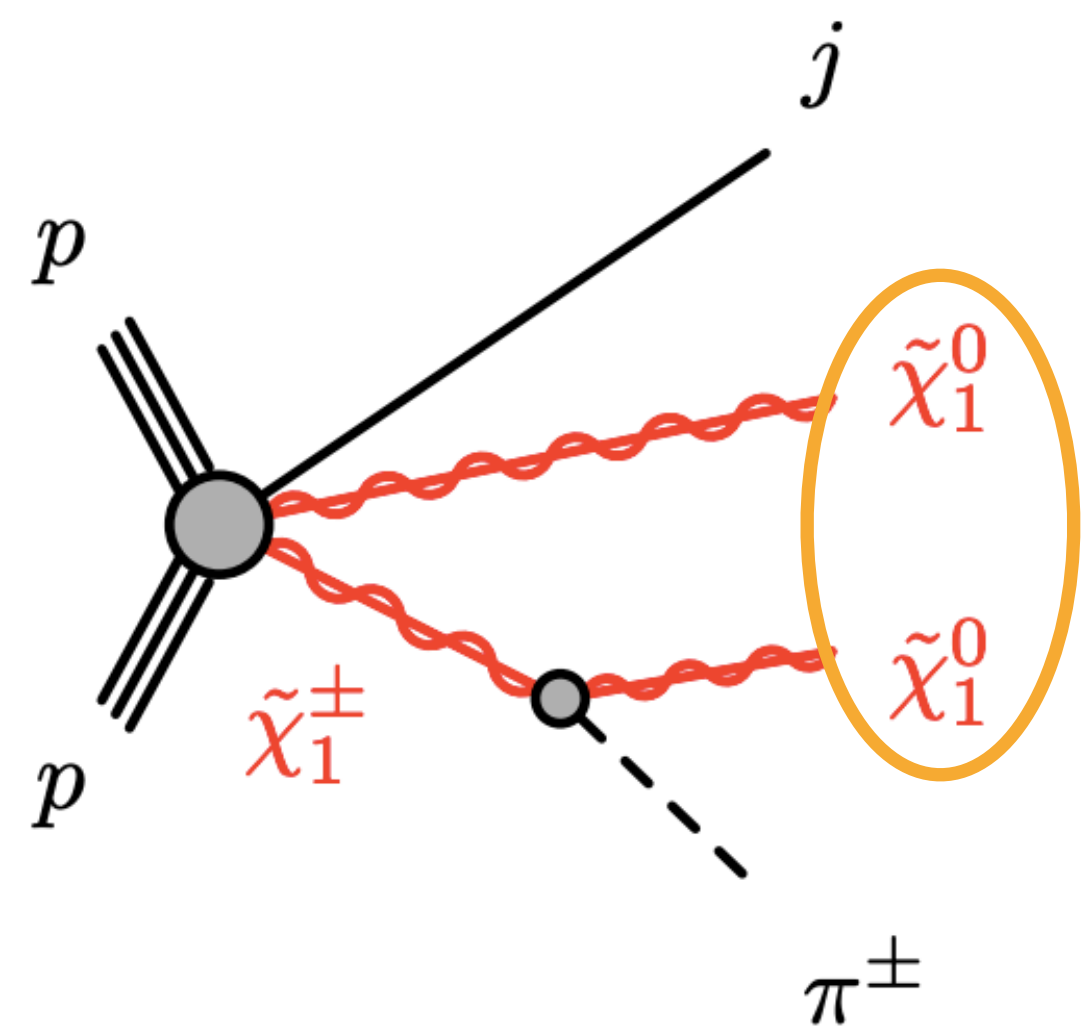
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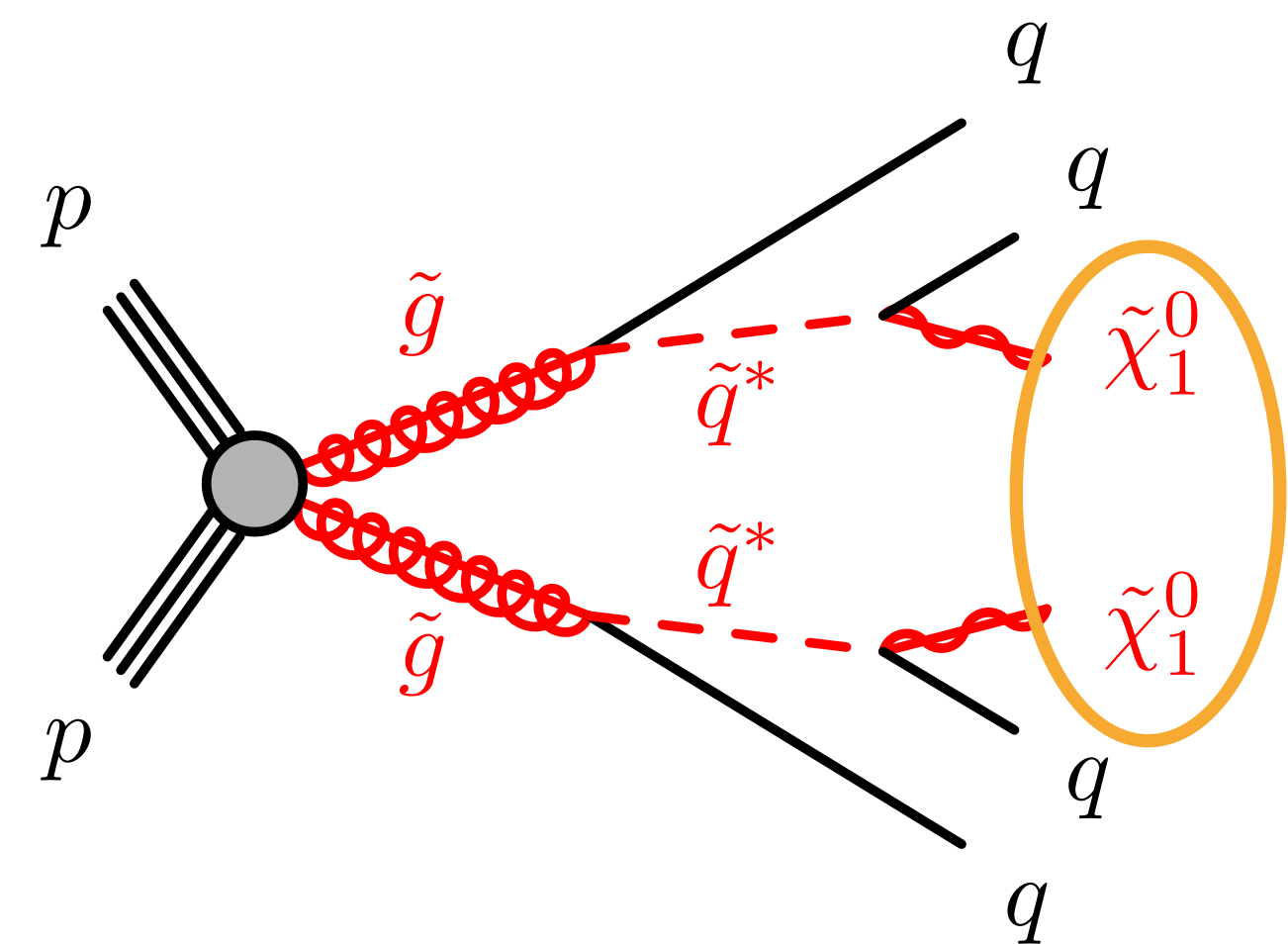
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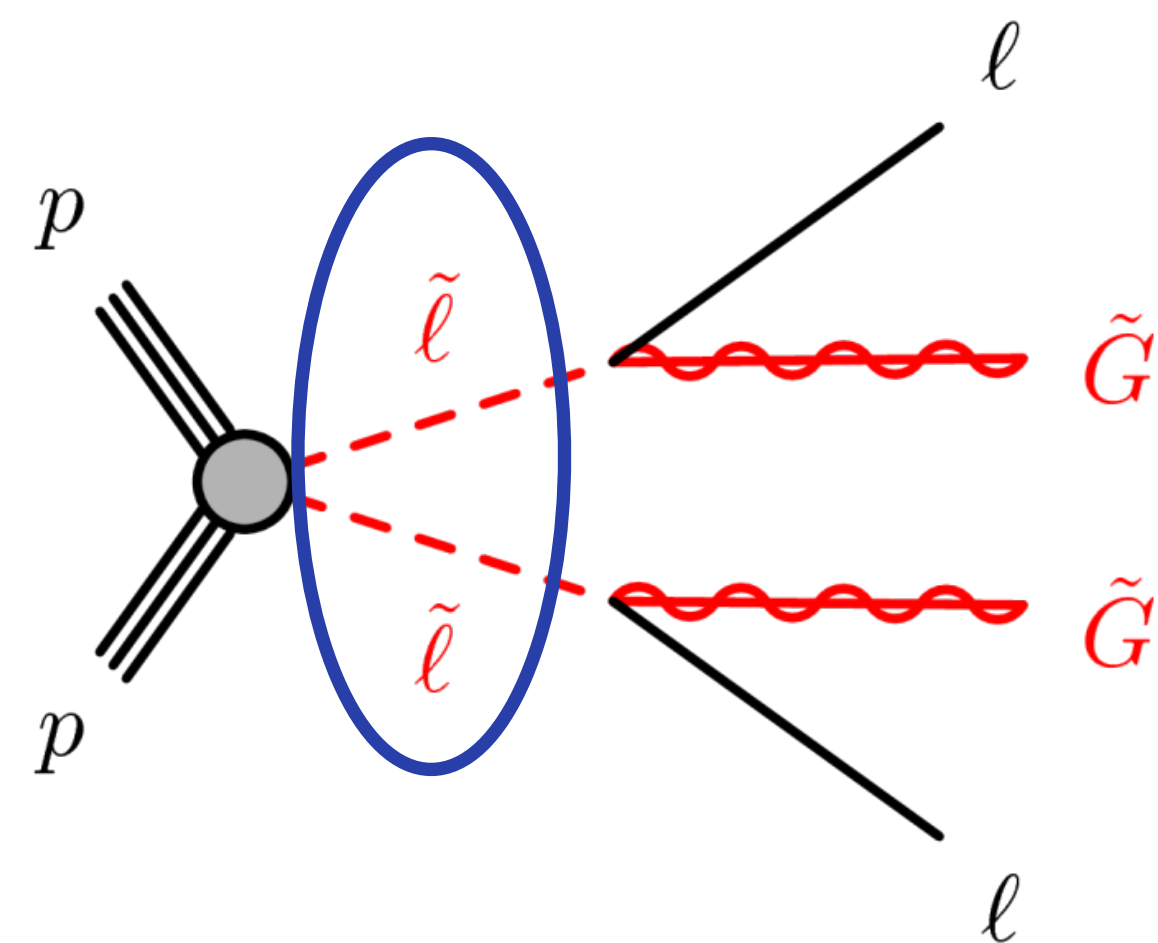


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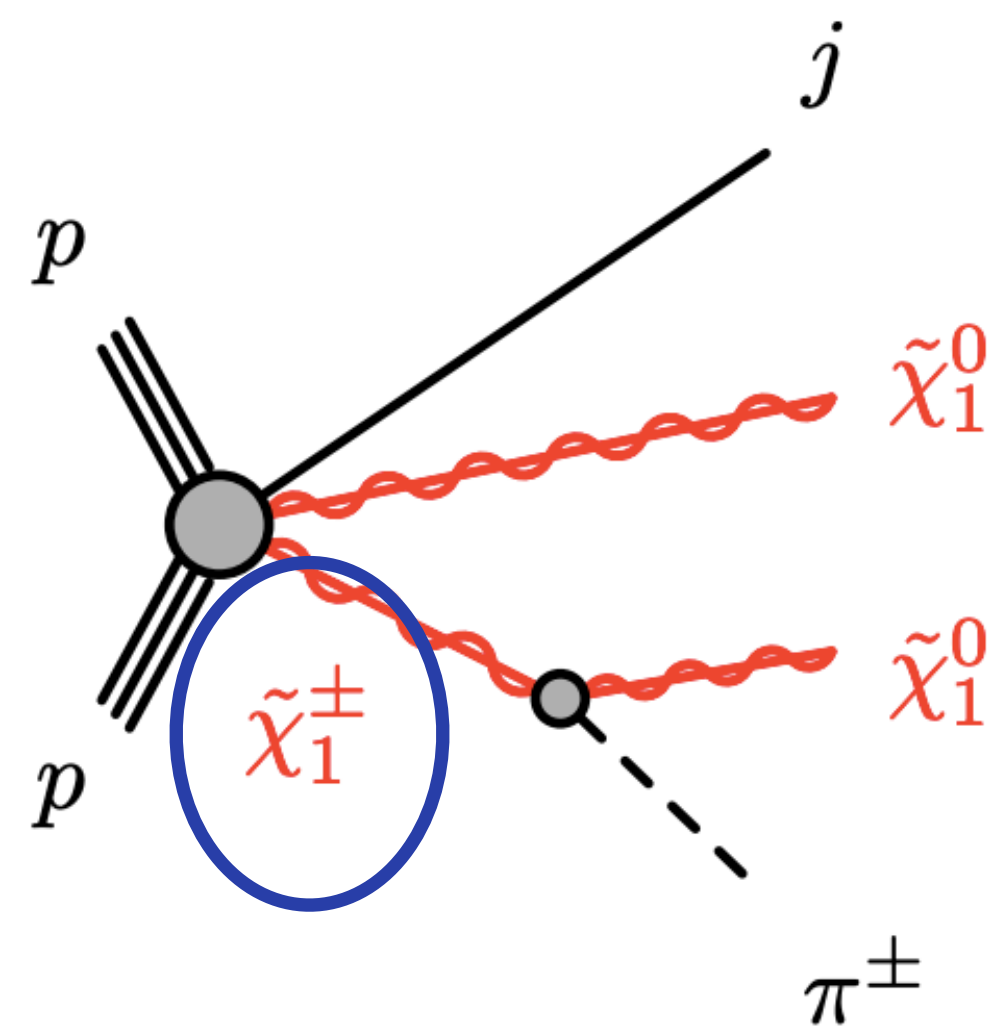
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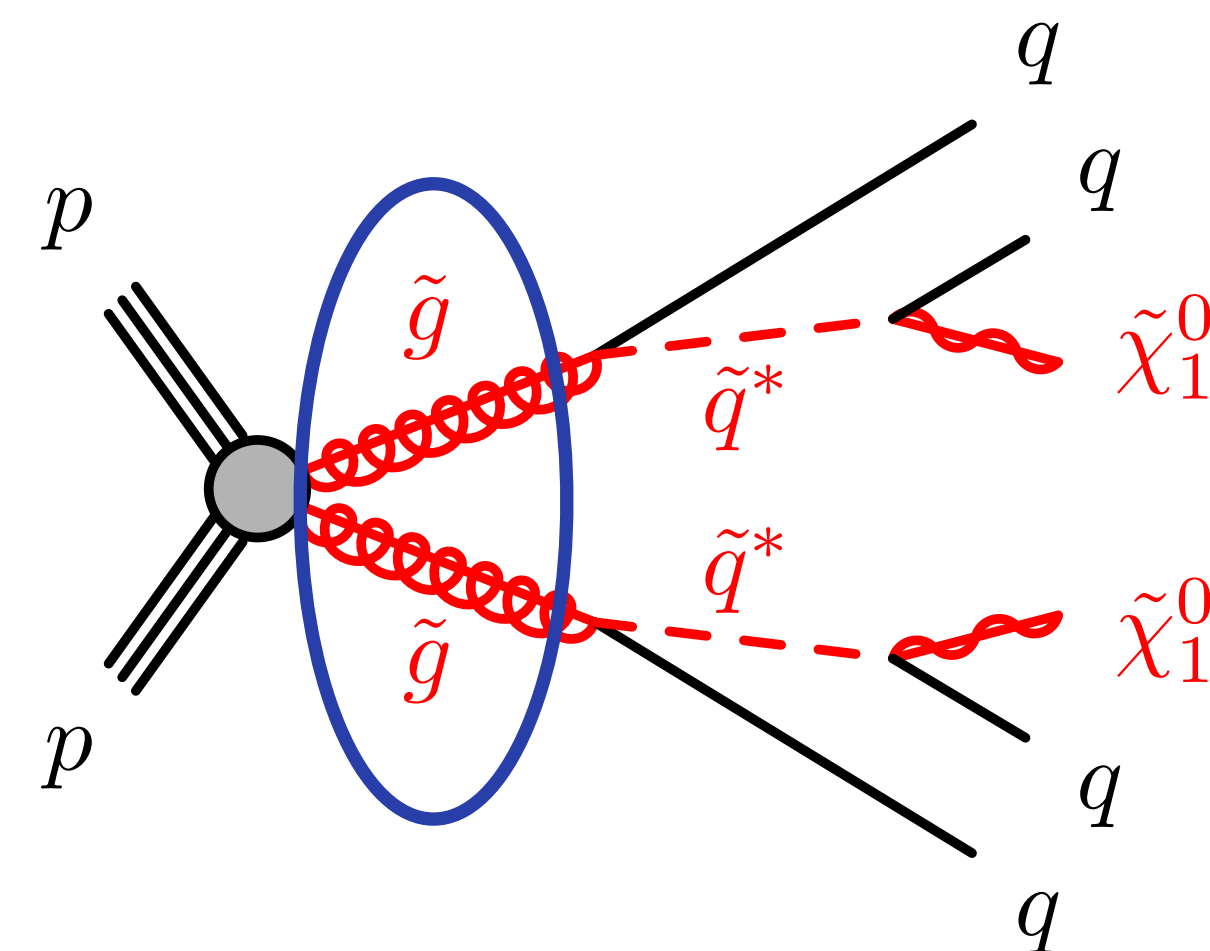
Gauge Mediated Symmetry Breaking

Small Mass Splittings



Anomaly Mediated Symmetry Breaking

Virtual Mediators



Split Supersymmetry

Happens in all kinds of theories of new physics, but I'll focus on one: **SUPERSYMMETRY**

### Why we love SUSY

Gives a solution to the **hierarchy** problem

Gives a **dark matter** candidate

Is a particularly **elegant** mathematical concept

Gives rise to many different particles with complex and ~unpredictable mass spectra  
(keeps us employed for decades)

### Why we don't love SUSY



Happens in all kinds of theories of new physics, but I'll focus on one: **SUPERSYMMETRY**

Why we love SUSY

Why we don't love SUSY

We thought it would be easier to find

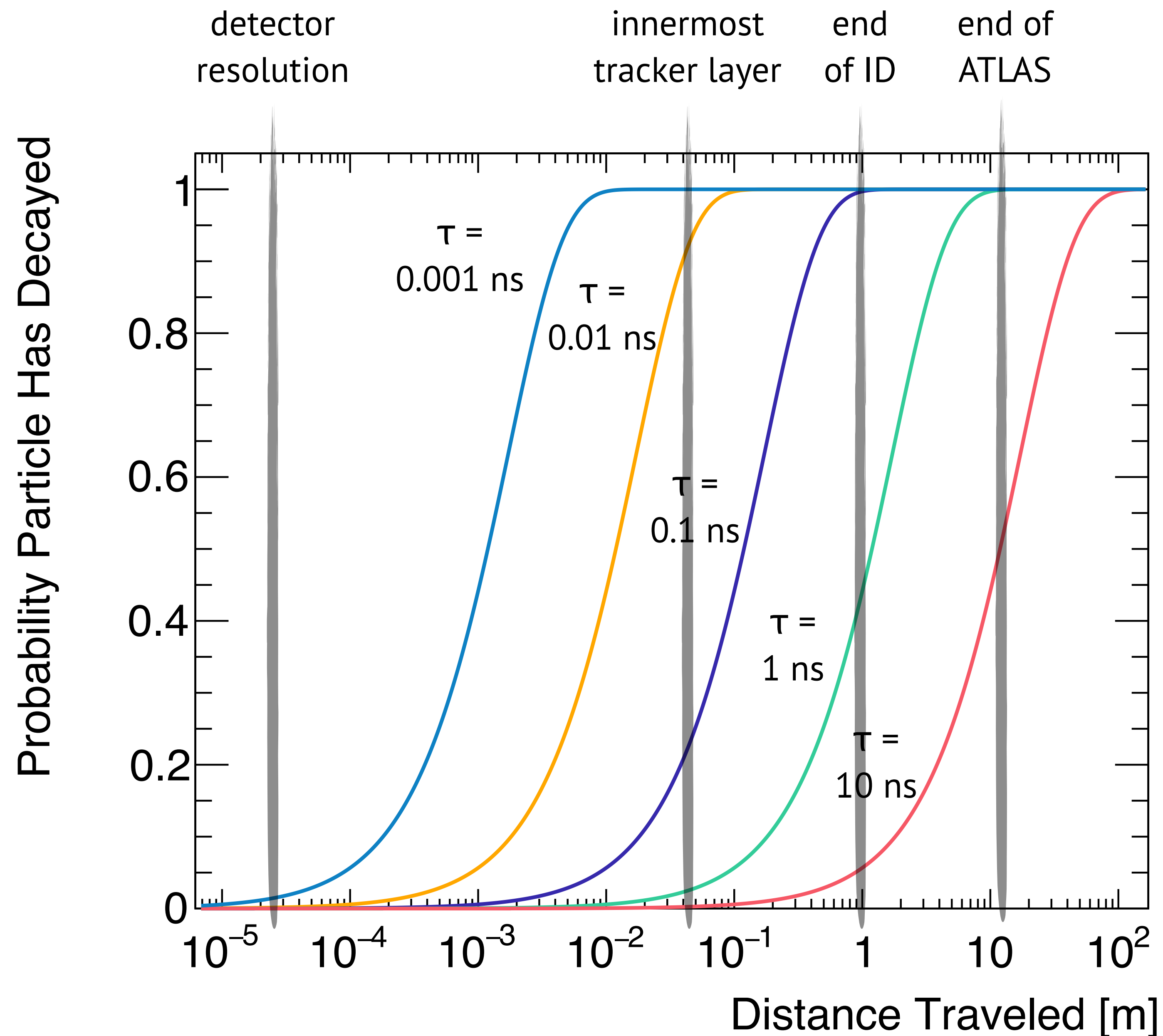
## What makes lifetime hard?

Varying lifetime creates completely different **signatures** in our detector

Many of those variations include signatures we **didn't design detectors** to reconstruct

Must have **dedicated** searches for each signature – adds a **new dimension** to the types of searches we need to cover all scenarios!

200 GeV particle  
energy: 500 GeV





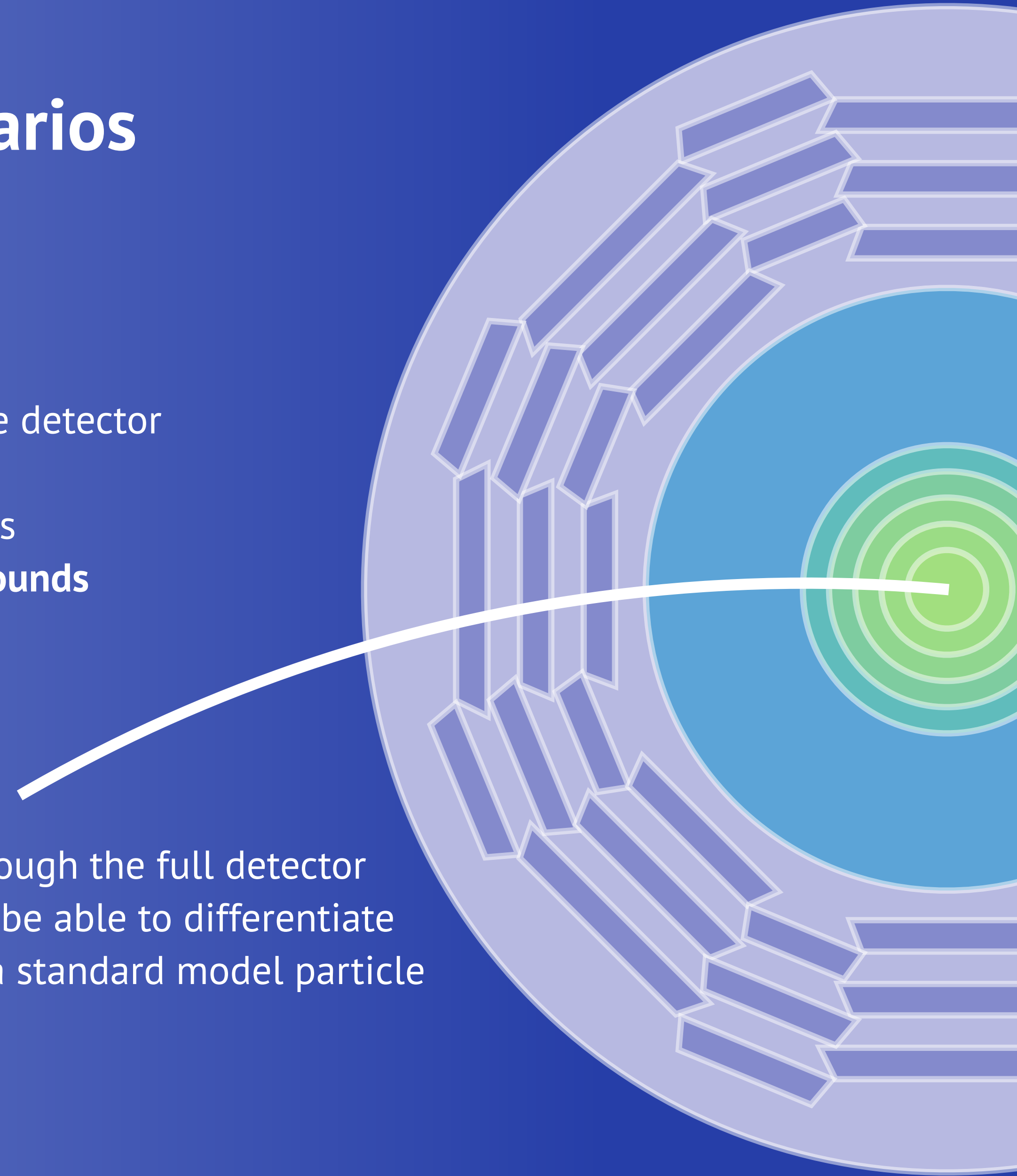
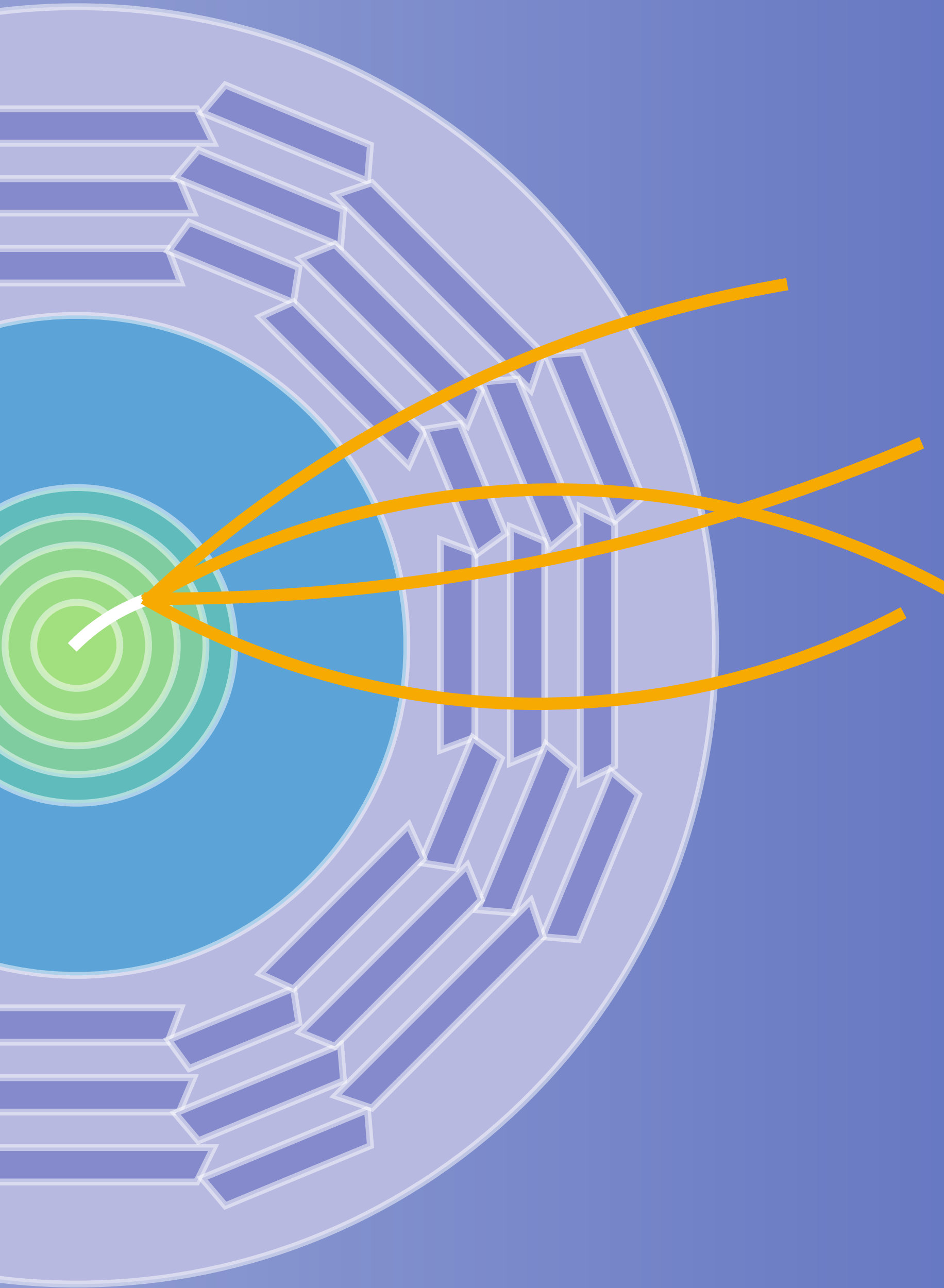
# Two main scenarios

## Metastable:

- Decays somewhere inside the detector
  - need different **algorithms** depending on where it decays
  - deal with different **backgrounds**

## Stable:

- Passes through the full detector
  - need to be able to differentiate this from a standard model particle



# Two main scenarios

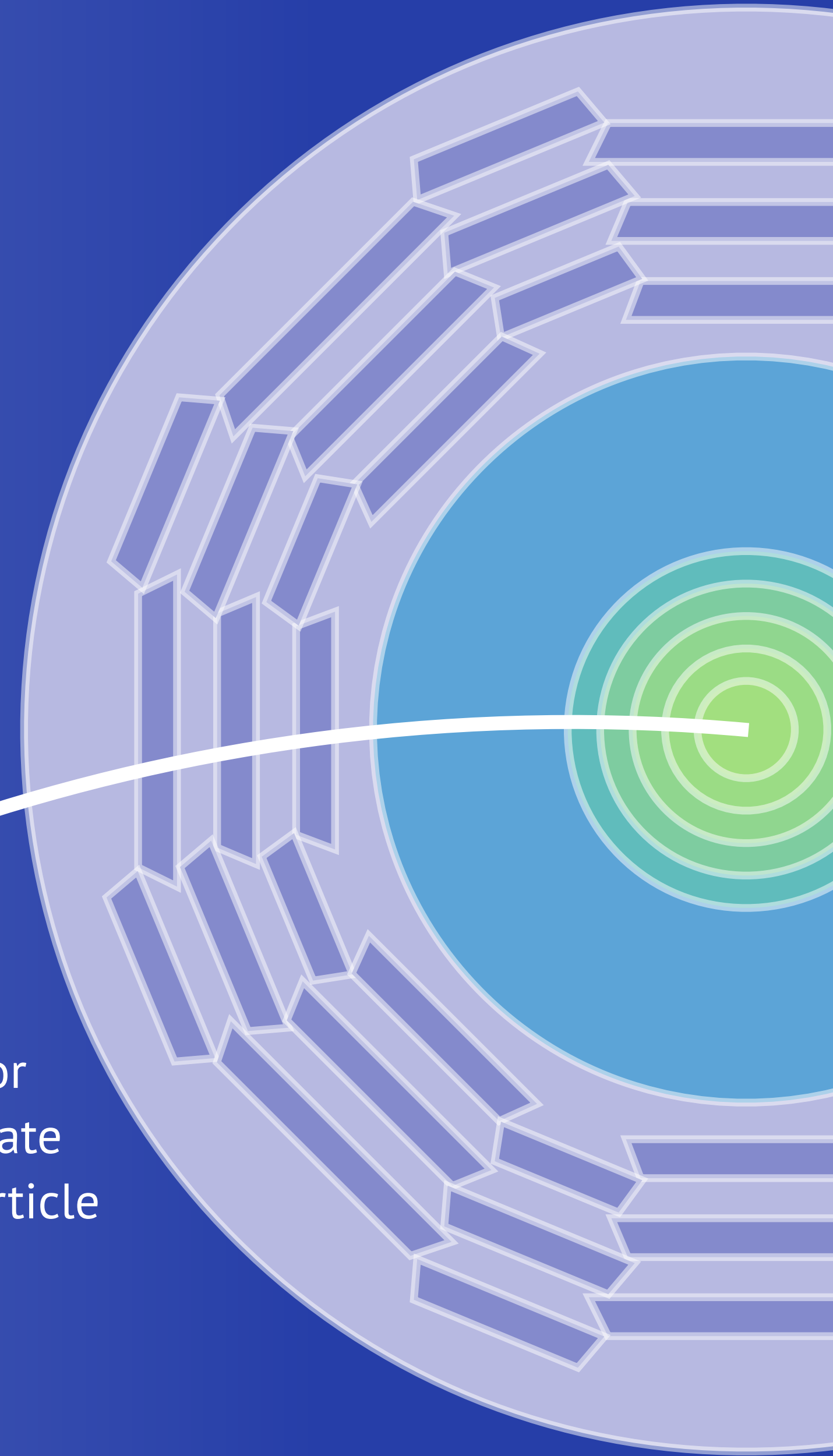
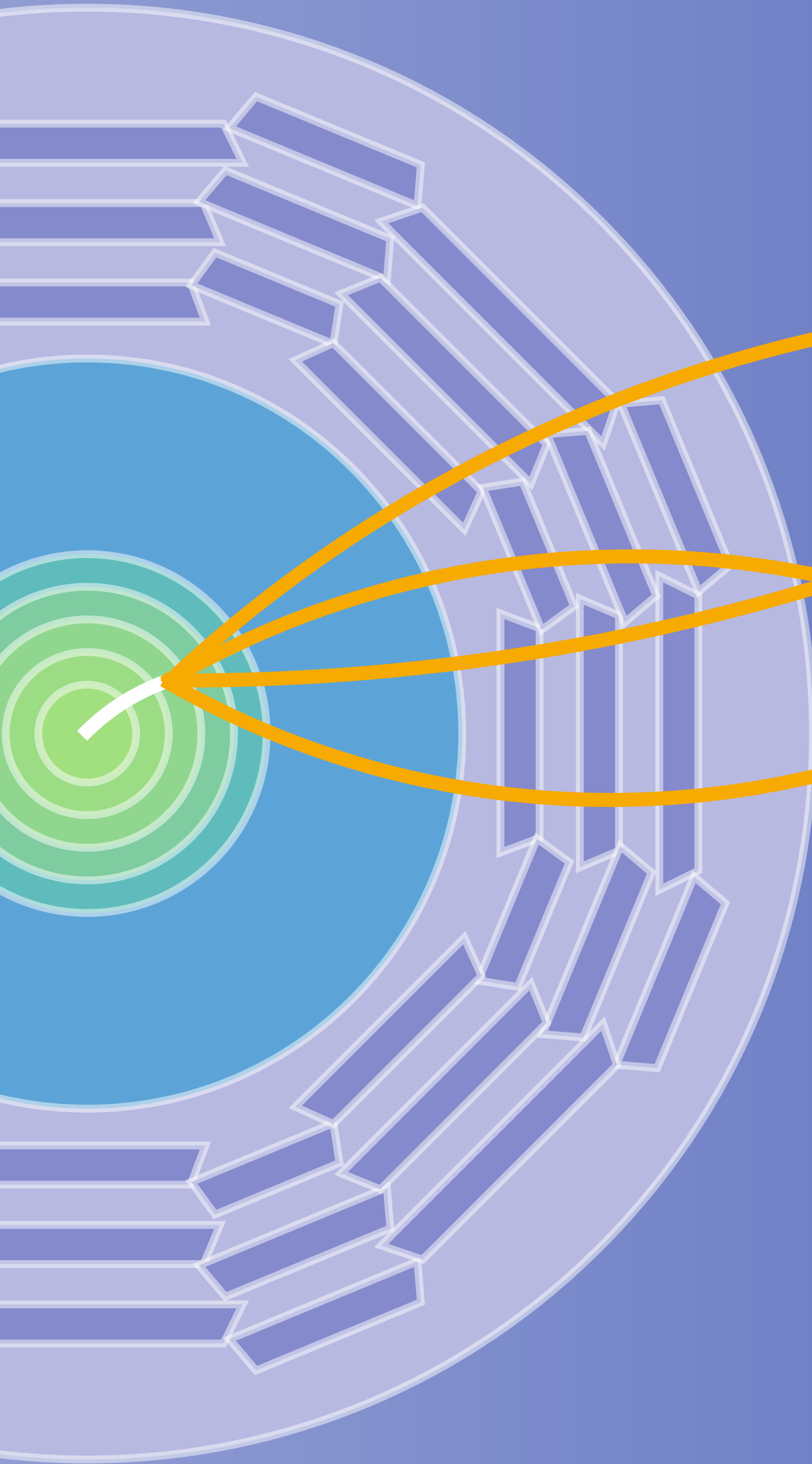
## Metastable:

Decays somewhere inside the detector  
– need different **algorithms**  
depending on where it decays  
deal with different backgrounds

Unusual signatures require dedicated  
(often time consuming) techniques

## Stable:

Passes through the full detector  
– need to be able to differentiate  
this from a standard model particle

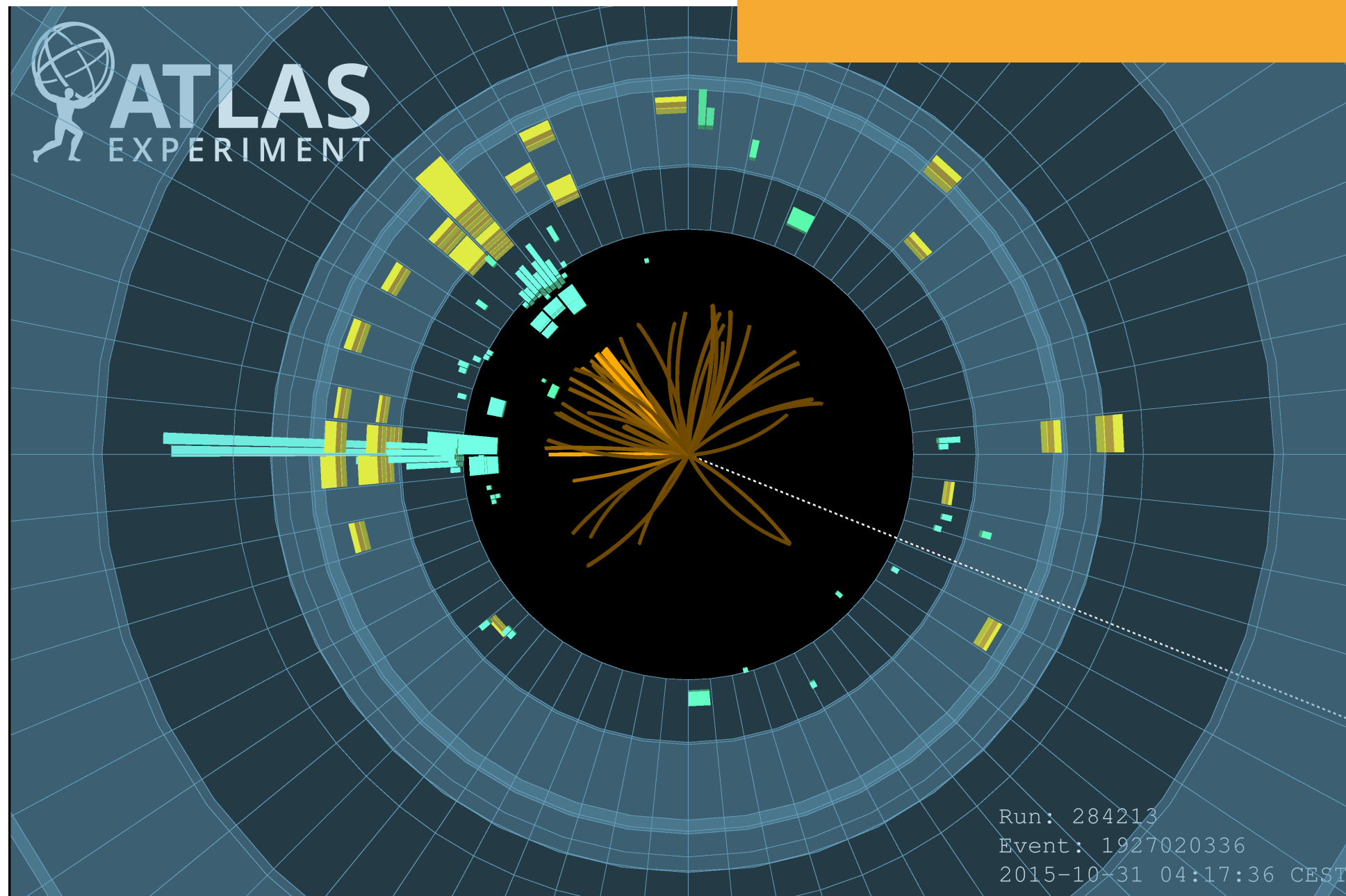




## Problem:

By the time you can do this special reconstruction, ATLAS has already thrown away more than 99.99% of collisions

## The ATLAS Trigger



ATLAS public event displays

Final decision on what to keep is made in around 250 ms

**How do we decide if this event is worth keeping?**

(image of an event with analysis-level “offline” reconstruction)

Level 1 trigger decisions are made with rough **calorimeter** and **muon** information

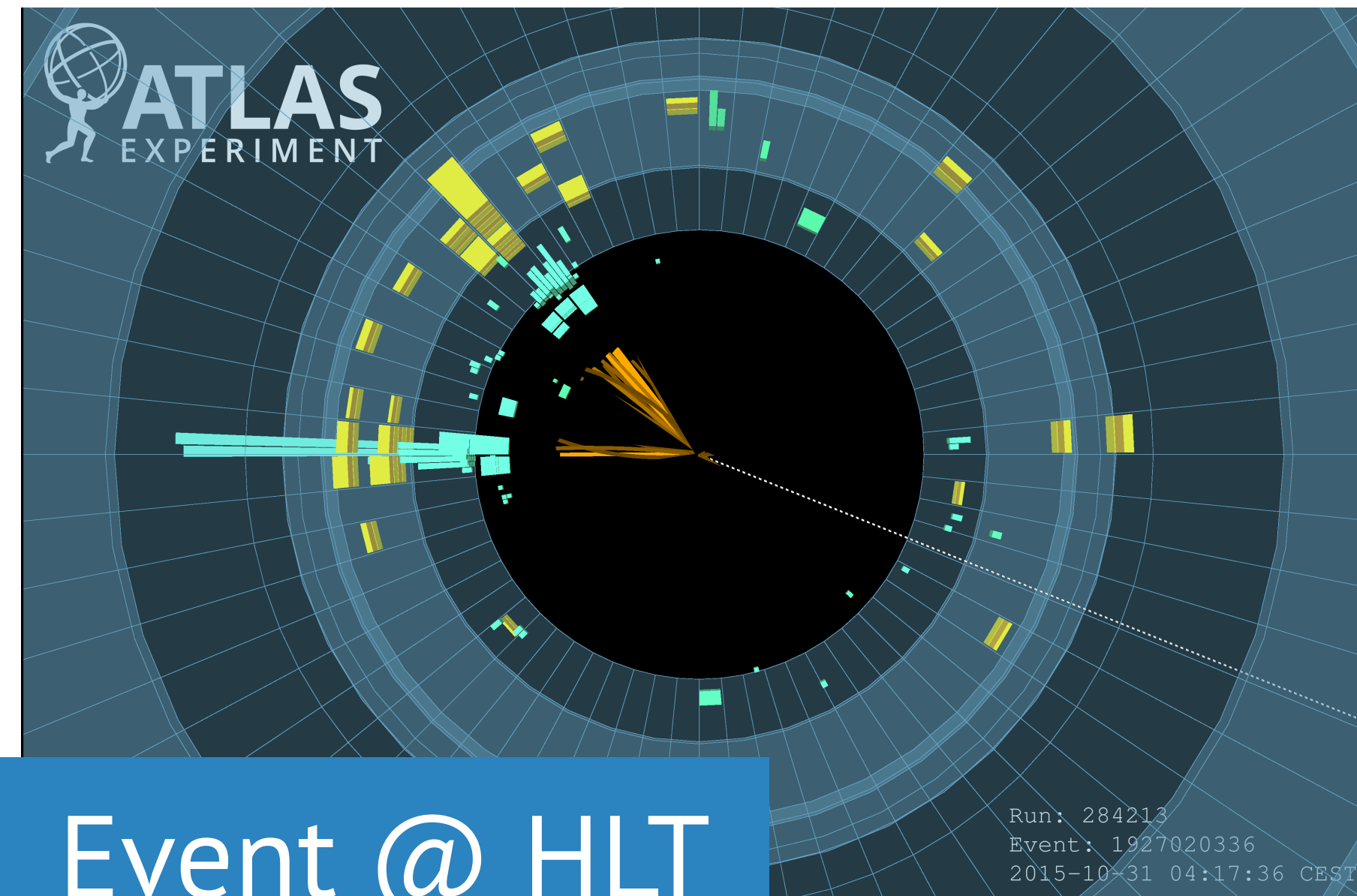
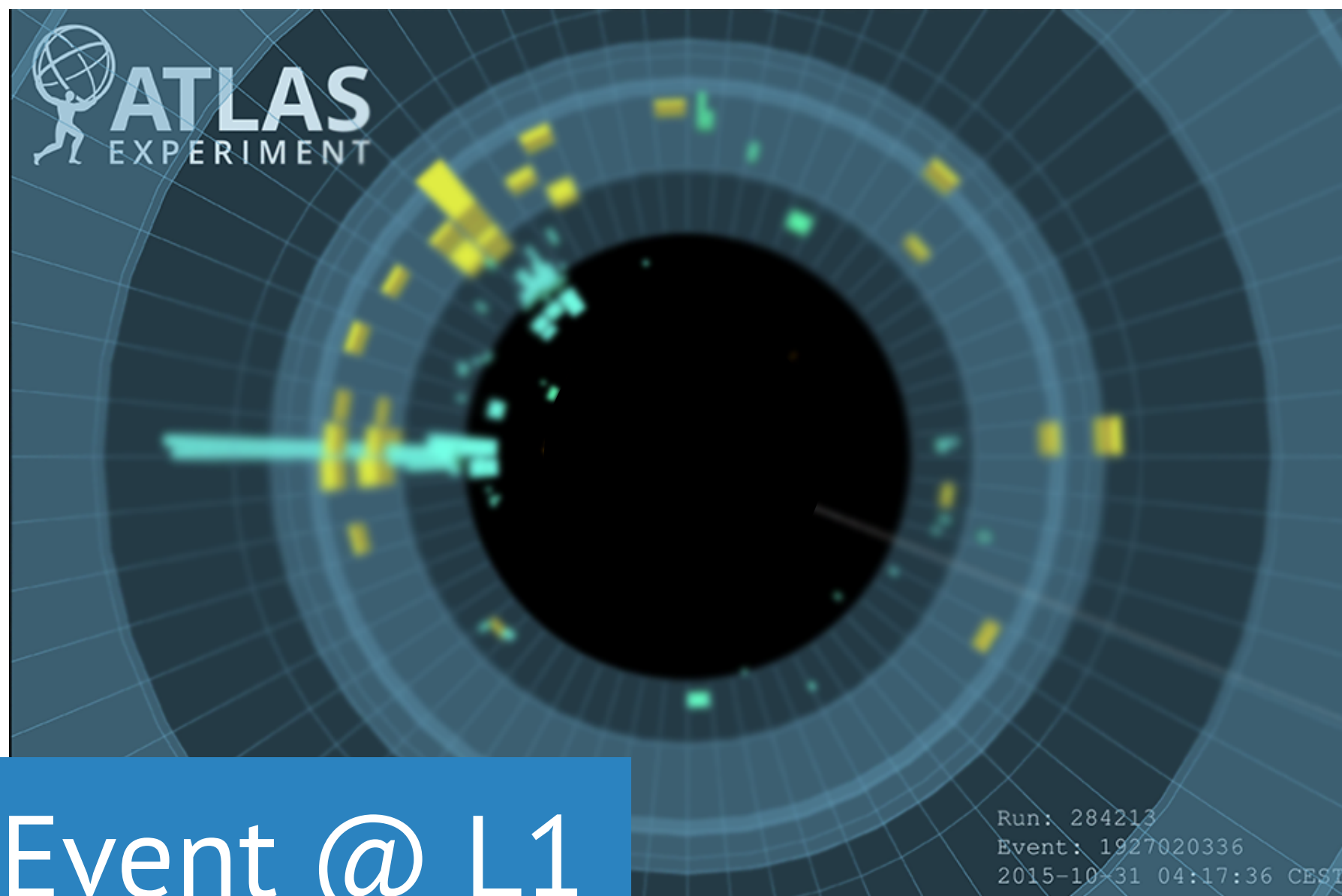
40 MHz →  
100 kHz

hardware

High Level Trigger uses **full precision** information in **small regions**

100 kHz →  
1 kHz

software



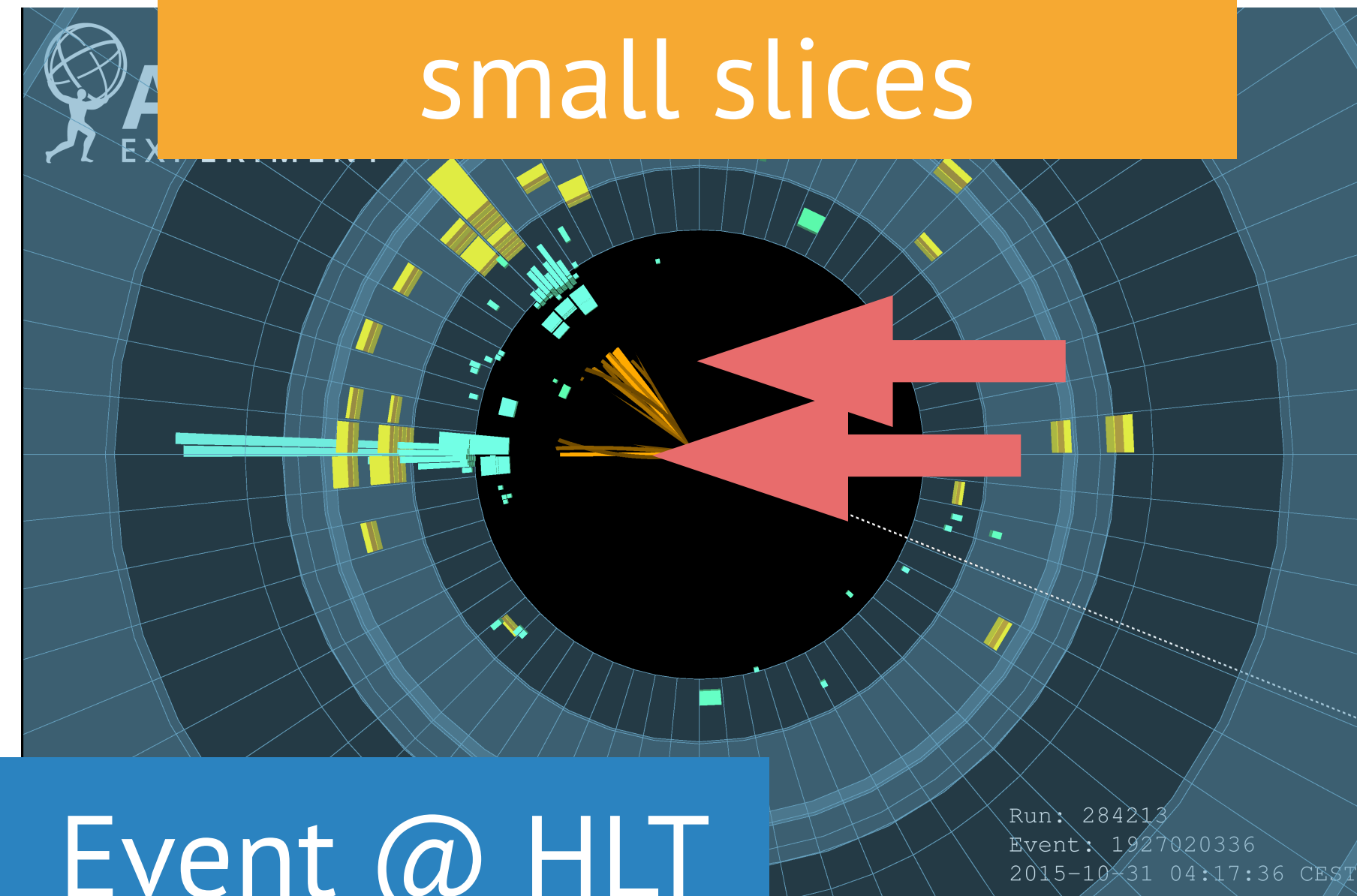
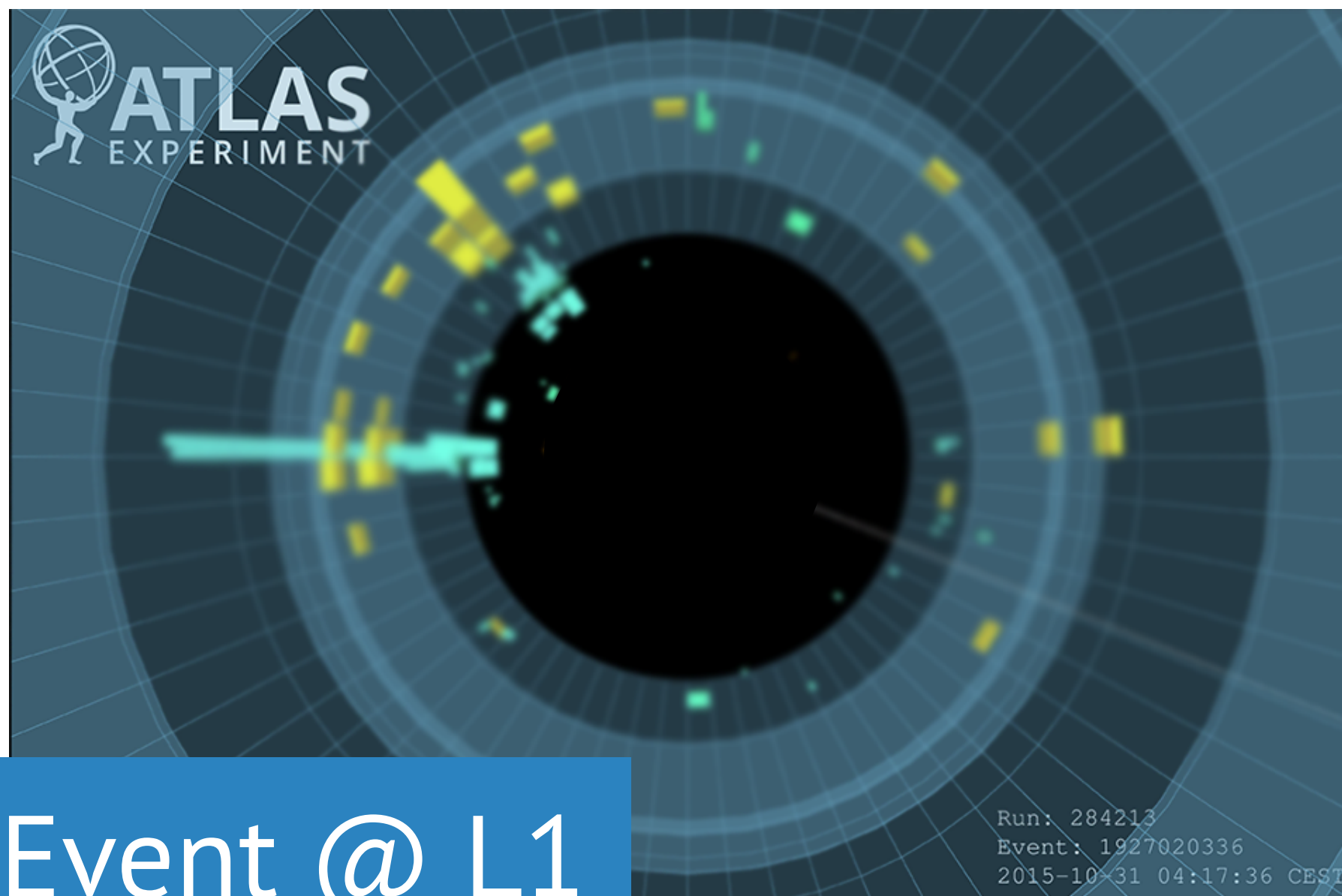


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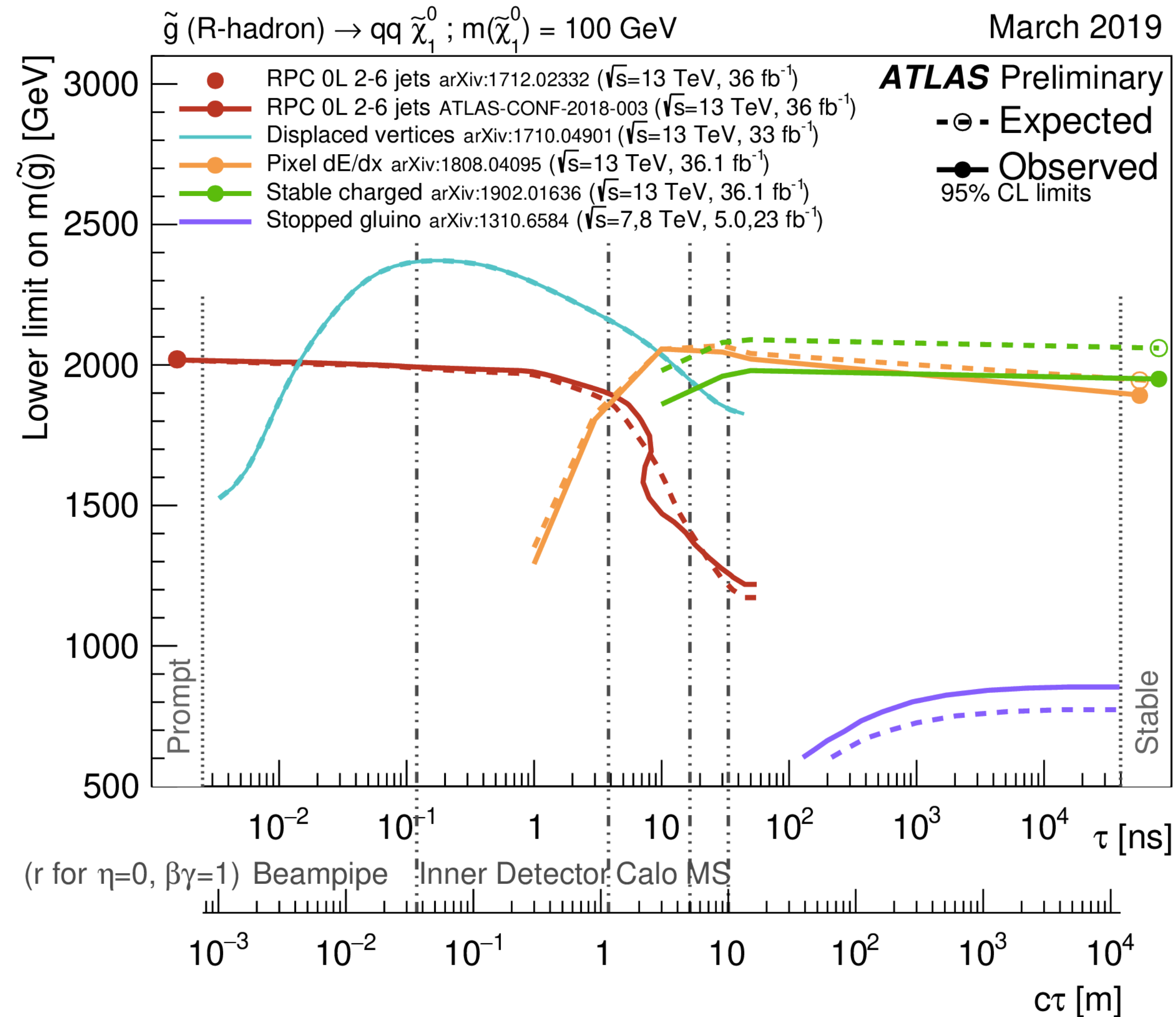
High Level Trigger uses **full pre**  
information in **small region**

tracking:  
best measurement of  
particle lifetime

only available in  
small slices

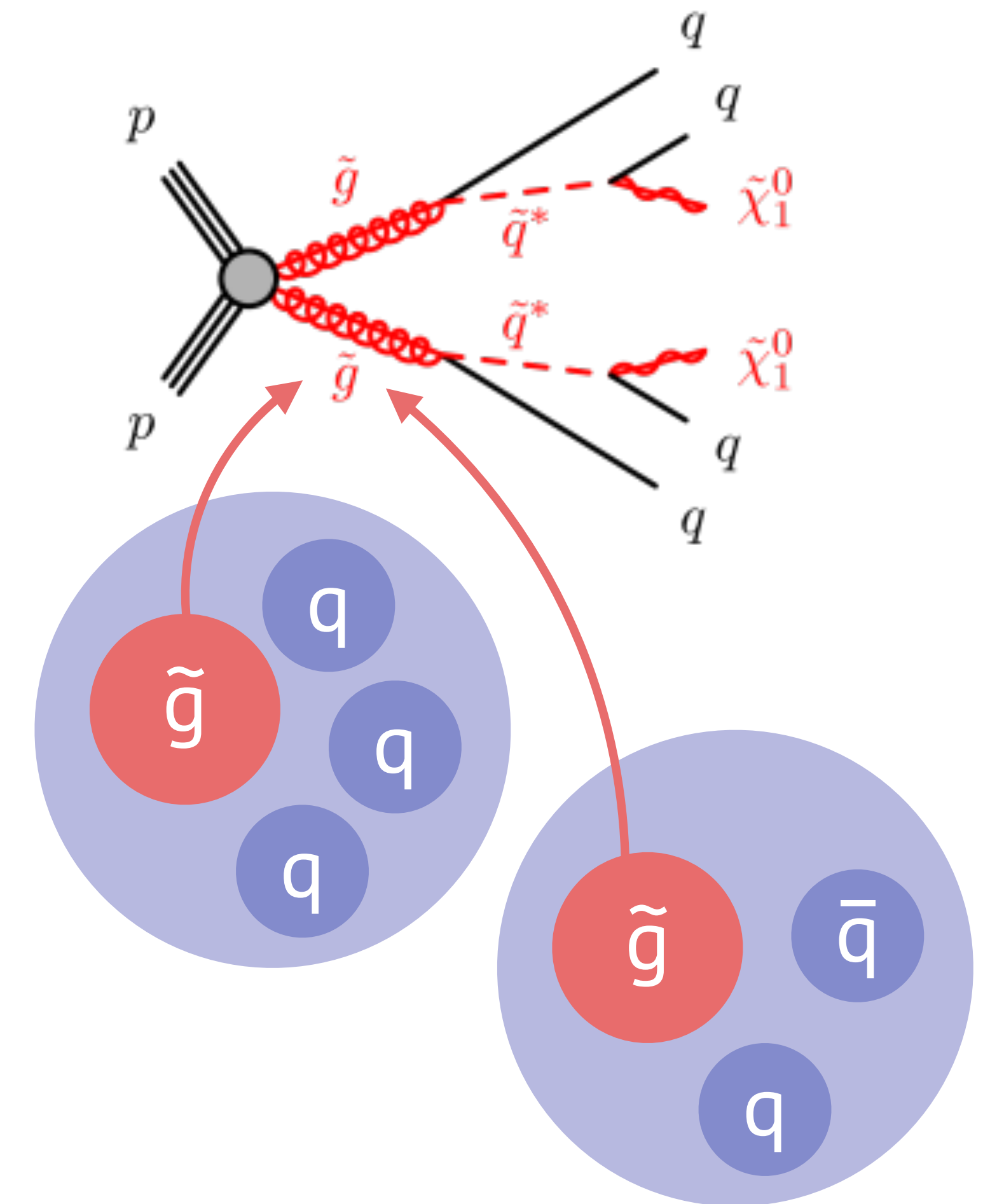


# ATLAS searches for long-lived particles



an example of success  
**R-hadrons**

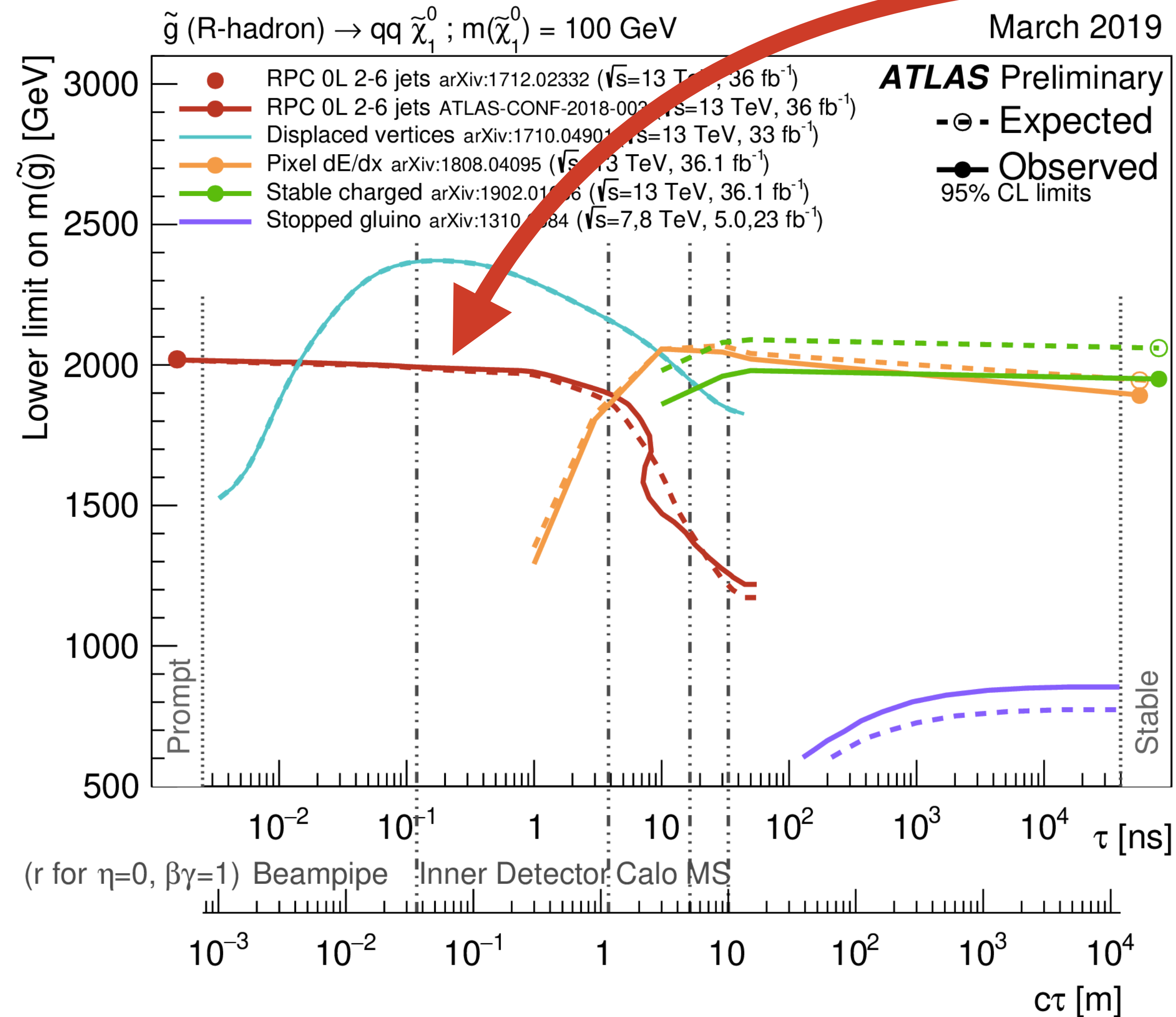
hadron formed with a  
metastable SUSY particle



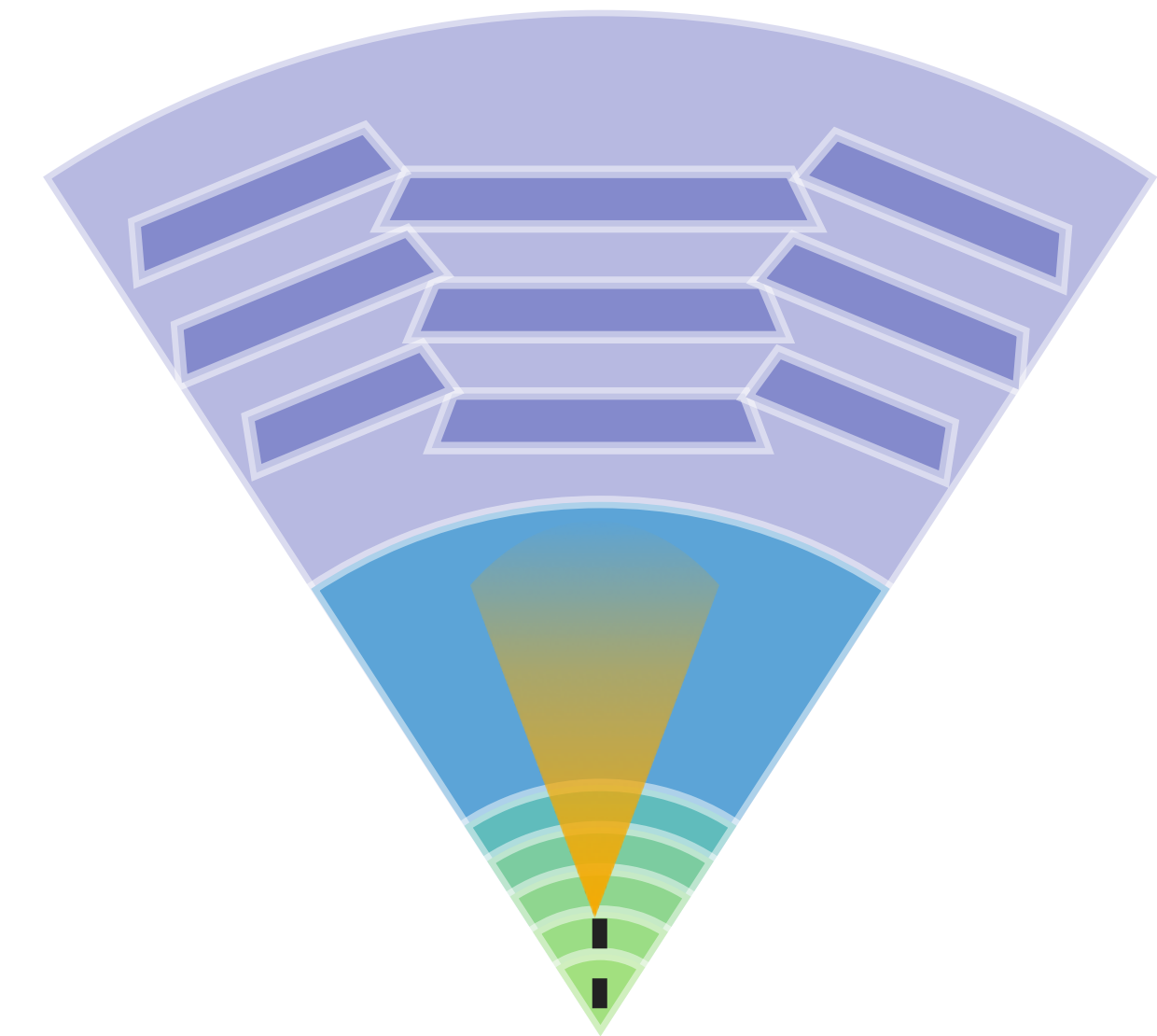


# ATLAS searches for long-lived particles

an example of success  
R-hadrons



## Prompt Search

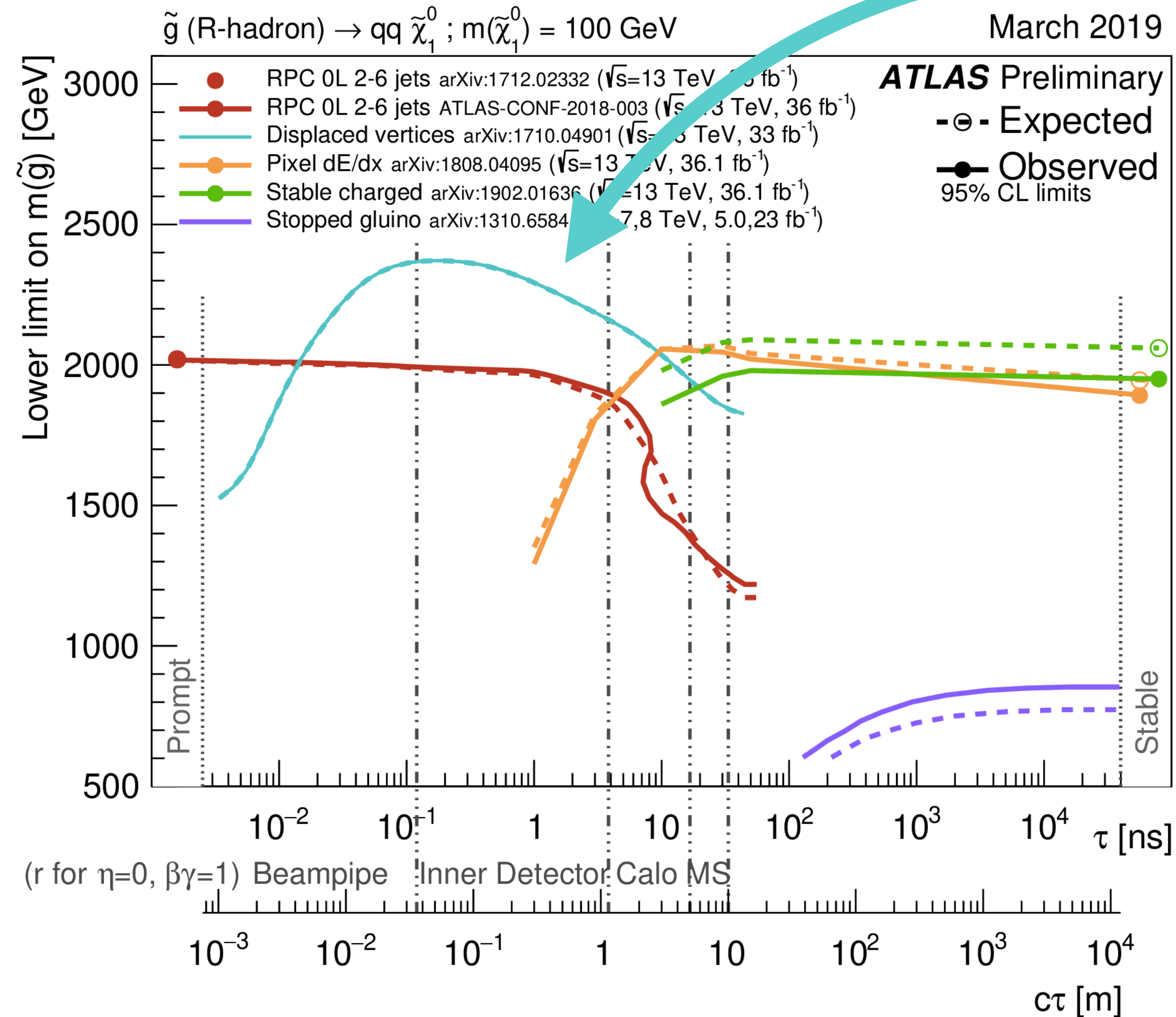


triggers on missing energy  
from neutralino

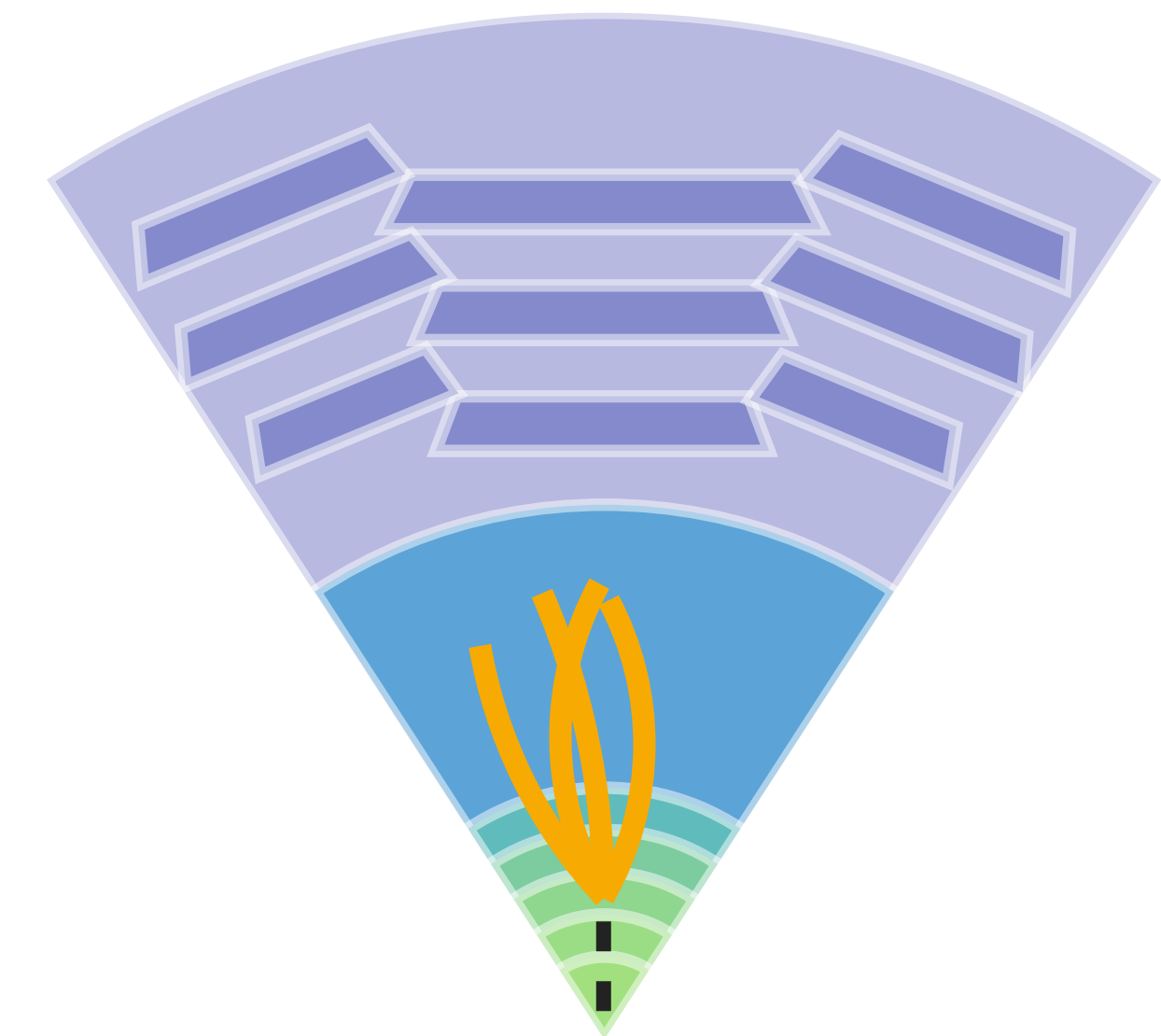
original search  
required decay products  
to have prompt tracks  
**modified to increase acceptance**

# ATLAS searches for long-lived particles

an example of success  
R-hadrons



Displaced Vertices



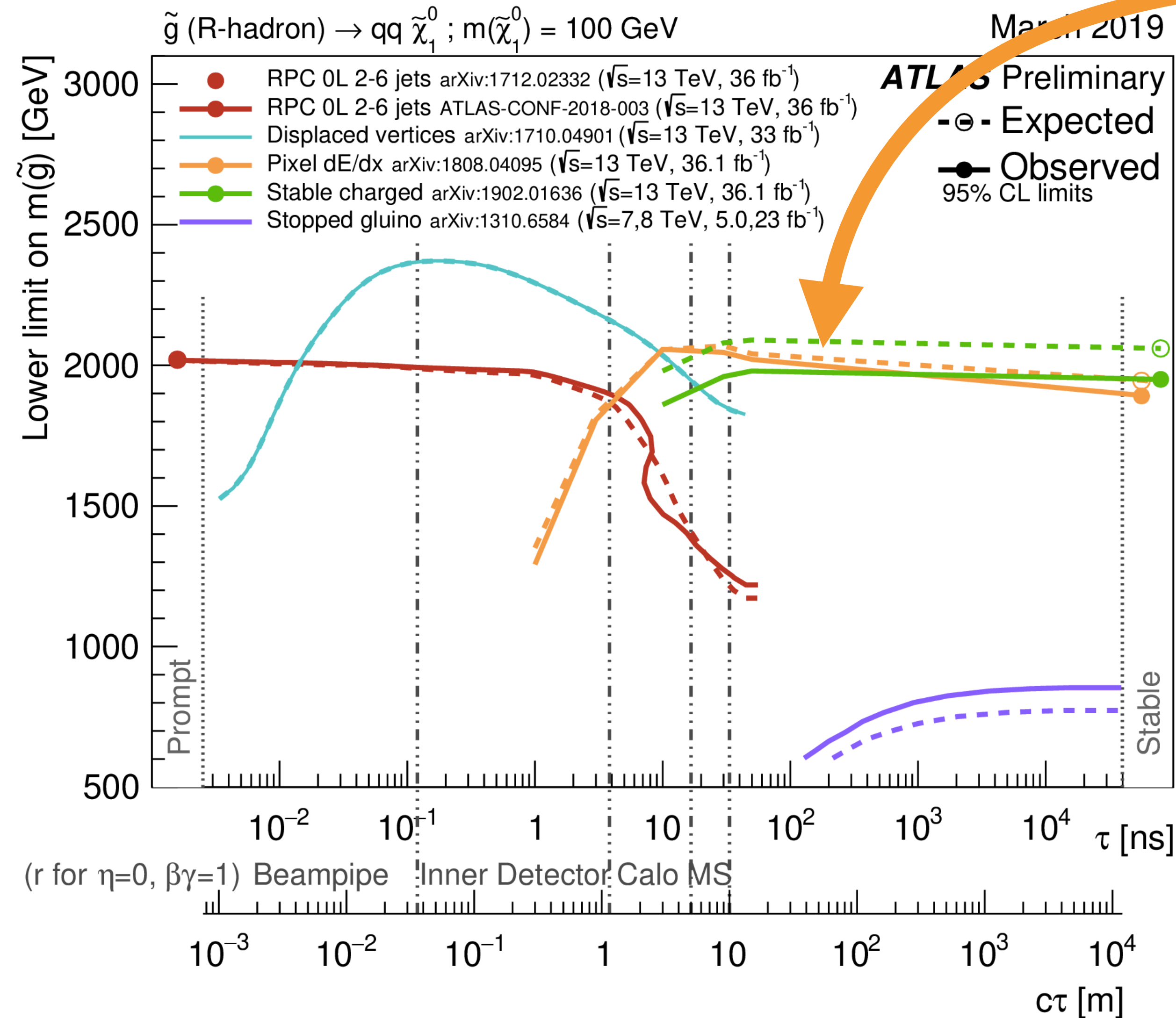
triggers on missing energy  
from neutralino

reconstructs displaced  
vertices with dedicated  
tracking algorithms

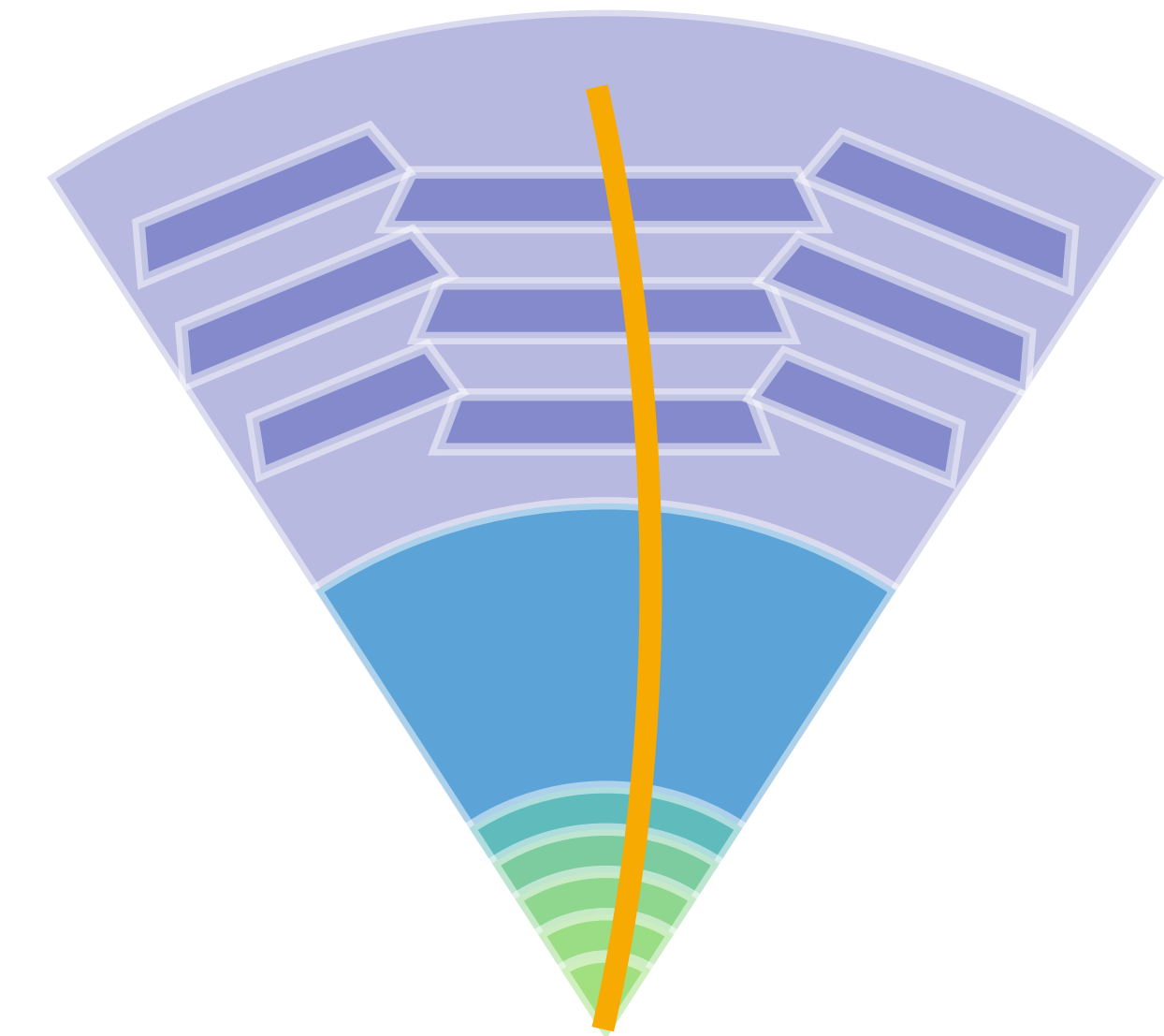


# ATLAS searches for long-lived particles

an example of success  
R-hadrons



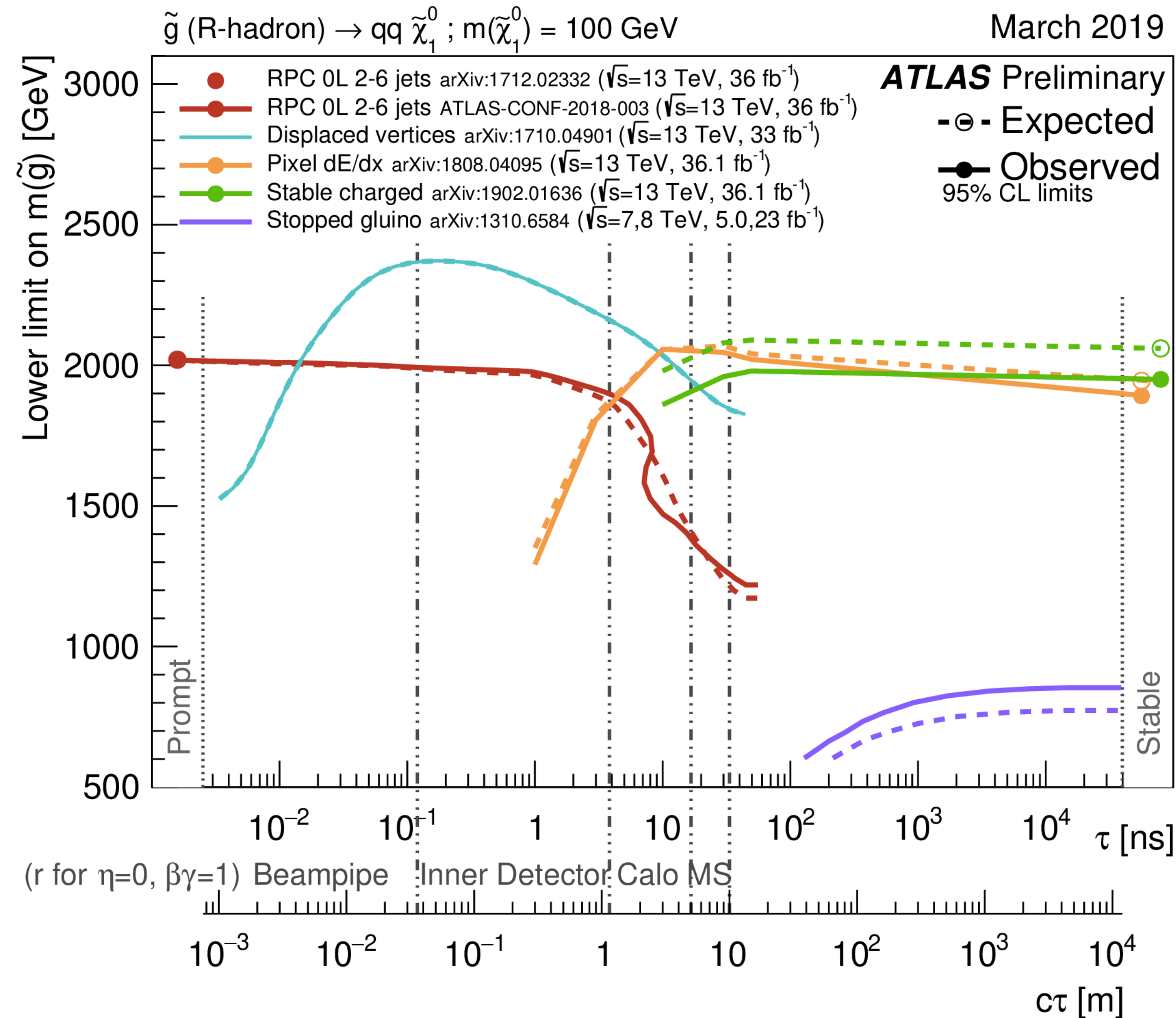
dE/dx



triggers on missing energy  
from neutralino  
(works less well when stable)

calculate energy deposition  
in silicon sensors to find  
high-mass particles

# ATLAS searches for long-lived particles



**Many different dedicated searches for R-hadrons**

Together they cover the full lifetime range

Each signature has different strategy, specialized techniques

- ▷ modified track requirements
- ▷ specialized track and vertex reconstruction
- ▷ dE/dx calculation

None use a trigger related to long lifetimes

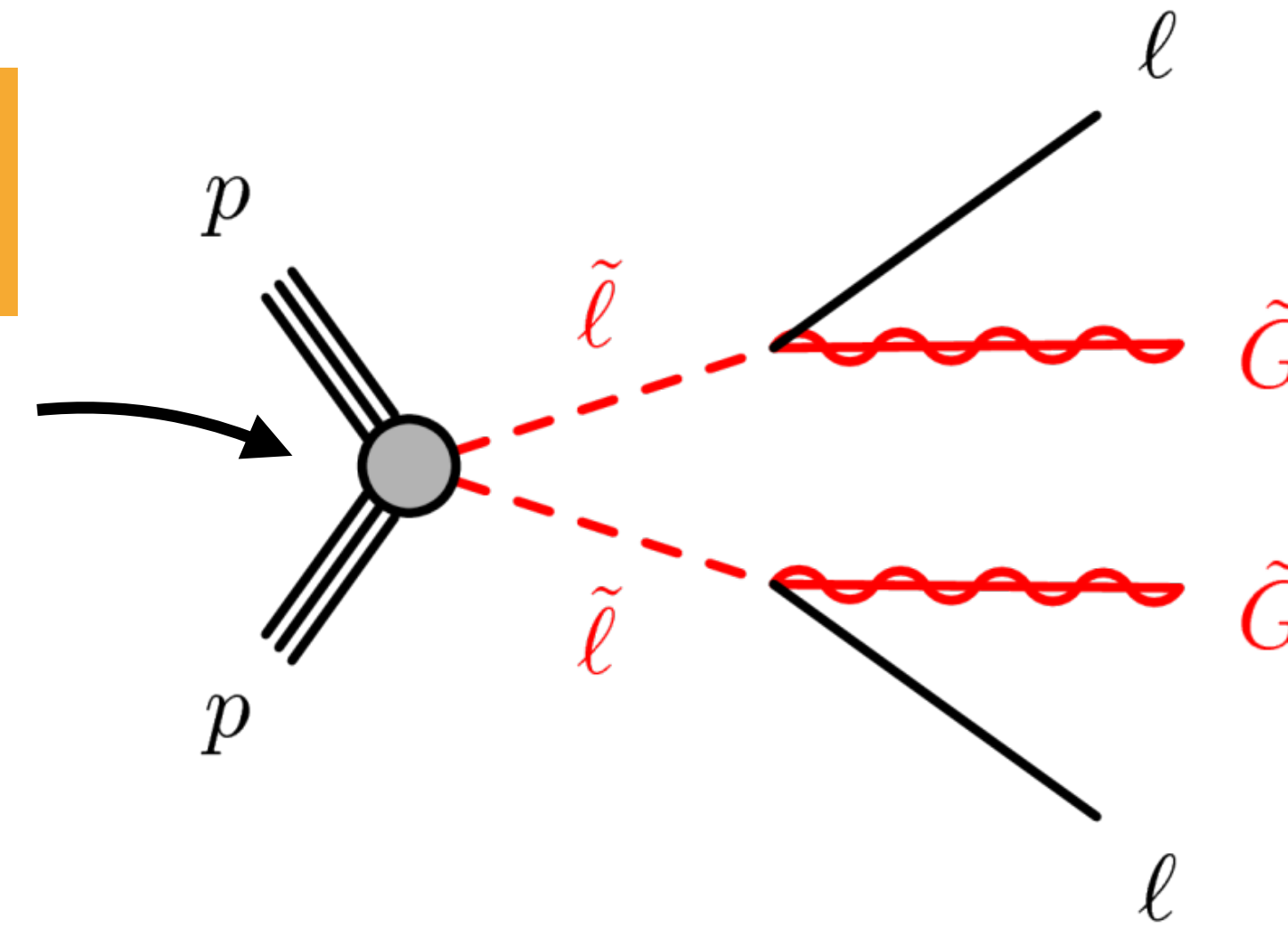
But in this case, we've  
done great



## Uncovered territory

**Still models that hadn't been looked at, like**

For intermediate lifetimes, tightest limit from OPAL  
(15 years ago, a fraction of the LHC energy)



One we don't have limits  
is because it's so **rare**

If they exist, the LHC would have produced in Run 2:

500 GeV gluinos: 4.7 million

500 GeV sleptons: 94

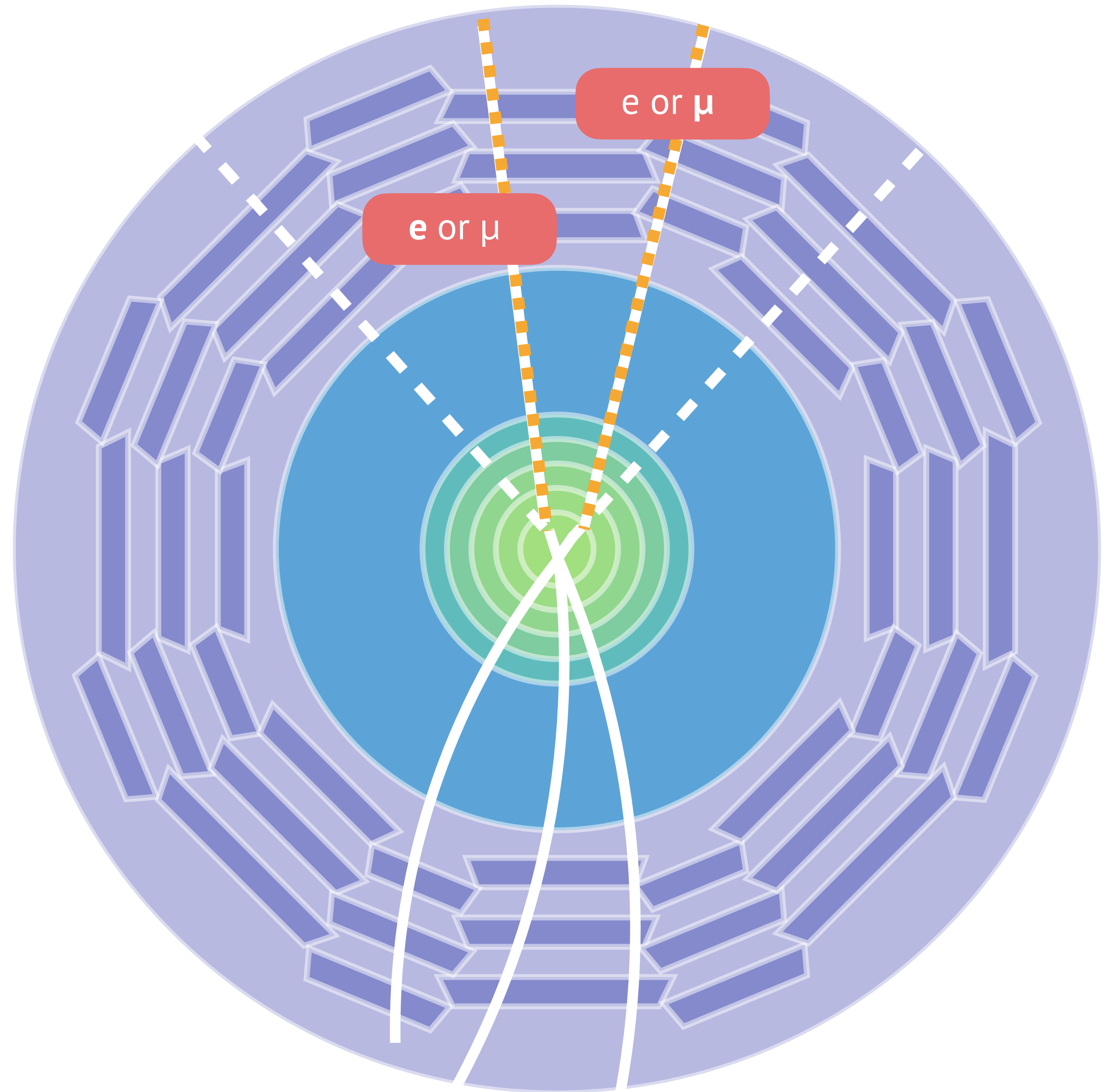
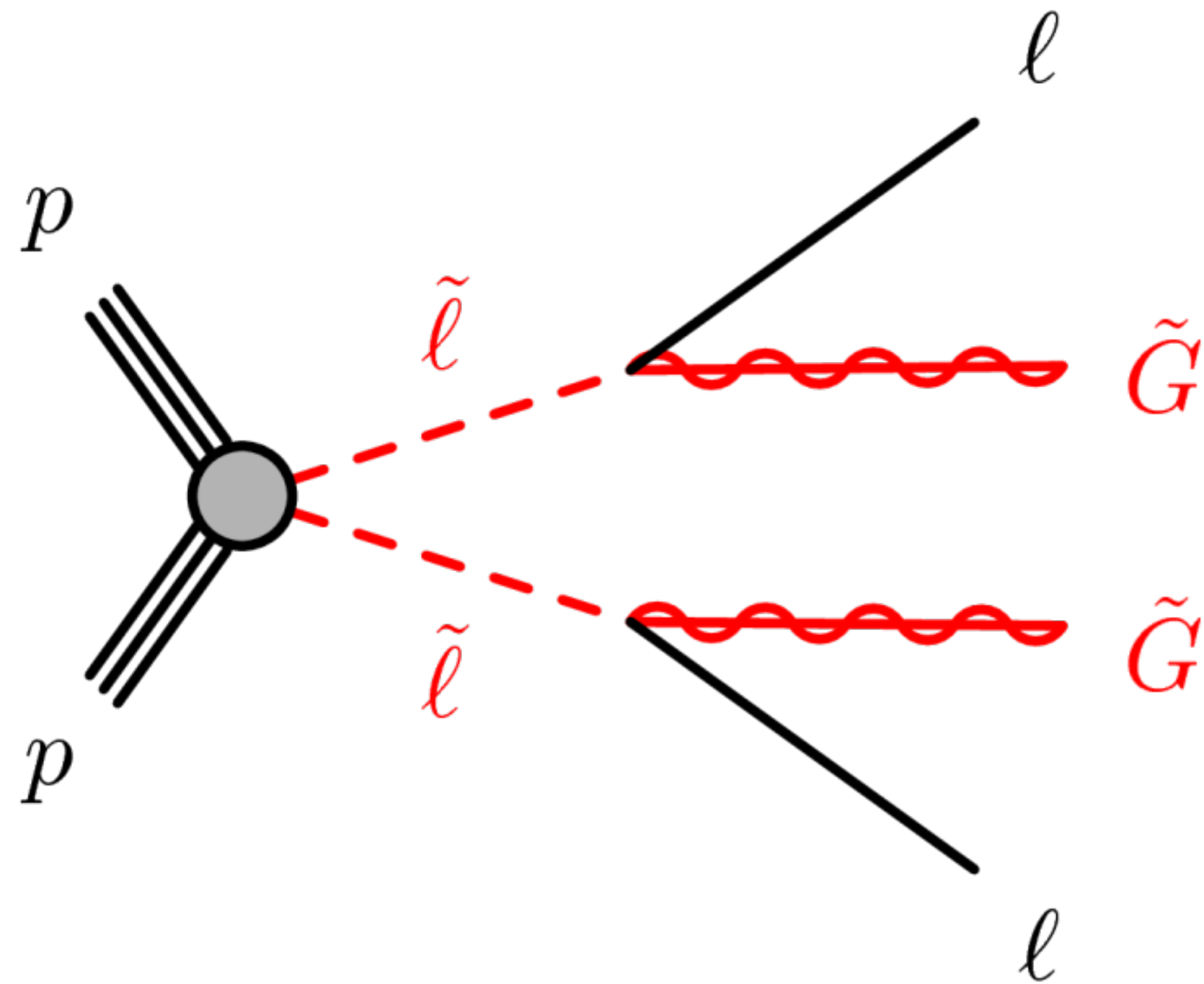
In order to pick out these events, in the sea of LHC data, you have to be really looking for them!

Previously we'd only been able to rule out these particles at **< 90 GeV**

Required a **dedicated search** to go further

# what I've been working on

**Displaced leptons – and nothing else!**



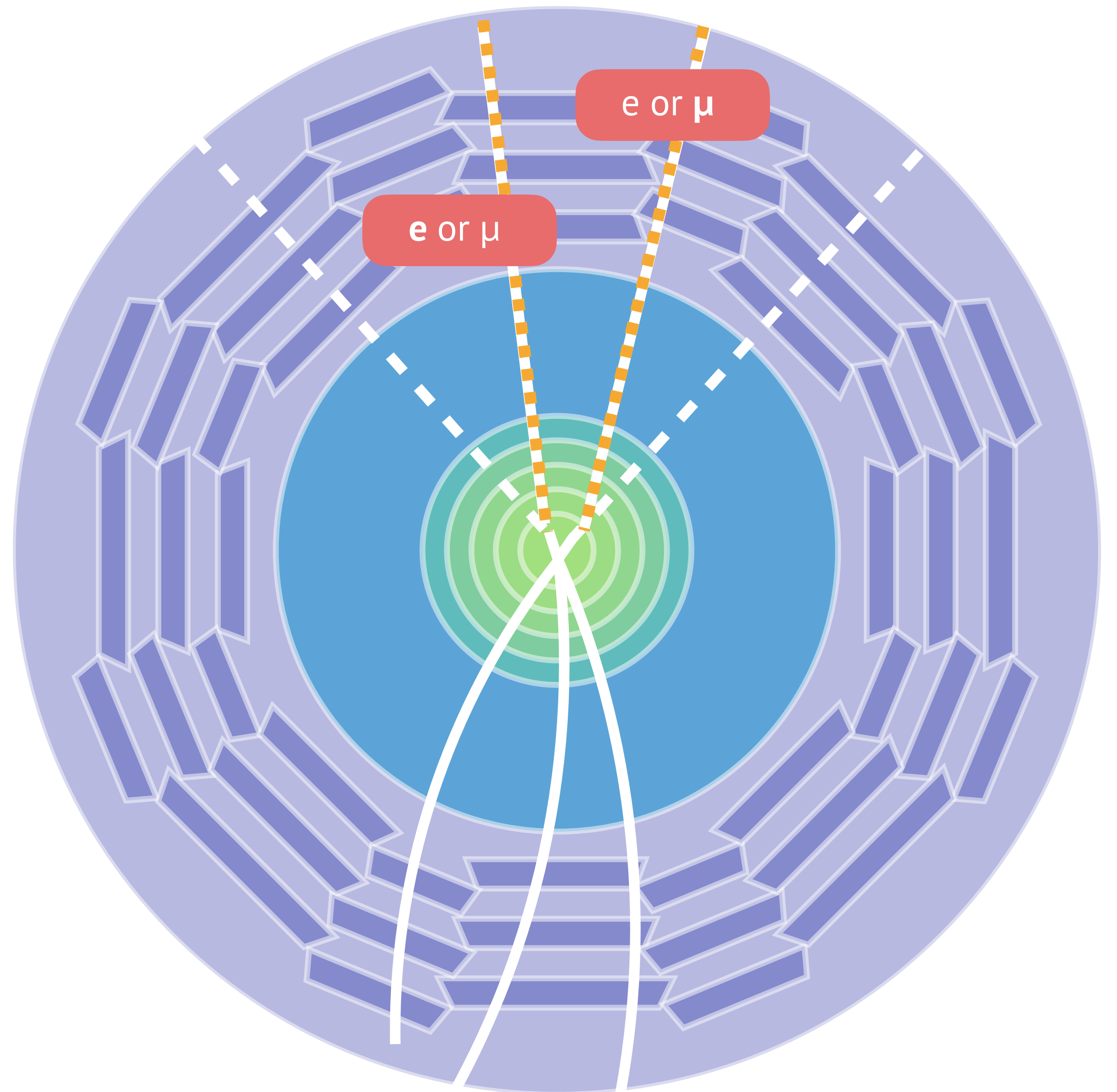


hasn't been done for a  
reason: **it's hard**

Very displaced leptons means **standard reconstruction algorithms** fail

Which may sound like it isn't a big deal,  
but in these huge experiments, those  
algorithms were made by **thousands** of  
people working to perfect them

Not easy to re-do as a tiny team!



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reason: **it's hard**

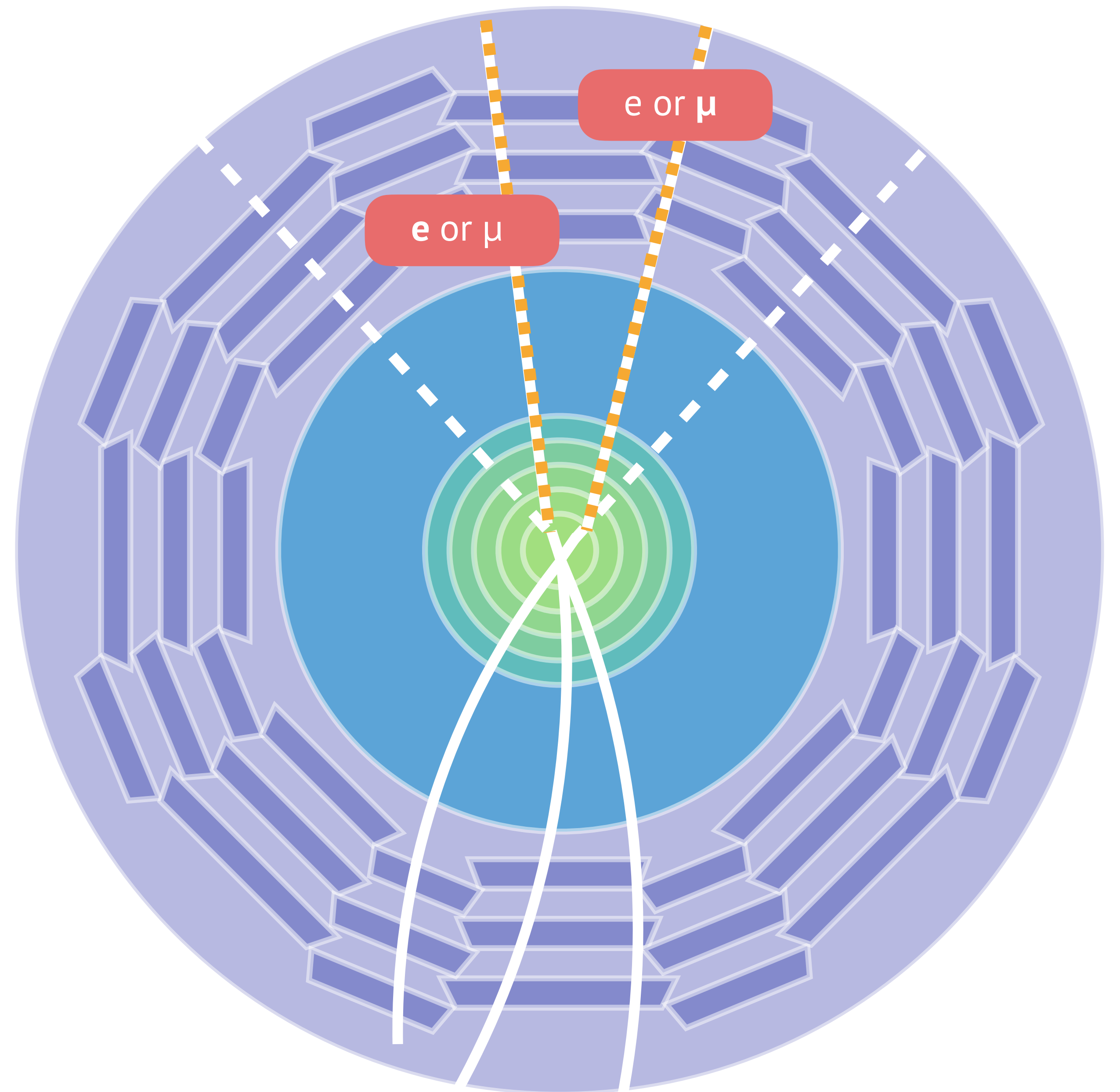
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Which may sound like it isn't a big deal,  
but in these huge experiments, those  
algorithms were made by **thousands** of  
people working to perfect them

Not easy to re-do as a tiny team!

First step: Record the data!

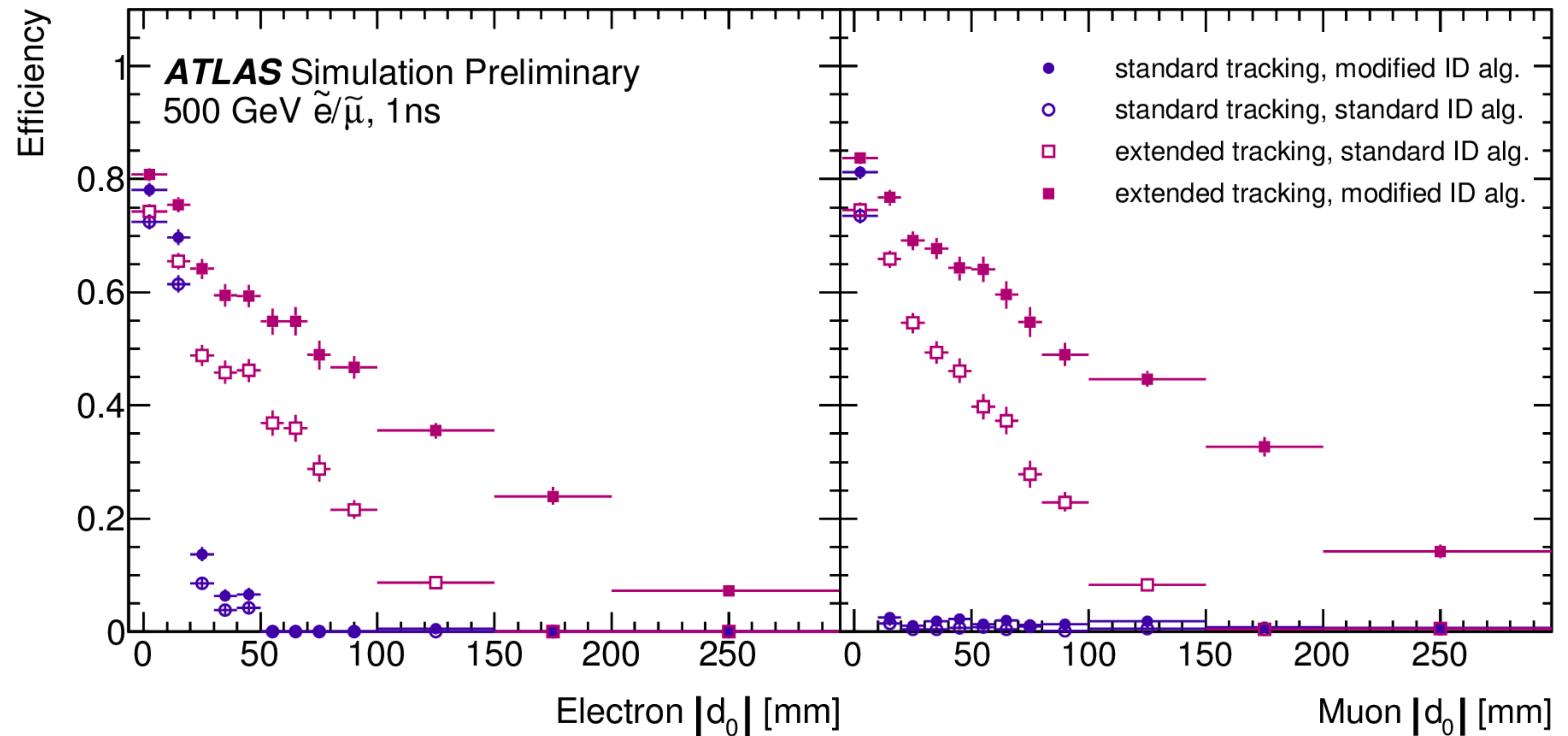
Need **special triggers** for displaced objects  
→ no tracking information so much harder  
to pick out from backgrounds





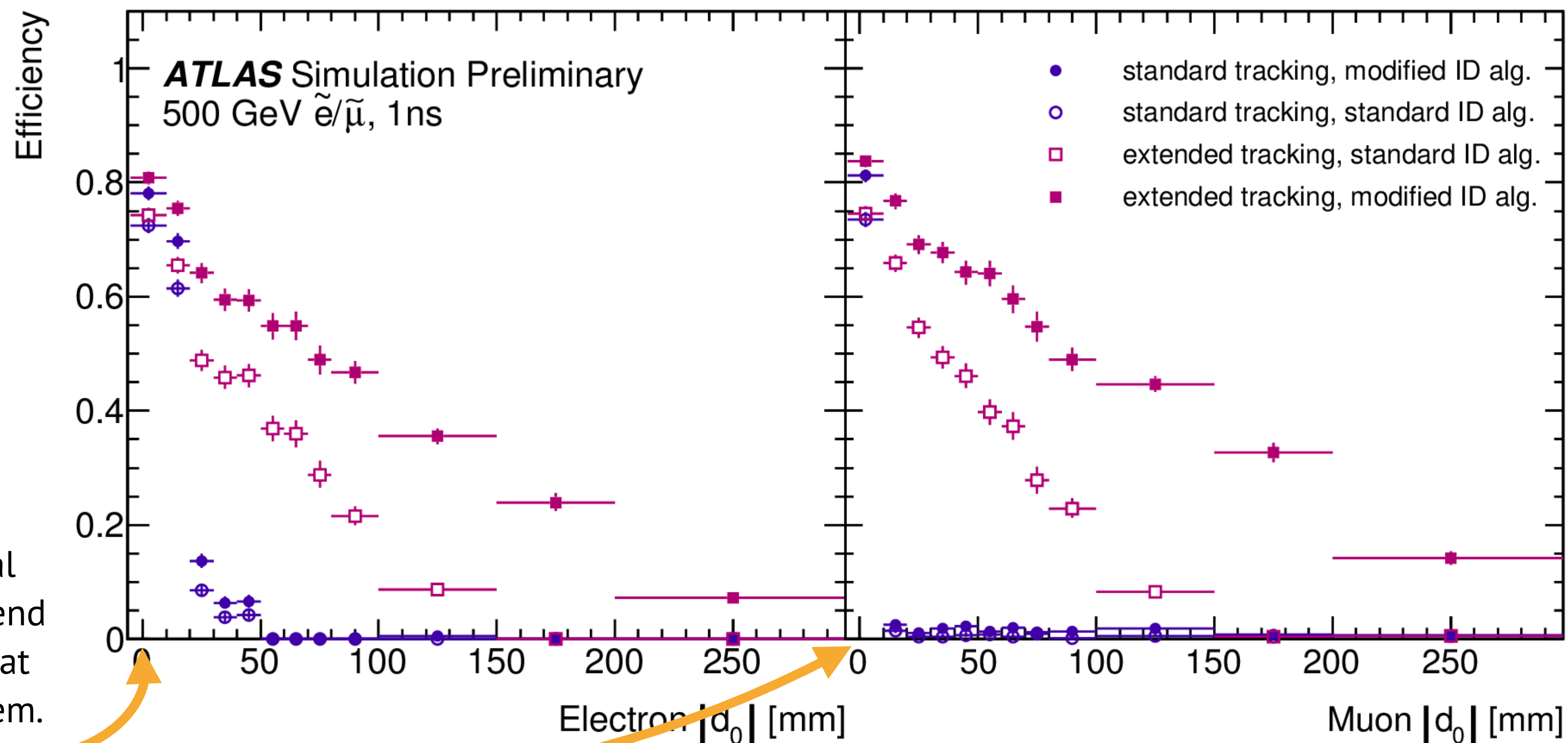
hasn't been done for a  
reason: **it's hard**

Next step: make sure you can **reconstruct** your signal  
Able to get good efficiency out to high values of  
displacement



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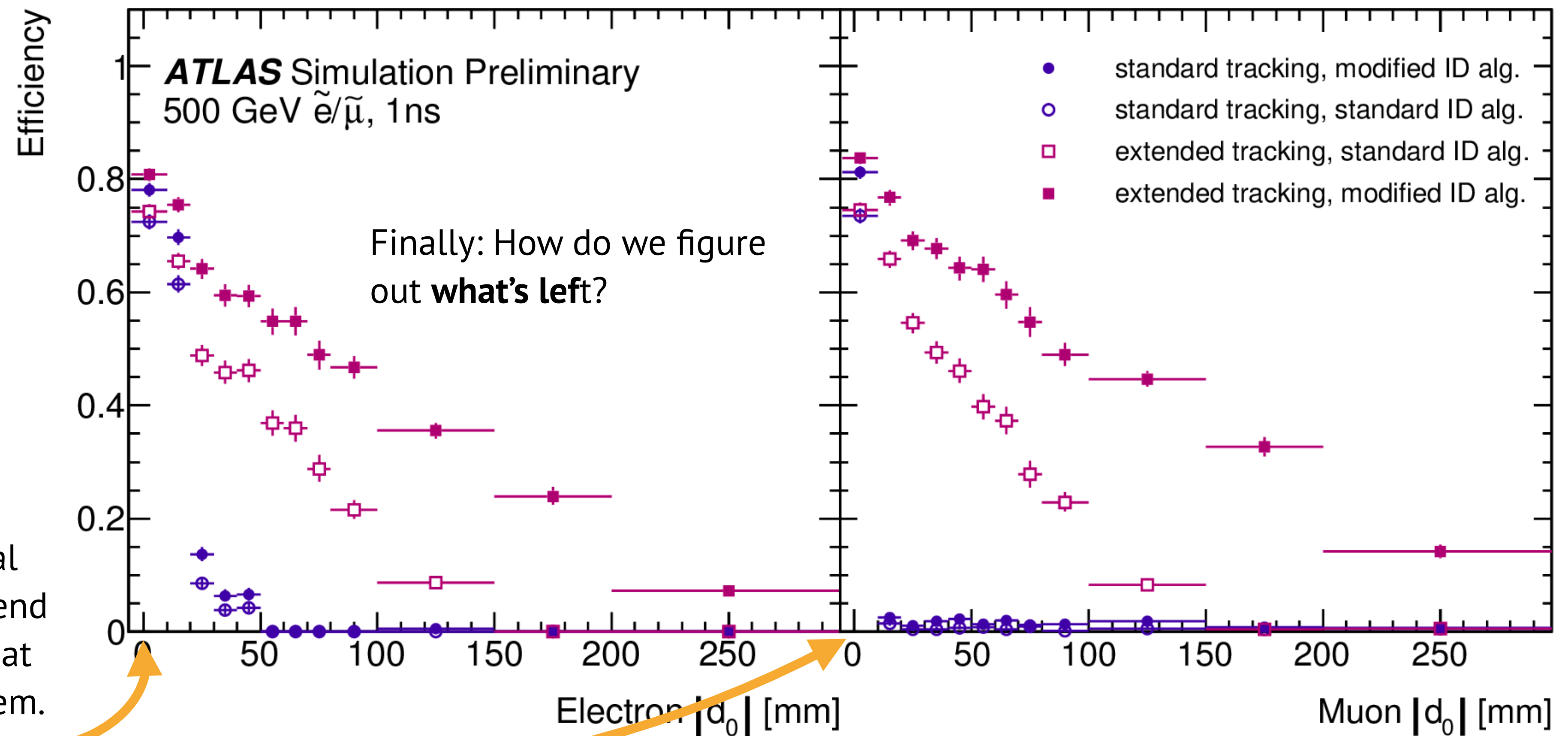


3rd step: **Isolate** the signal  
Almost all SM processes end  
up in the first bin. We cut at  
 $|d_0| > 3$  mm to remove them.

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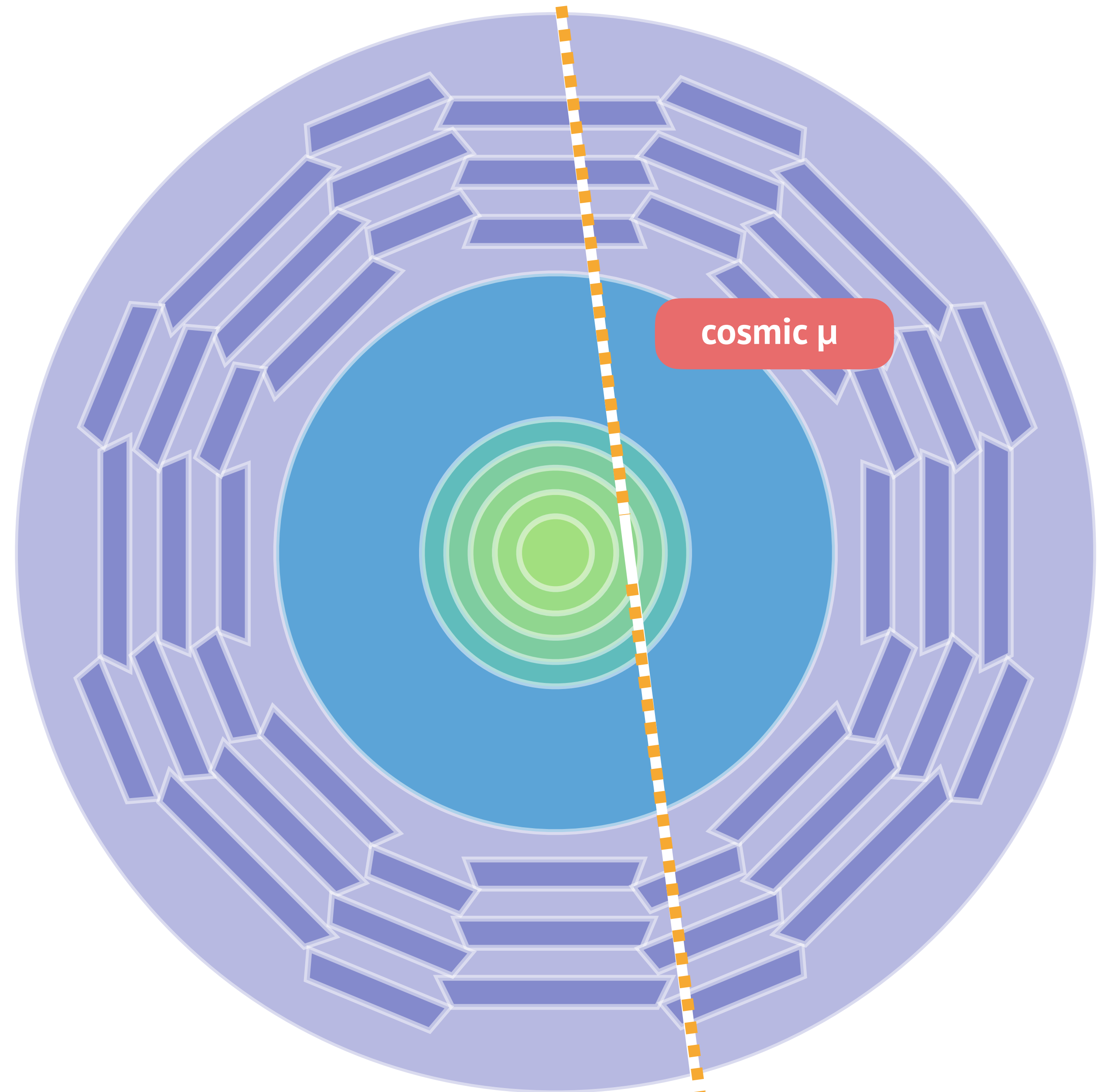




hasn't been done for a  
reason: **it's hard**

When you have a search with tiny  
backgrounds, it can be more **complicated**  
than you expect

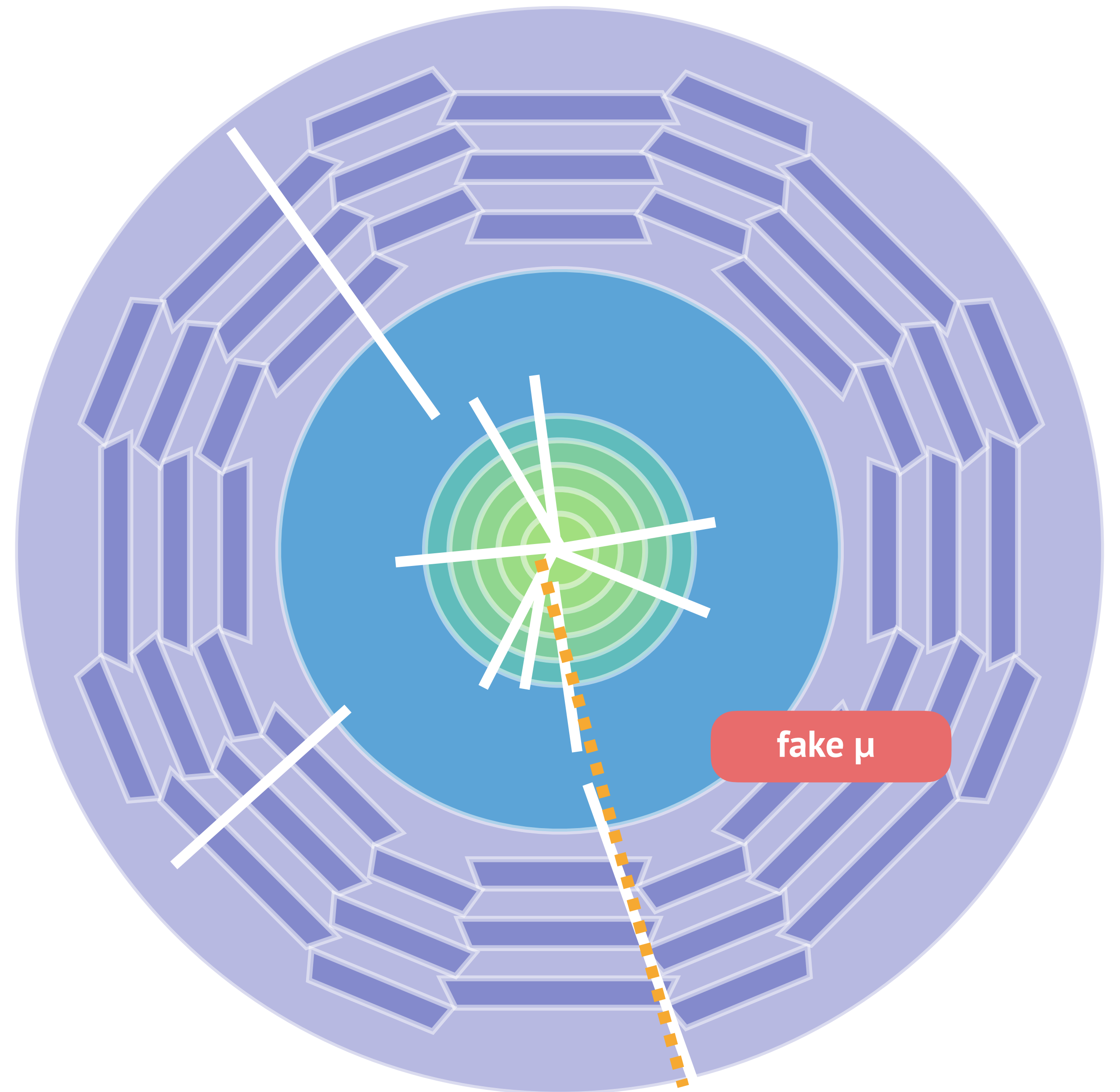
Dominant background for  $\mu\mu$   
**cosmics**



hasn't been done for a  
reason: **it's hard**

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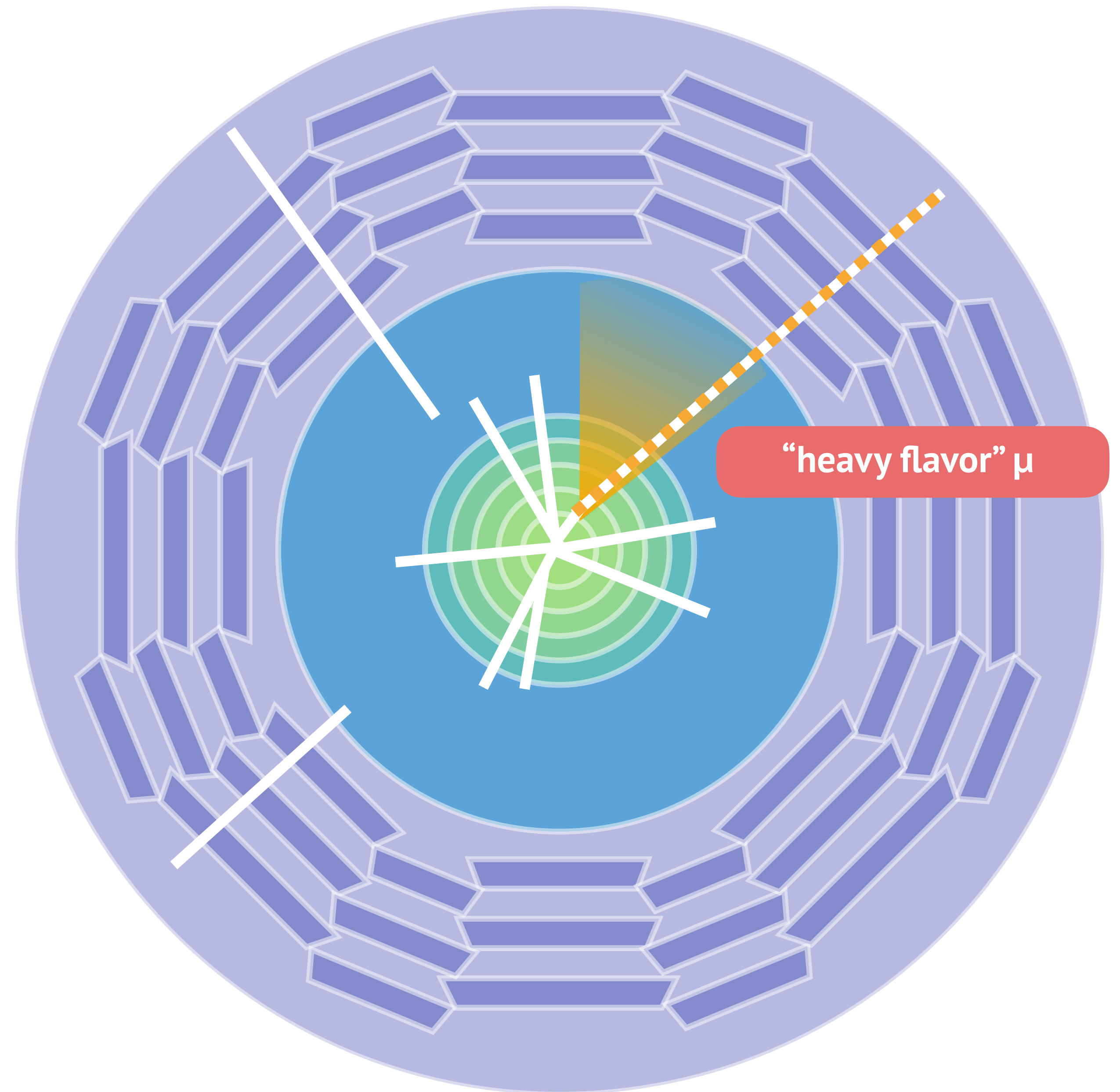
Dominant background for  $\mu\mu$   
**cosmics**  
**fake muons**



hasn't been done for a  
reason: **it's hard**

When you have a search with tiny  
backgrounds, it can be more **complicated**  
than you expect

Dominant background for  $\mu\mu$   
**cosmics**  
**fake muons**  
**rare standard model processes**

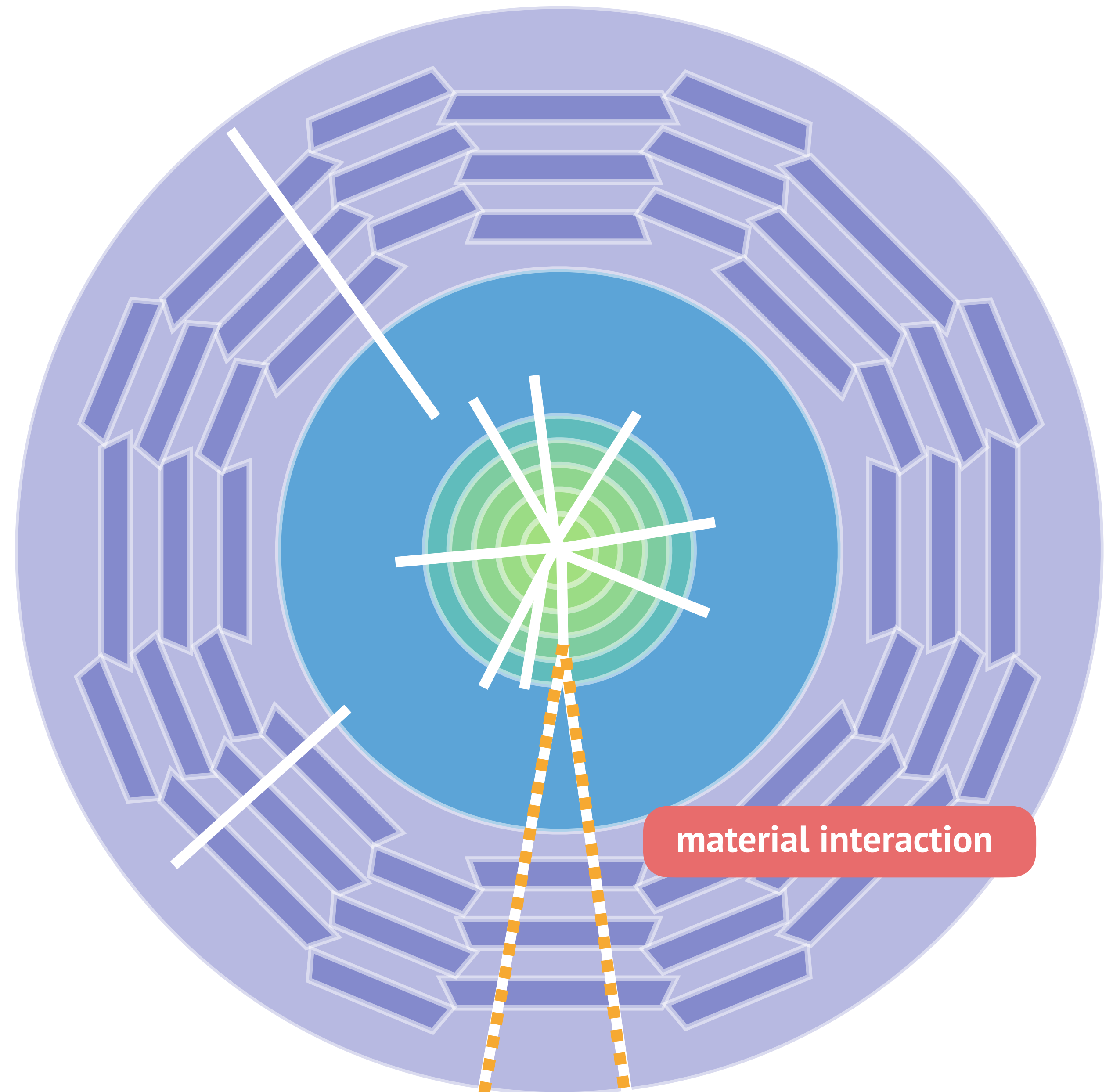




hasn't been done for a  
reason: **it's hard**

When you have a search with tiny  
backgrounds, it can be more **complicated**  
than you expect

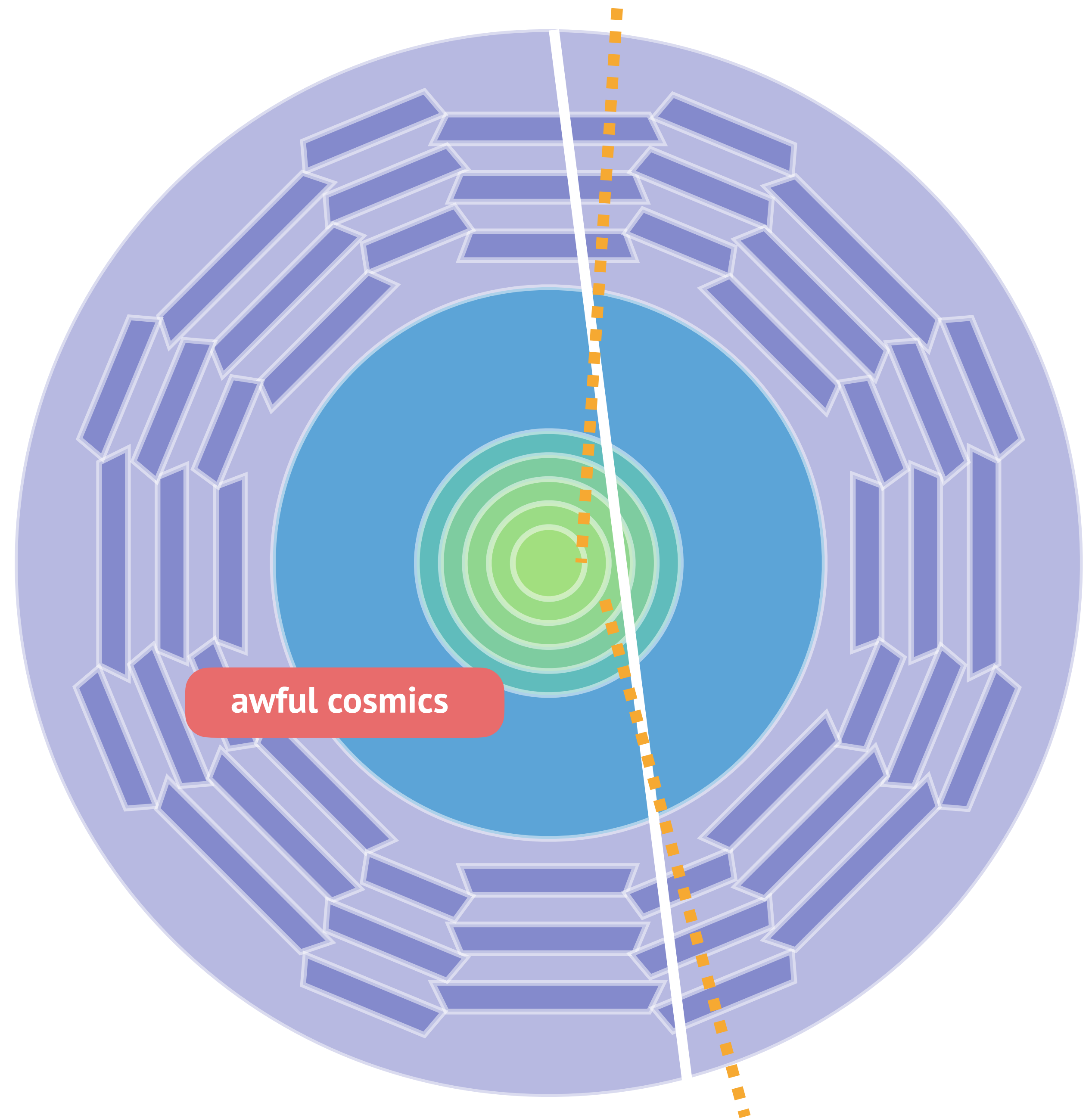
Dominant background for  $\mu\mu$   
~~cosmics~~  
~~fake muons~~  
~~rare standard model processes~~  
material interactions



hasn't been done for a  
reason: **it's hard**

When you have a search with tiny  
backgrounds, it can be more **complicated**  
than you expect

Dominant background for  $\mu\mu$   
**cosmics**  
~~fake muons~~  
~~rare standard model processes~~  
~~material interactions~~  
massive reconstruction failures



# estimating backgrounds

This is the kind of logic you go through  
in each channel -  $ee$ ,  $\mu\mu$ ,  $e\mu$

Figure out what backgrounds really contribute

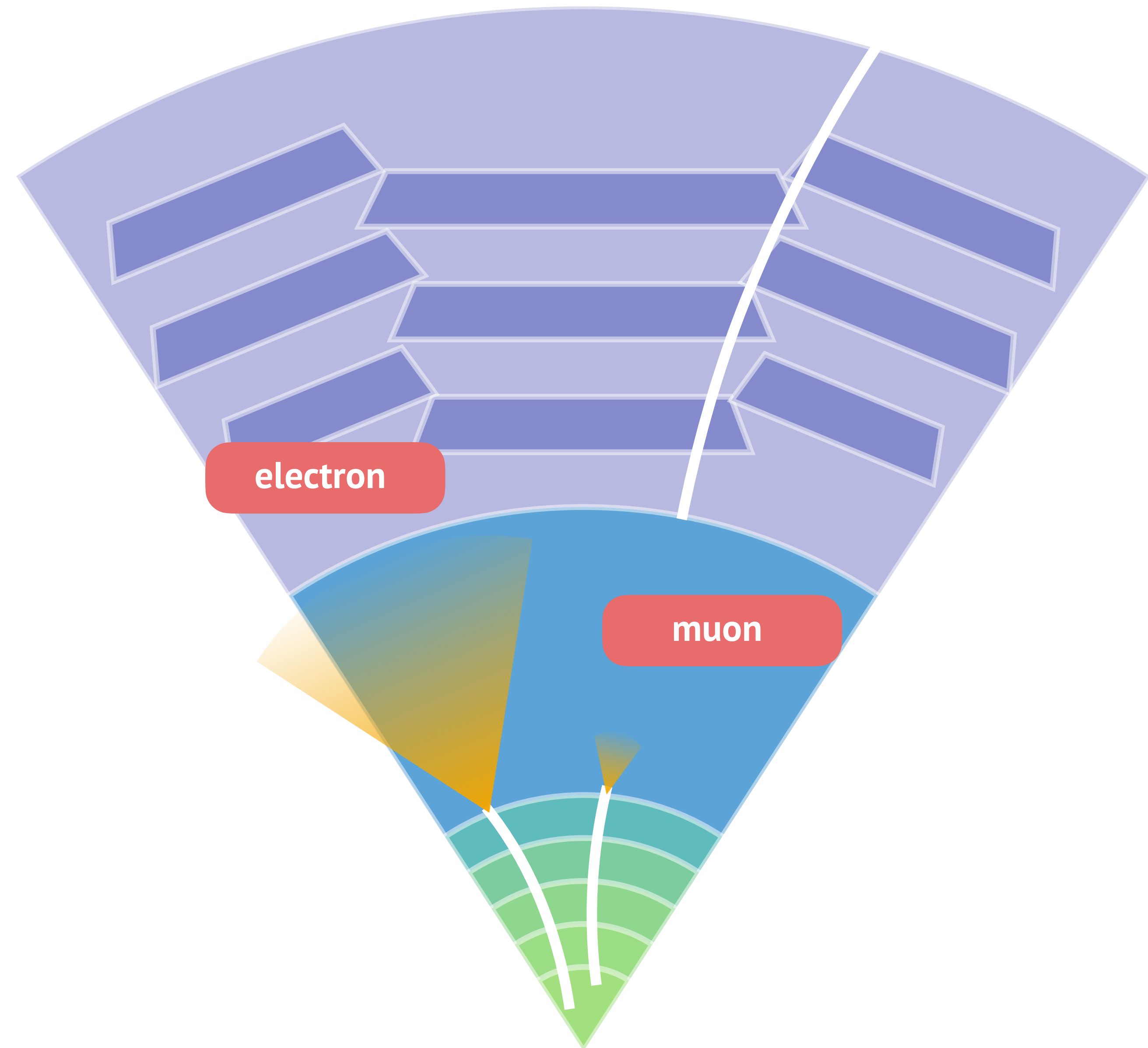
$\mu\mu$ : mismeasured **cosmic** muons

$ee$ : photons associated with **fake** tracks

$e\mu$ : **fake** leptons and **heavy flavor**

Figure out how to estimate these backgrounds

**Can't use simulation!**





# estimating backgrounds

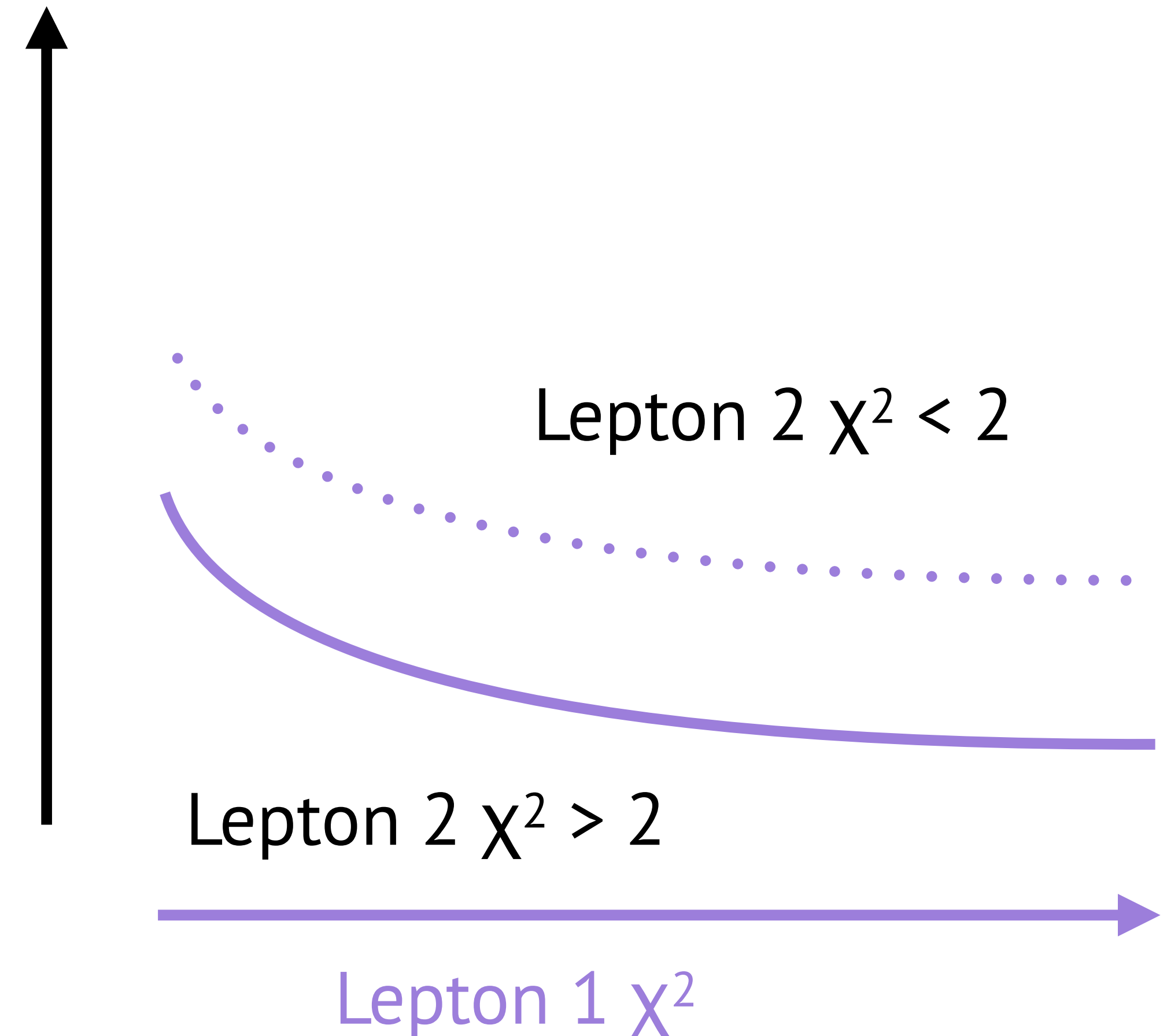
Data-driven approaches designed to model each background:

## fakes and heavy flavor:

take advantage of the **uncorrelated** nature of the two leptons in the event

cut on the **quality** of one – it doesn't affect the **shape** of the other distribution

both pass =  $\frac{(1^{\text{st}} \text{ fails}, 2^{\text{nd}} \text{ passes}) * (2^{\text{nd}} \text{ fails}, 1^{\text{st}} \text{ passes})}{\text{both fail}}$



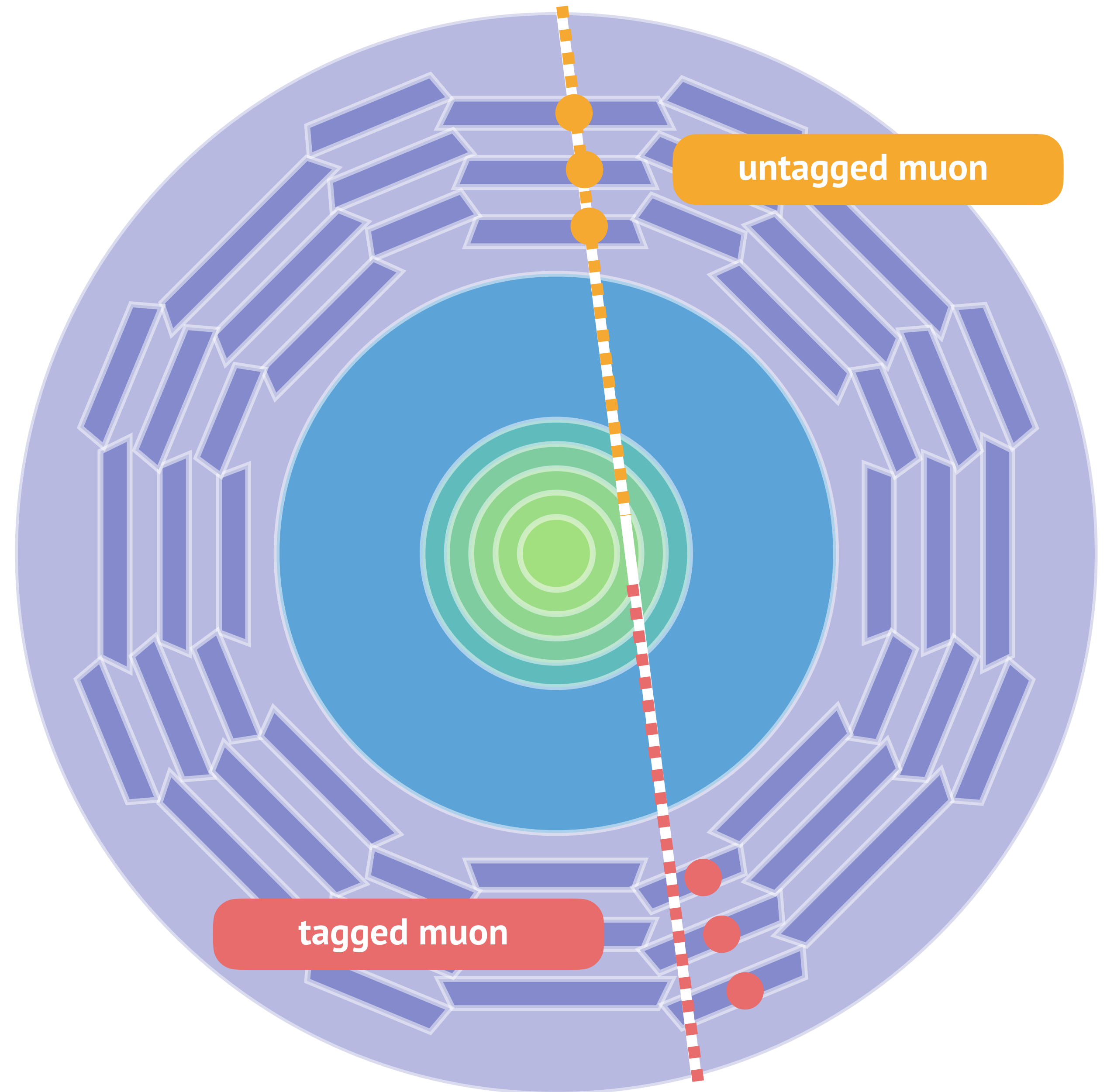
## estimating backgrounds

Cosmic estimate follows a similar strategy, but is less intuitive

Reliable ID tracks, but timing means often **missing information in the MS**

Combined muon has correct measurement, but can't find MS activity to tag  
**mismeasured** muon is **tagged**  
well measured muon isn't

**Quality of muon is independent of probability to be tagged** (given the same muon quality on the opposite side)



# estimating backgrounds

**Measure ratio** of “good quality” to “poor quality” muons in events with one cosmic tag

**Apply** that ratio to events with an untagged poor quality muon to estimate the background

cosmic-tagged

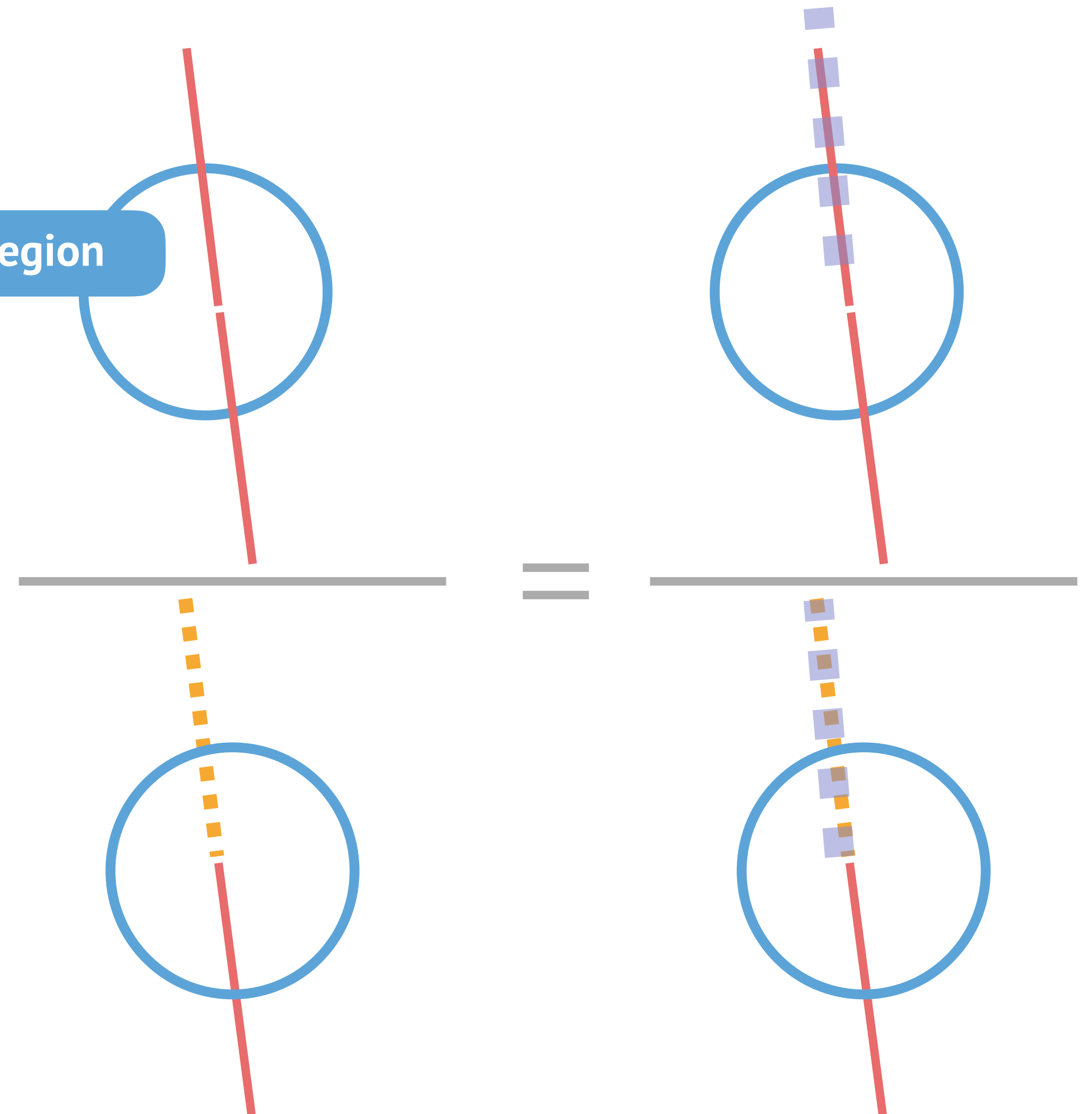
good quality muon

(used in signal region)

poor quality muon

(missing measurements)

signal region





## The last steps...

Able to reduce our backgrounds to **fractions of a event**, estimate all backgrounds  
And finally, can look in our **signal regions**

Region	SR- $ee$	SR- $\mu\mu$	SR- $e\mu$
Fake + Heavy-Flavor	$0.46 \pm 0.10$	–	$0.007^{+0.019}_{-0.007}$
Cosmics	–	$0.11^{+0.20}_{-0.11}$	–
Expected Background	$0.46 \pm 0.10$	$0.11^{+0.20}_{-0.11}$	$0.007^{+0.019}_{-0.007}$

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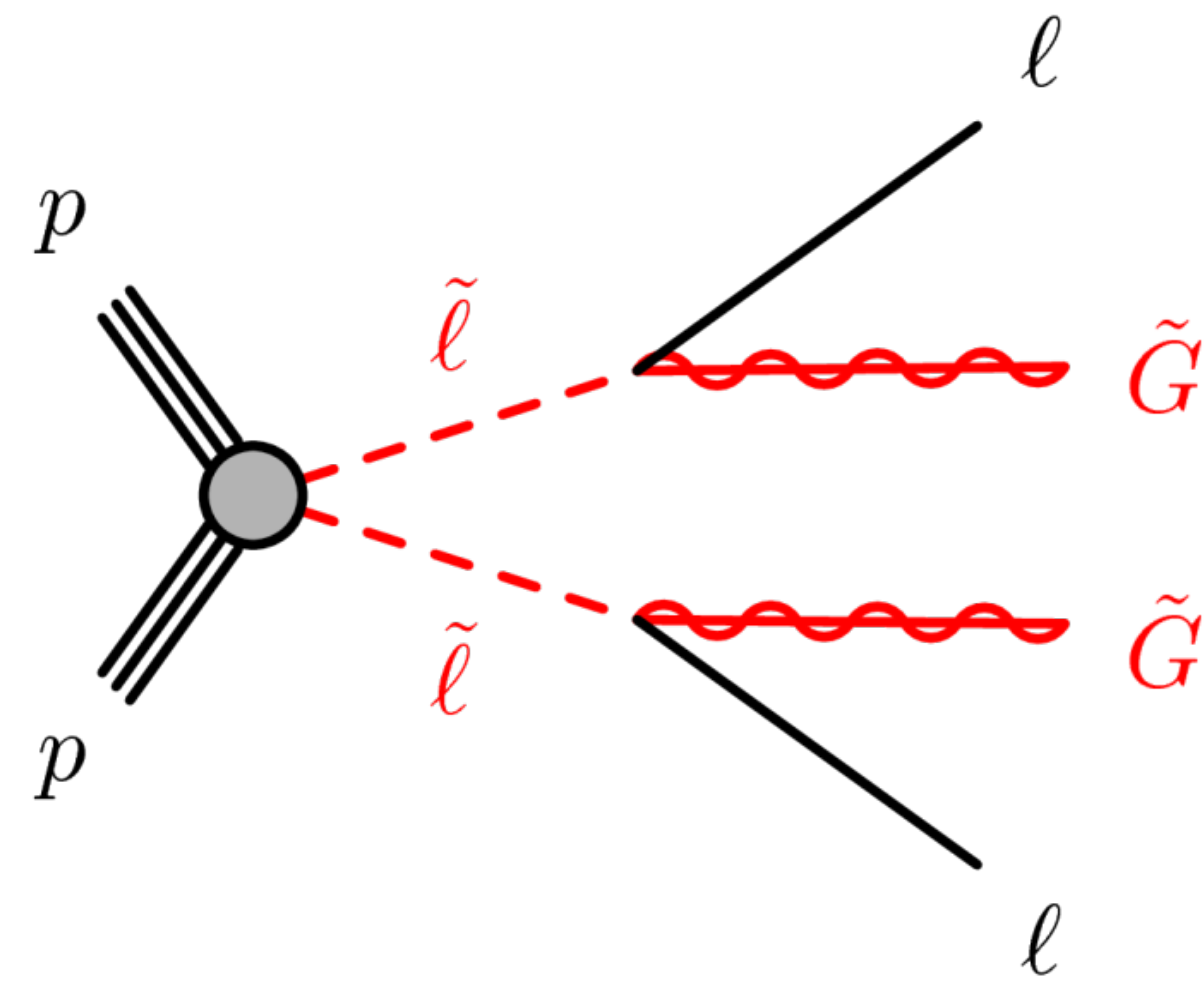
Region	SR- $ee$	SR- $\mu\mu$	SR- $e\mu$
Fake + Heavy-Flavor	$0.46 \pm 0.10$	–	$0.007^{+0.019}_{-0.007}$
Cosmics	–	$0.11^{+0.20}_{-0.11}$	–
Expected Background	$0.46 \pm 0.10$	$0.11^{+0.20}_{-0.11}$	$0.007^{+0.019}_{-0.007}$
Observed events	0	0	0

And see... **nothing!**

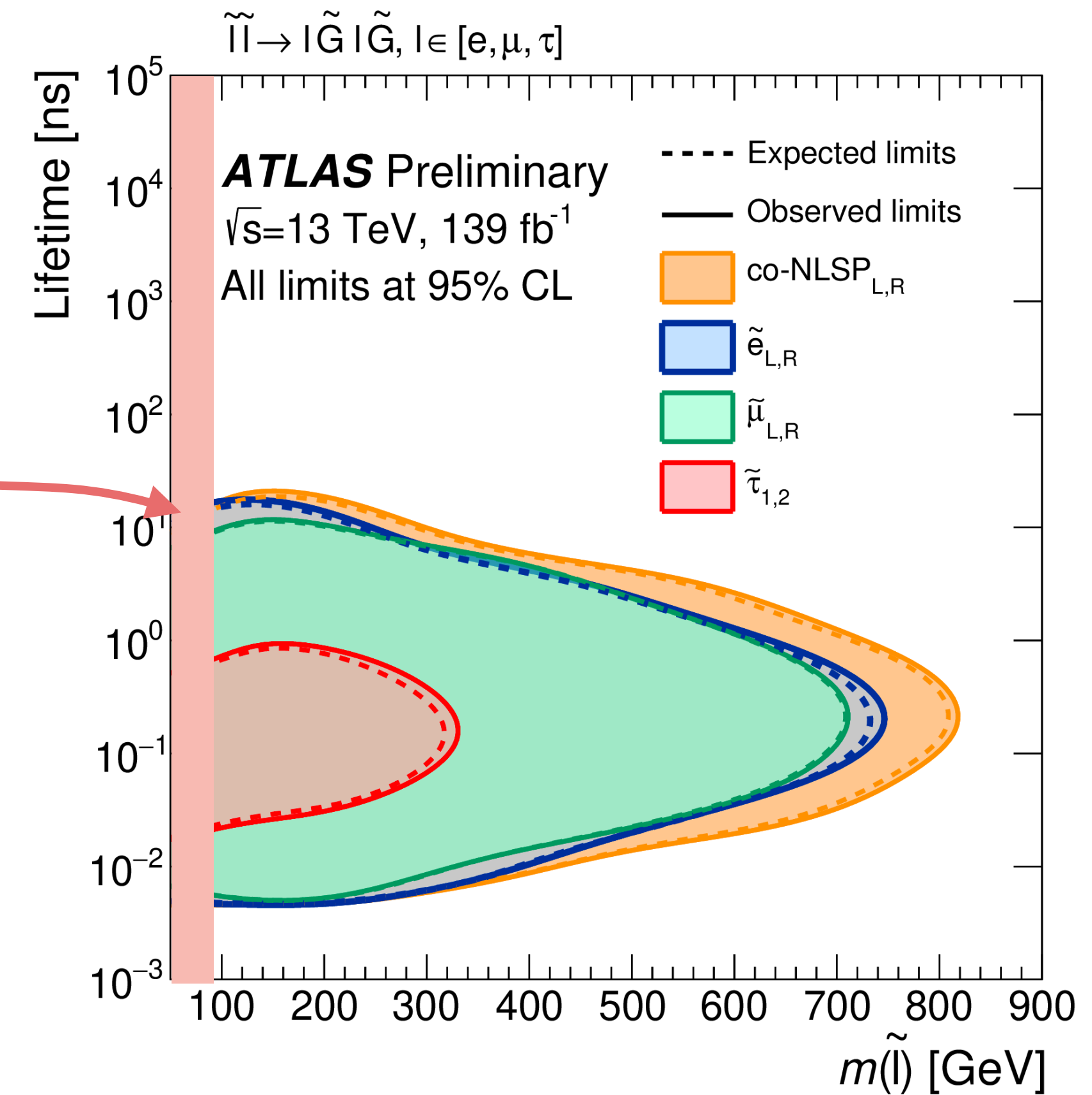
## The last steps...

Able to reduce our backgrounds to **fractions of a event**, estimate all backgrounds

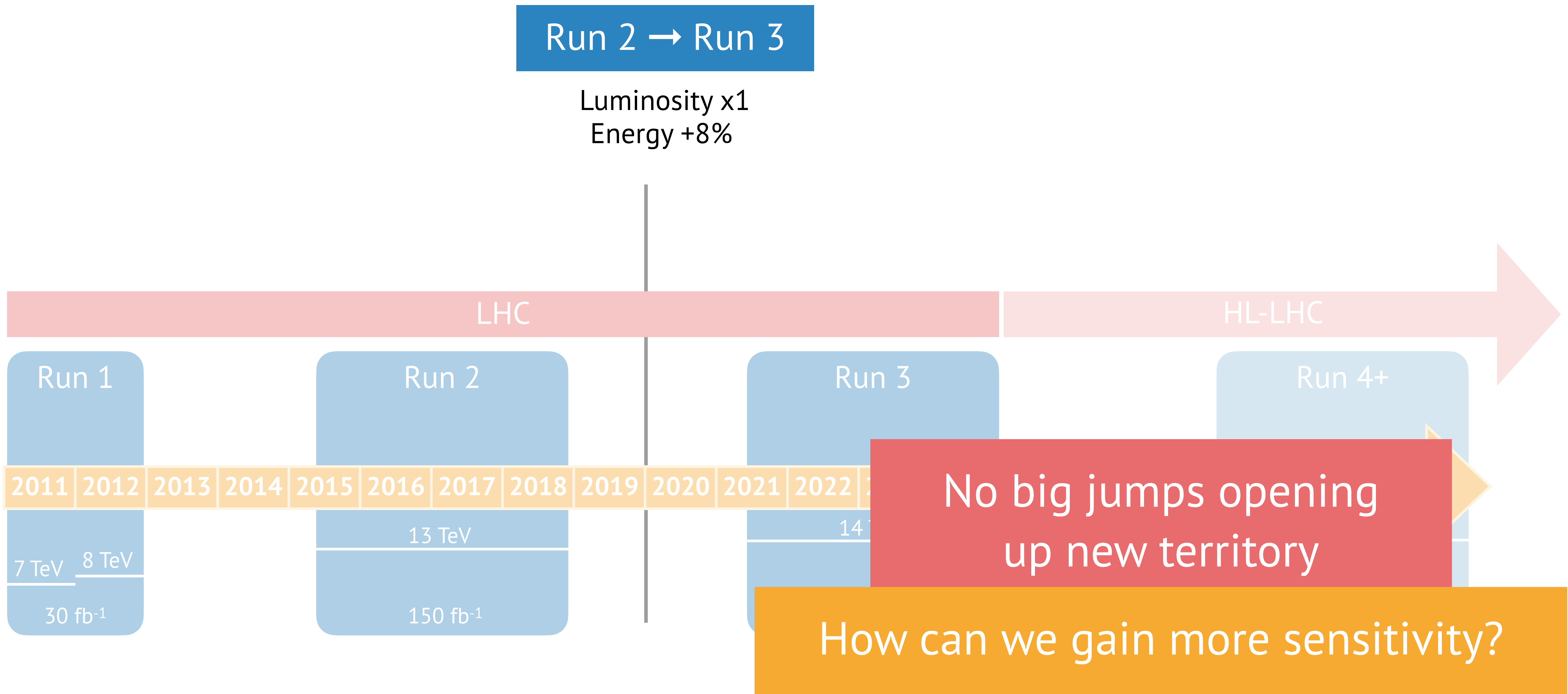
And finally, can look in our **signal regions**



previous limit







To improve long-lived searches

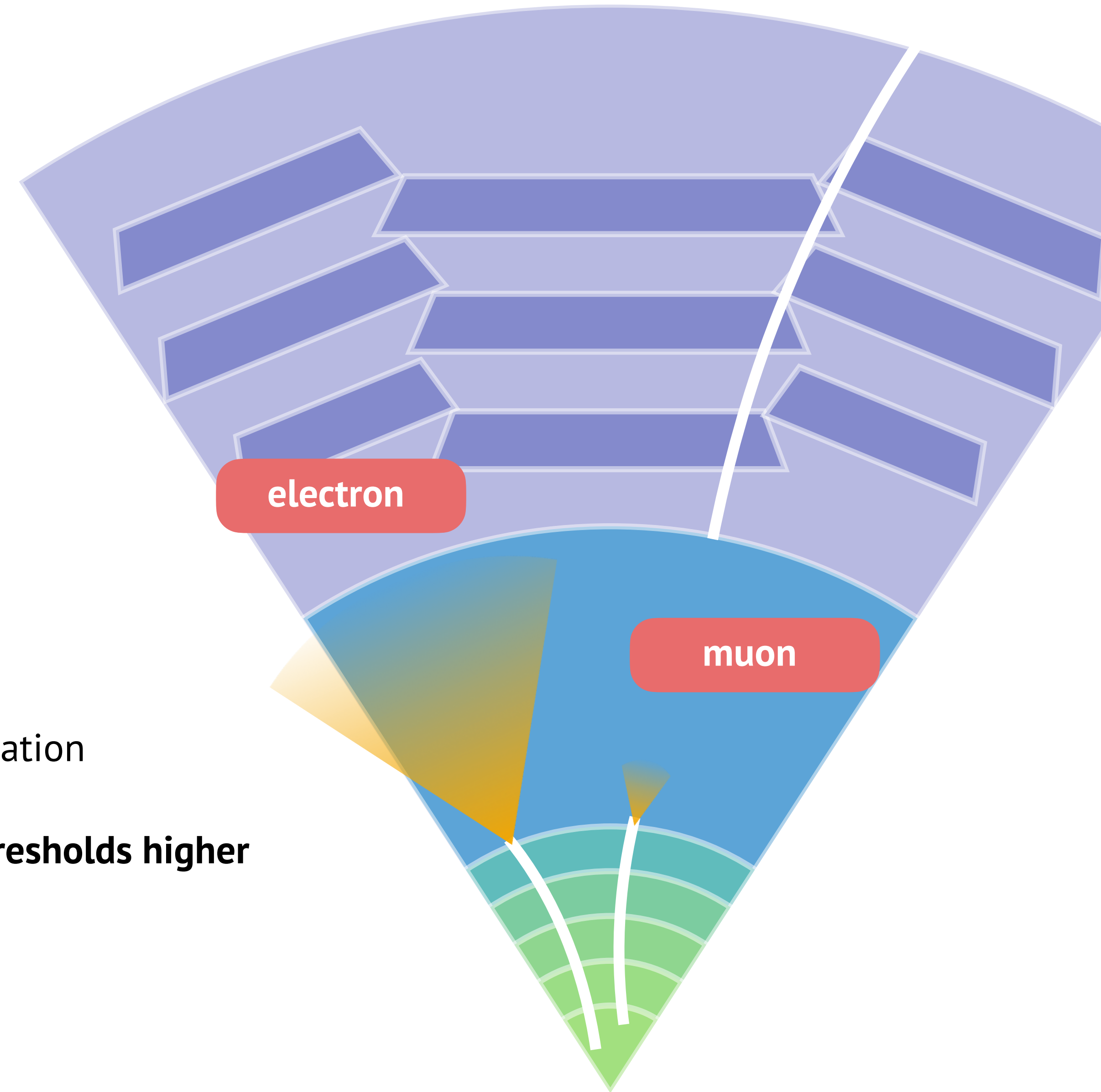
must improve the **Trigger**

For our displaced lepton search...

For some parameter space that we haven't yet excluded trigger acceptance is **as low as 3%**!

When you can't use tracks, throw away **half of your lepton** information

Can't reject backgrounds as well → have to make your **energy thresholds higher**



To improve long-lived searches

must improve the **Trigger**

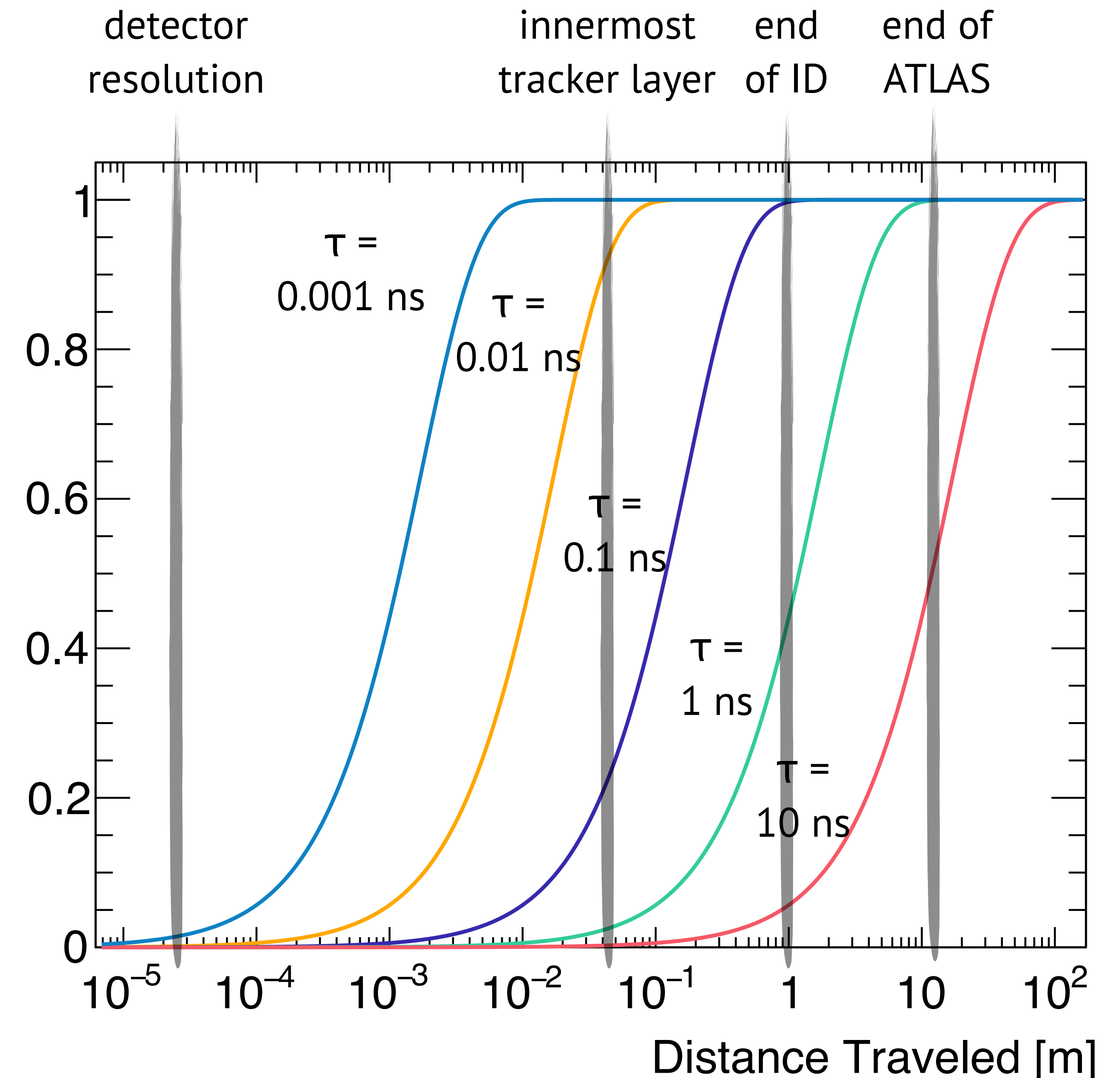
200 GeV particle  
energy: 500 GeV

what can we do for  
Run 3...

No big changes to the detector  
→ can only make **software** changes

Working on developing smart strategies to  
**quickly** produce approximations of these  
long-lived particle track signatures  
what **rough information** can we add  
that makes a big difference?

Probability Particle Has Decayed





To improve long-lived searches

200 GeV particle  
energy: 500 GeV

must improve the **Trigger**

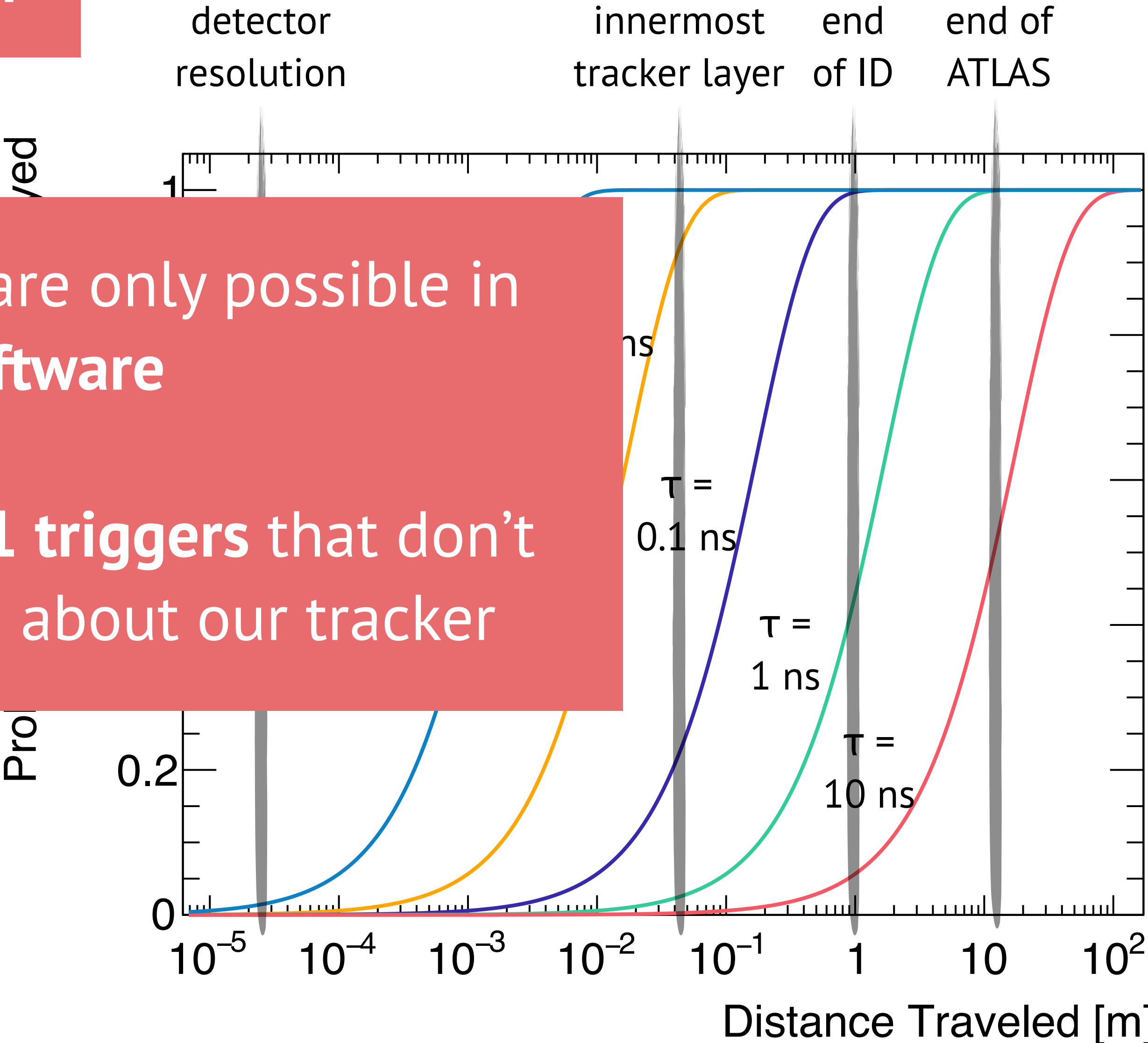
what can we do  
Run 3...

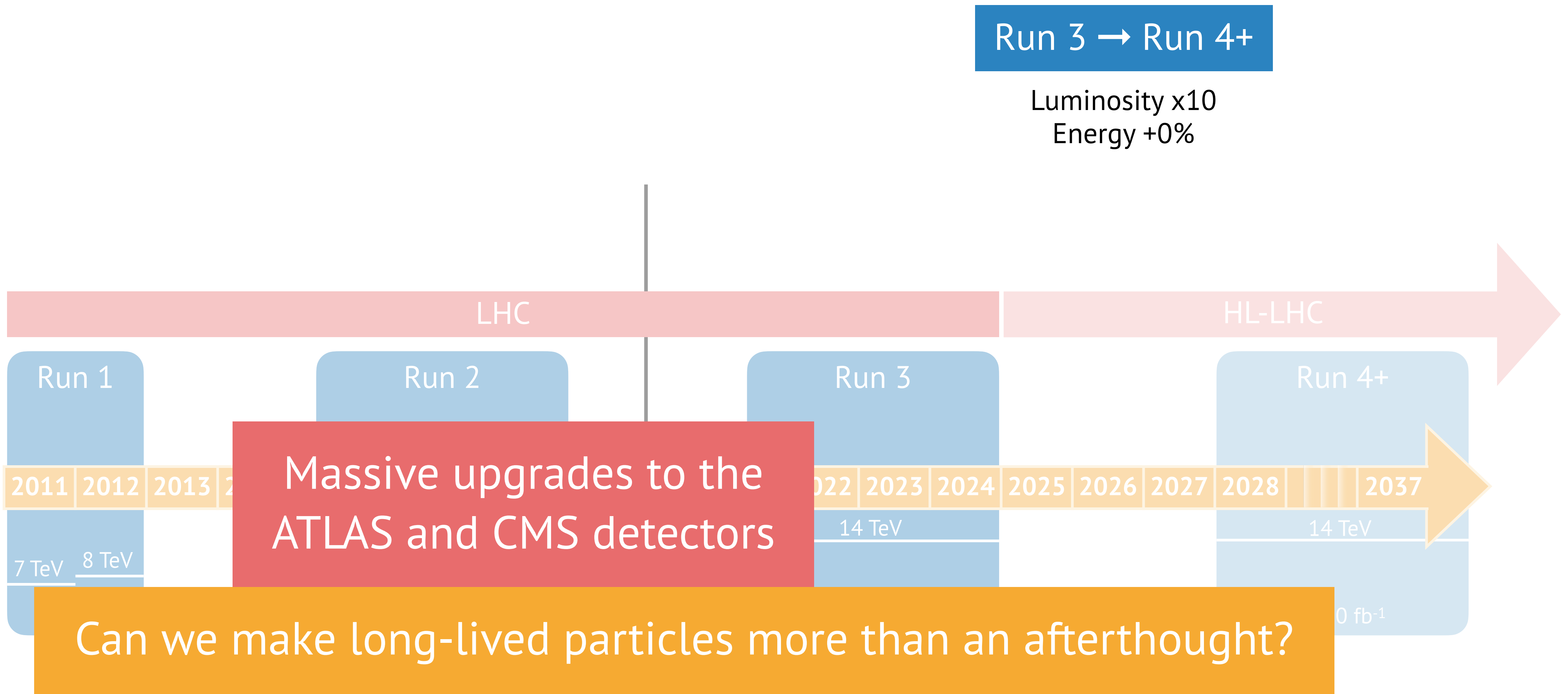
No big changes to the detector  
→ can only make **software**

Working on developing smarter  
**quickly** produce approximate

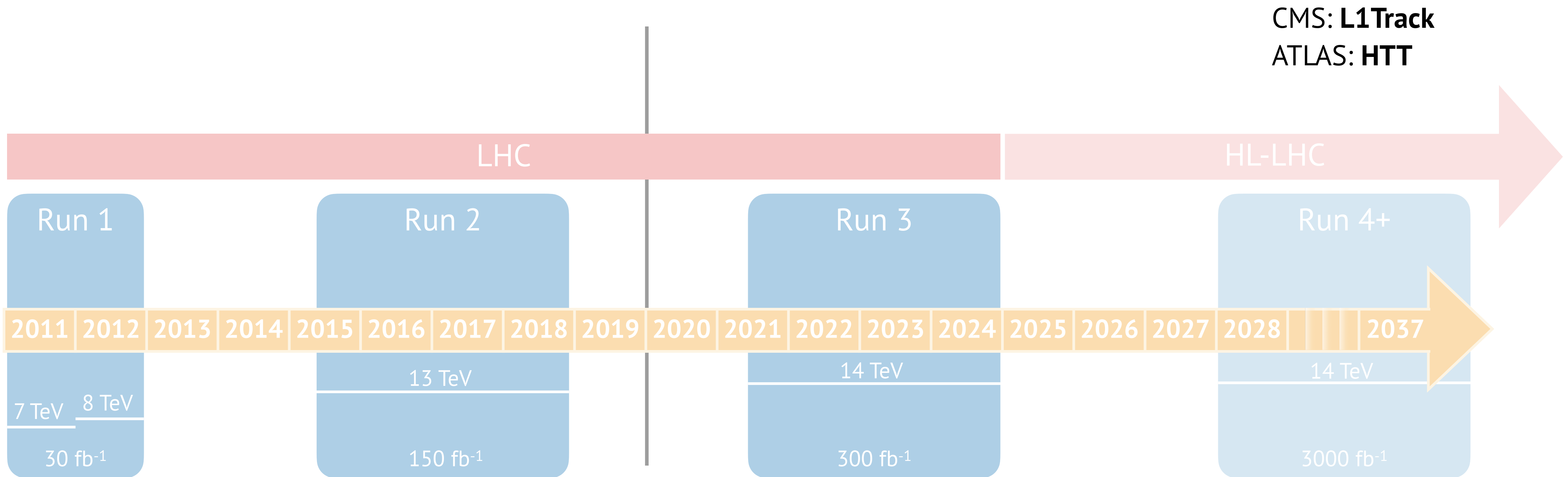
long-lived particle track signatures  
what **rough information** can we add  
that makes a big difference?

these changes are only possible in  
**software**  
still need **level 1 triggers** that don't  
know anything about our tracker



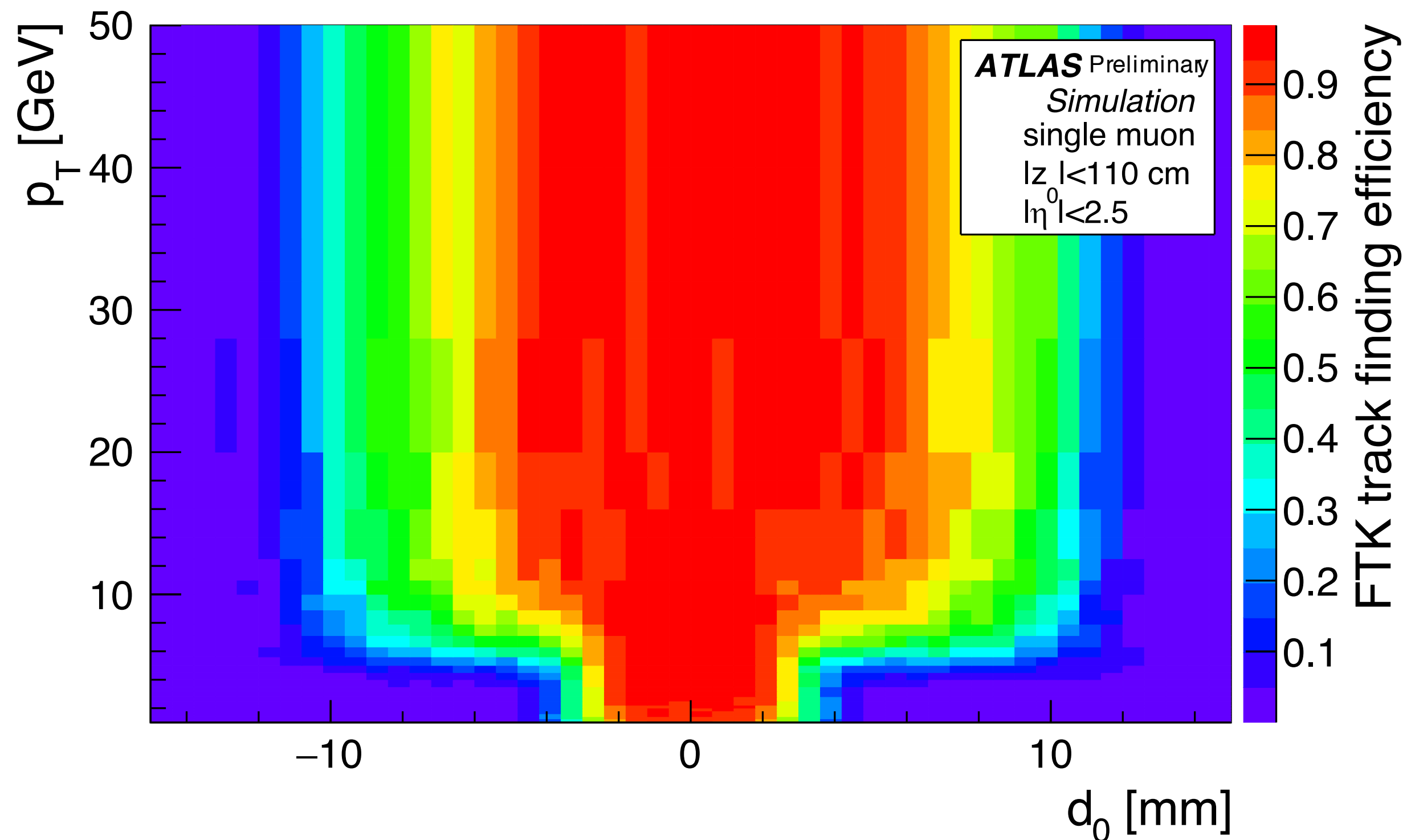


Both ATLAS and CMS are working  
on a hardware-based tracker





## Example of tracking efficiency for a hardware tracker



And on ATLAS, I worked on  
demonstrating that these trackers  
could help us find long-lived particles

For some signatures, this hardware tracker configuration  
resulted in **up to 7x trigger acceptance**

But it's not just trackers...

the tricky thing about long-lived particles  
is how diverse their signatures are

## Tracker

Displaced tracking  
Anomalous charge deposits  
Trackless signatures

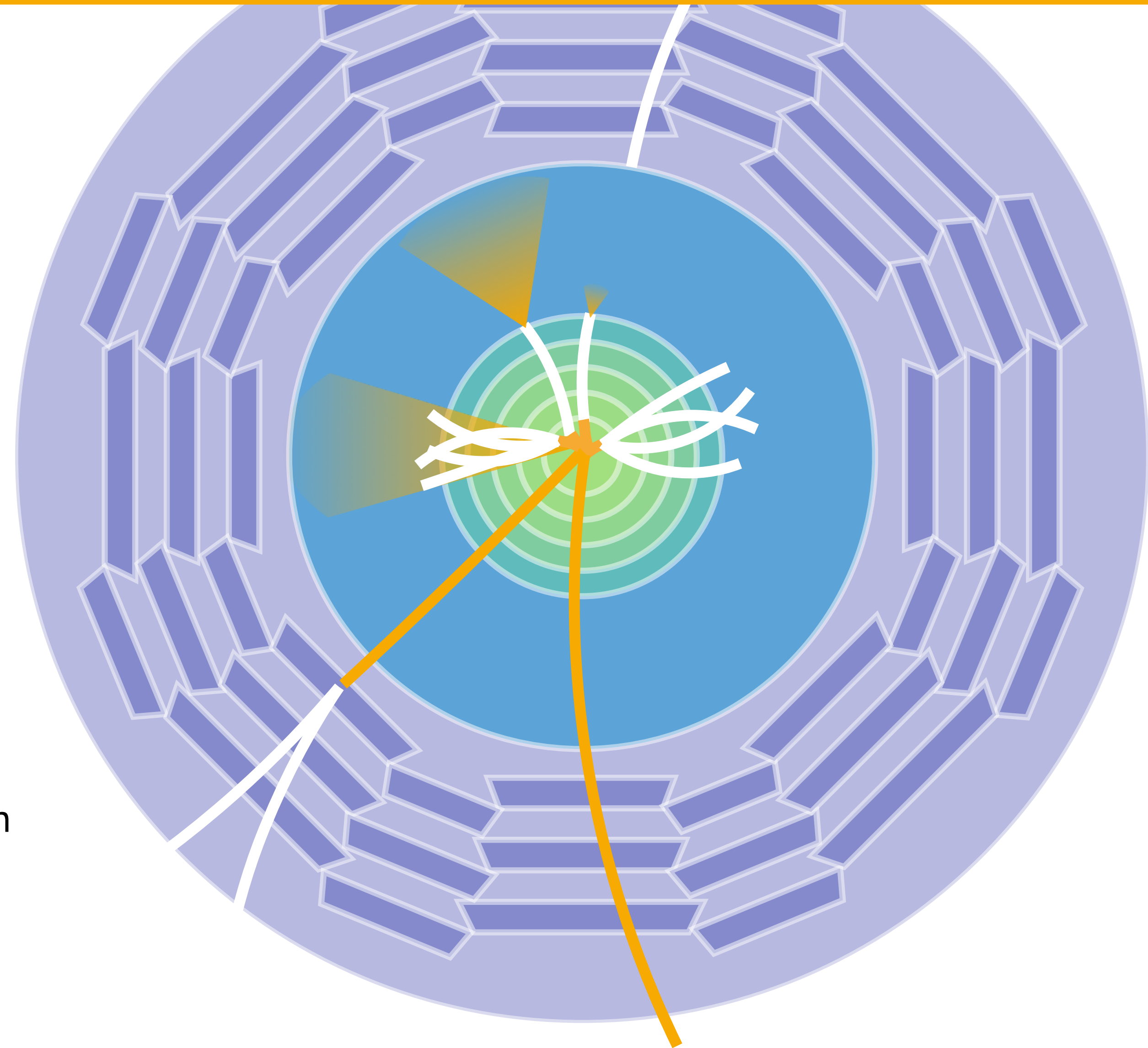
## Muon System

Displaced tracking  
Timing information to find  
slow moving heavy particles

## Calorimeter

Timing information to find  
slow particles, to allow us to  
trace particle trajectories in  
4D (including time!) and reduce  
backgrounds

Fine-grained pointing information  
to look for displaced decays



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Displaced tracking  
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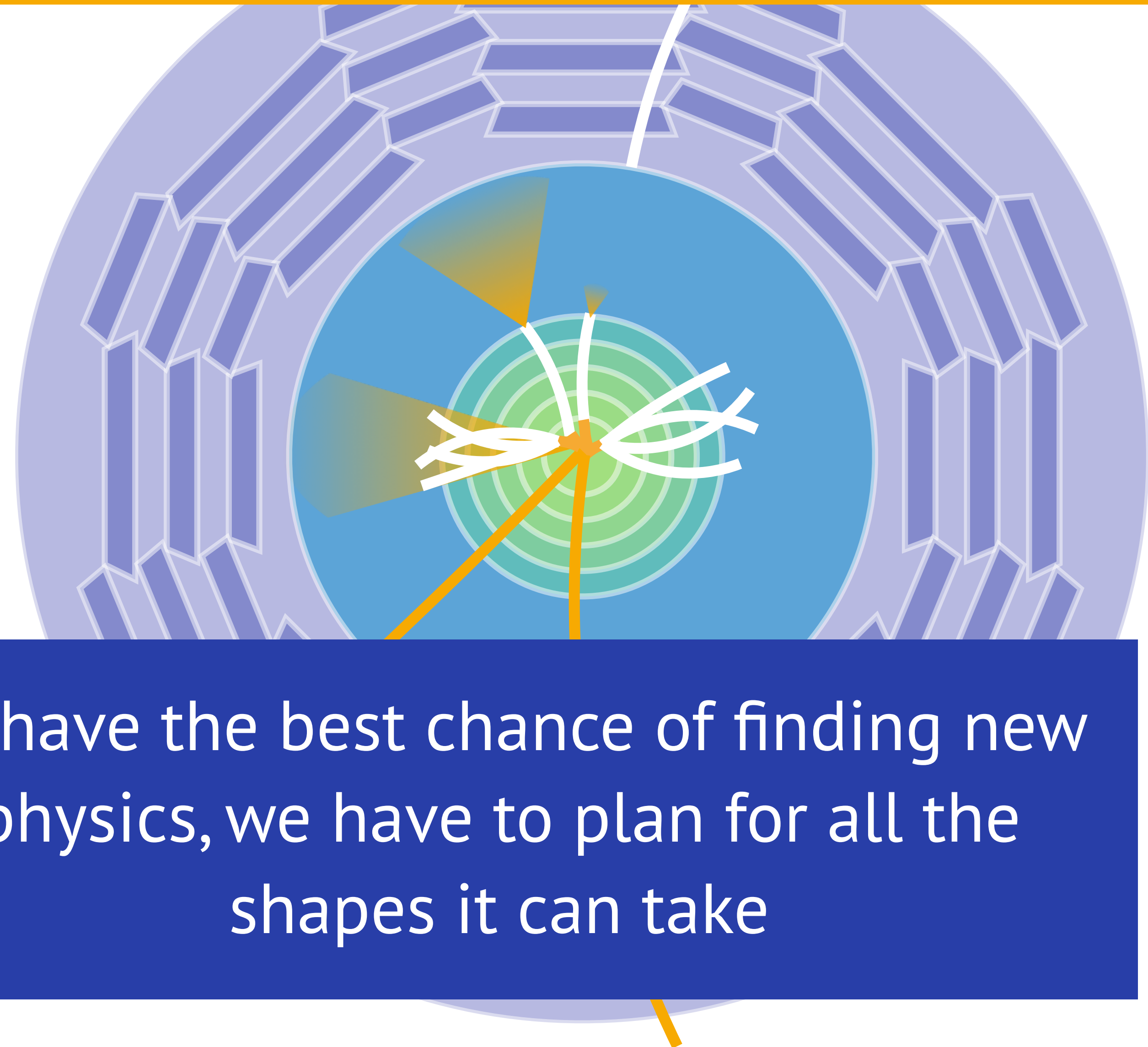
Displaced tracking  
Timing information to find  
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## Calorimeter

Timing information to find  
slow particles, to allow us to  
trace particle trajectories in  
4D (including time!) and  
backgrounds

Fine-grained pointing info  
to look for displaced decays

To have the best chance of finding new  
physics, we have to plan for all the  
shapes it can take





The background of the slide is a complex visualization of a particle detector event. It features a dense network of yellow and orange lines radiating from a central point, representing particle tracks. Interspersed among these tracks are numerous small white and red dots, likely representing individual particles or interaction points. The overall pattern is circular and symmetrical, resembling a starburst or a complex web. The text is overlaid on this background in white and orange boxes.

In conclusion...

As we move forward with LHC physics, **rare and hard to reconstruct** processes are the most interesting place to look

**Long-lived particles** are often both of these things

Still lots of room to find well-motivated new physics, if we just **expand our imagination** (and algorithms) beyond the LHC beamspot

If we act now, can open up **new sensitivity** in the next generation of detectors!





THANK YOU!